



# Nucleon decay searches in JUNO

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On behalf of JUNO Collaboration

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CONFERENCE ON SCIENCE  
AT THE

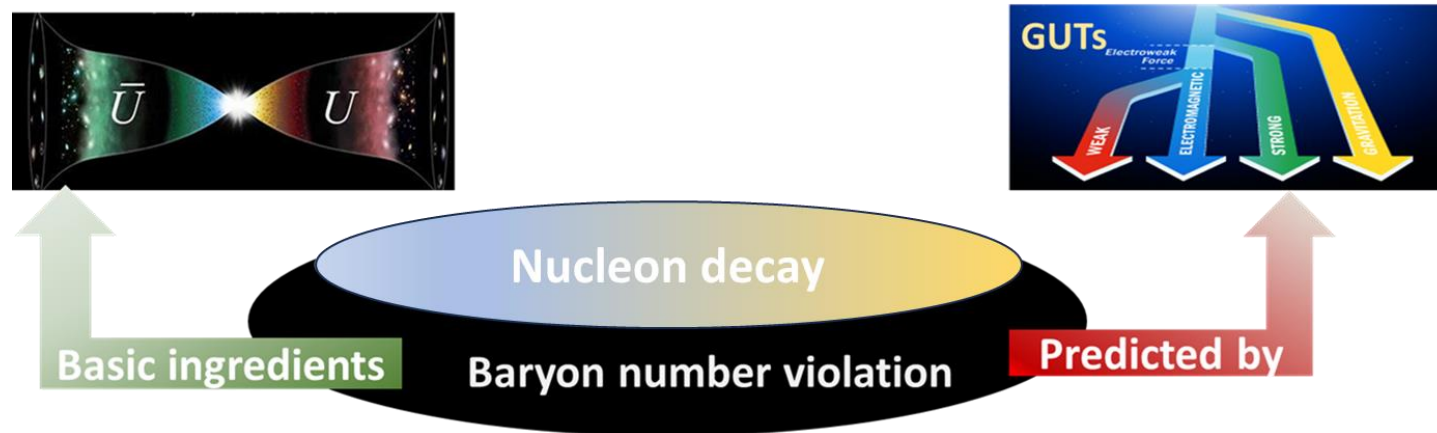
**SANFORD UNDERGROUND  
RESEARCH FACILITY**

South Dakota Mines campus, May 14-16, 2024

- Introduction
- JUNO experiment
- Nucleon decay
  - $p \rightarrow \bar{\nu} + K^+$
  - Invisible neutron decay
- Summary

# Introduction

- GUTs unify the strong, weak and electromagnetic interactions
  - Predict the instability of nucleon
    - Search the nucleon decay to test the GUTs
- Matter-antimatter is asymmetric in universe
  - Sakharov conditions(to explain the asymmetry)
    - **Baryon number violation  $\Delta B \neq 0$**
    - C and CP violation
    - departure from thermal equilibrium

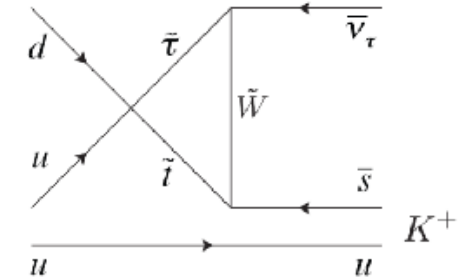
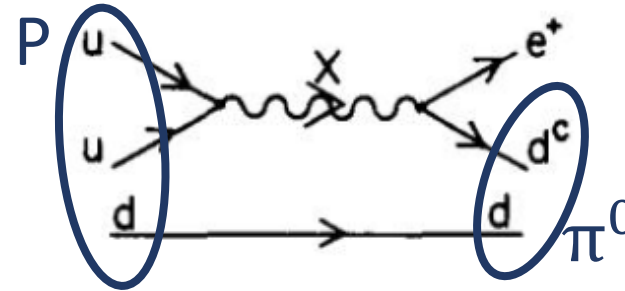


## ➤ GUTs predict instability of nucleon

- Protons can decay into lighter subatomic particles(hypothesis)

### • Examples

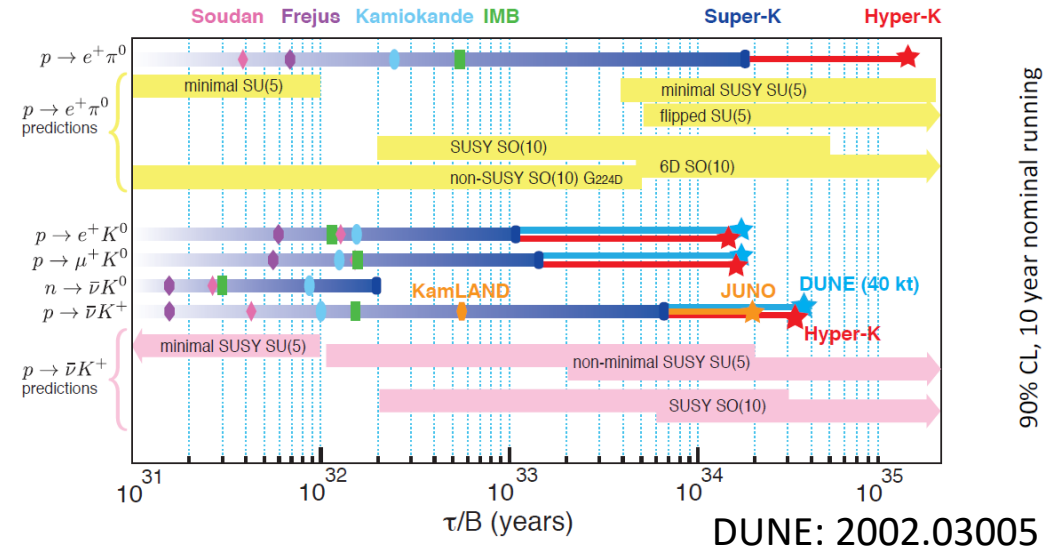
- $p \rightarrow e^+ + \pi^0$  (Non-SUSY GUTs)
- $p \rightarrow \bar{\nu} + K^+$  (SUSY GUTs)



- Invisible neutron decay: neutron decay into **undetected** particles (hypothesis)

### • Examples

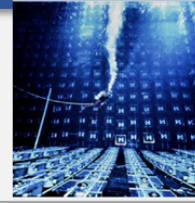
- $n \rightarrow \text{neutrino}, n \rightarrow \text{dark fermions} \dots$
- Phys. Rev. D 67, 075015 (2003).
- Phys. Lett. B 662, 259 (2008).
- Phys. Rev. D 98, 035049 (2018).



DUNE: 2002.03005

## IMB

- 3.3 kton water-Cherenkov detector, 2000 PMTs
- $\tau/B(p \rightarrow e^+ \pi^0) > 5.5 \times 10^{32} \text{ years}$  (1990)
- No proton decay have been found



## KamiokaNDE

- 0.88 kton water-Cherenkov detector, 948 PMTs
- $\tau/B(p \rightarrow e^+ \pi^0) > 2.6 \times 10^{32} \text{ years}$  (1989)
- No proton decay have been found



## Super-Kamiokande (Super-K)

- 22.5 kton water-Cherenkov detector, 11146 PMTs
- $\tau/B(p \rightarrow e^+ \pi^0) > 2.4 \times 10^{34} \text{ years}$  (2020)
- $\tau/B(p \rightarrow \bar{\nu} K^+) > 5.9 \times 10^{33} \text{ years}$  (2014)
- No proton decay have been found

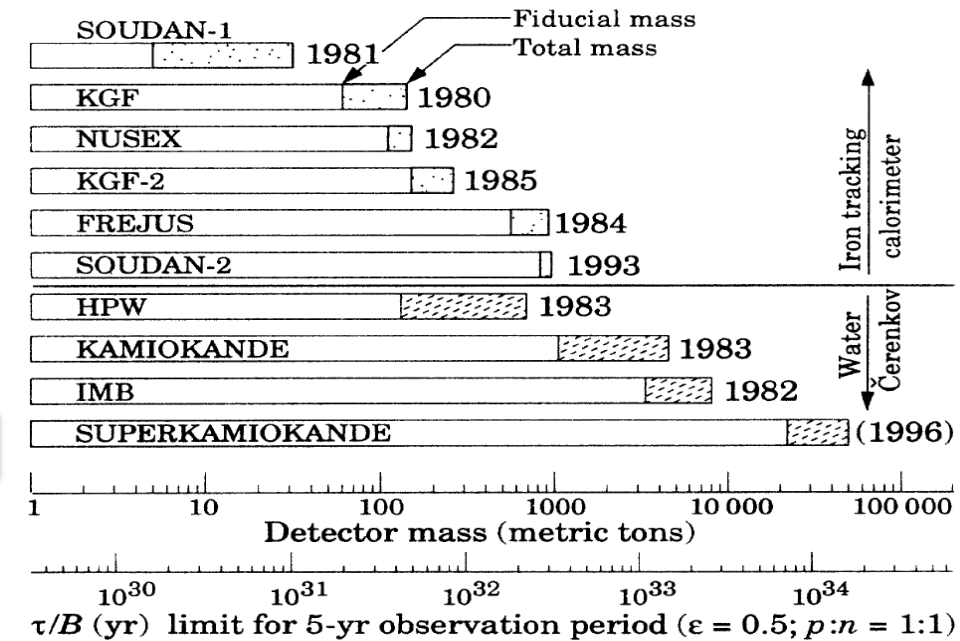


## KamLAND

- 0.9 kton liquid scintillator detector
- $\tau/B(p \rightarrow \bar{\nu} K^+) > 5.4 \times 10^{32} \text{ years}$  (2015)
- No proton decay have been found



## Past searches





## KamLAND

- 0.9 kton liquid scintillator detector
- $\tau/B(n \rightarrow inv) > 5.8 \times 10^{29} \text{ years}$  (2006)
- $\tau/B(nn \rightarrow inv) > 1.4 \times 10^{30} \text{ years}$  (2006)
- No invisible neutron decay have been found

## SNO+

- 0.9 kton liquid scintillator detector
- $\tau/B(n \rightarrow inv) > 9.0 \times 10^{29} \text{ years}$  (2022)
- $\tau/B(nn \rightarrow inv) > 1.5 \times 10^{28} \text{ years}$  (2022)
- No invisible neutron decay have been found



## Hyper-K



## DUNE



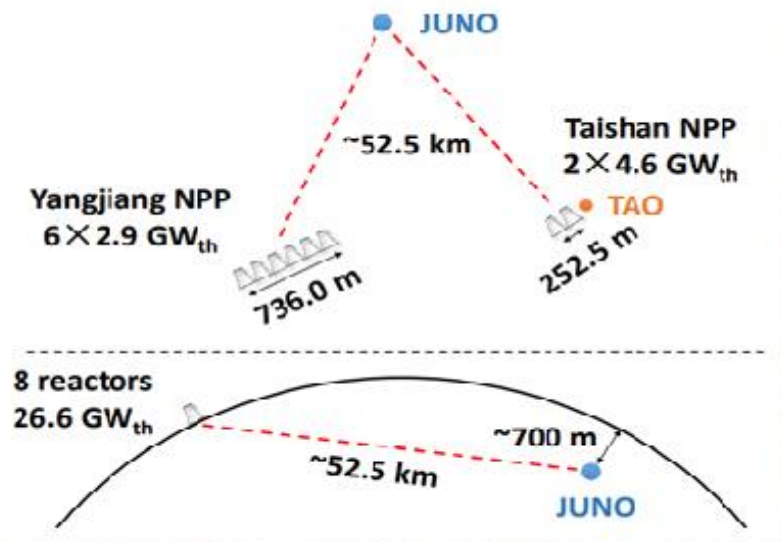
## JUNO



	Hyper-K	DUNE	JUNO
Mass (kton)	258 (186)	4*17 (4*10)	20
Target Nucleus	H2O	Ar40	12% H, 88% C12
Technology	Water Cherenkov	LAr TPC	Liquid Scintillator
Advantages	Large mass and cheap Good particle Identification Good direction resolution	Excellent track reconstruction Excellent particle Identification Good energy resolution	Excellent energy resolution 3% Excellent $E$ threshold 0.7MeV
Shortcomings	Cerenkov threshold	Complex FSI for Ar40	Direction information lost



# JUNO experiment



	JUNO
<b>Photon Statistics</b>	<b>1200 p.e./MeV</b>
<b>PMT coverage</b>	<b>75%</b>
<b>LS transparency</b>	<b>&gt; 20 m</b>
<b>Light yield(anthracene)</b>	<b>45%</b>
<b>Detection Eff.(QE×CE)</b>	<b>30%</b>

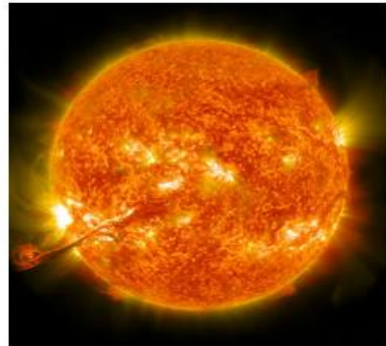
**Civil construction finished in Dec, 2021**



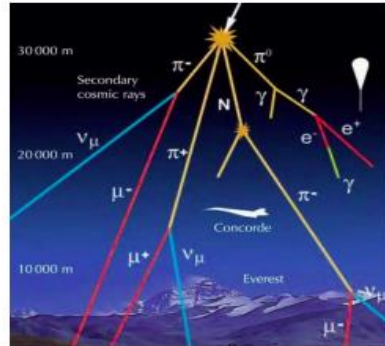
## ➤ Jiangmen Underground Neutrino Observatory (JUNO), a multi-purpose neutrino experiment



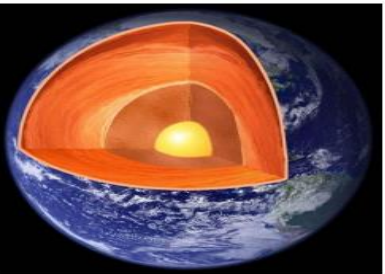
~ 50/day



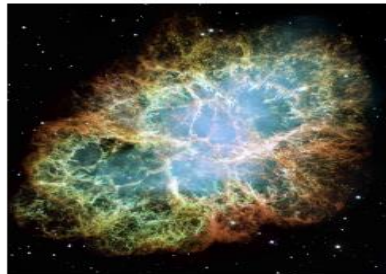
$\mathcal{O}(1000)$ /day



~ 10 - 20/day

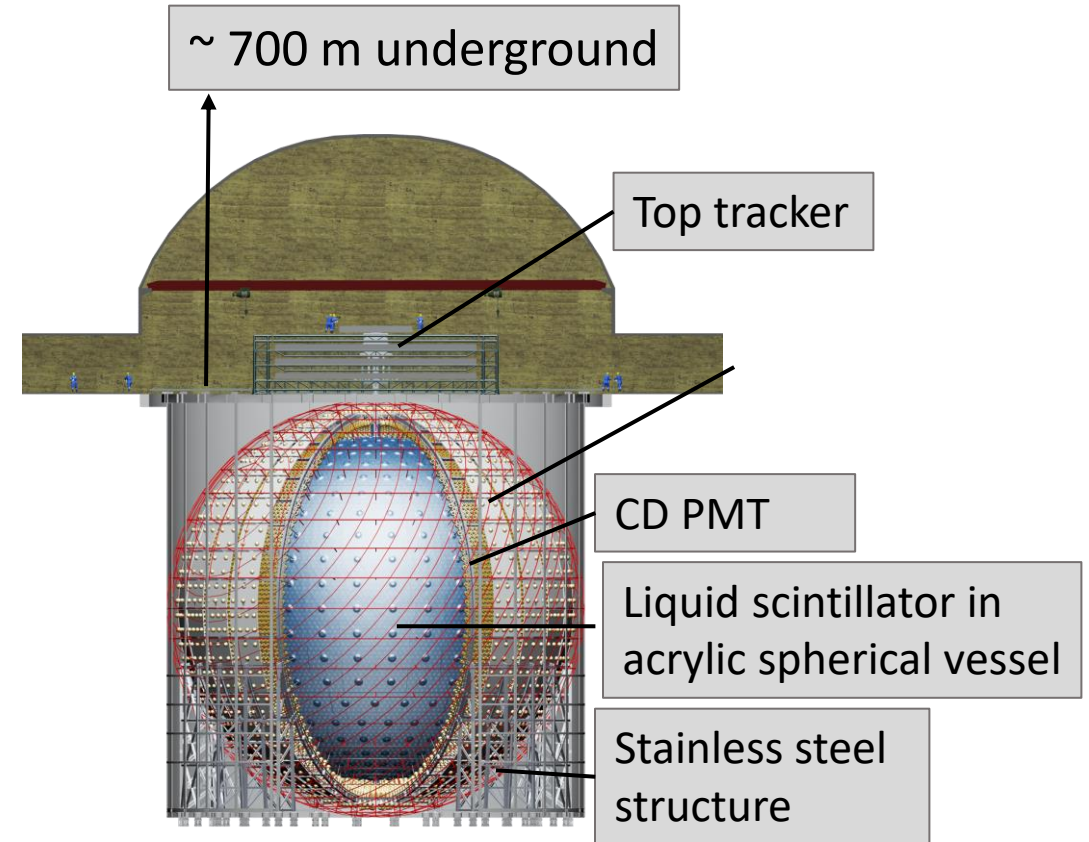


~ 1 - 2/day



CCSN @10kpc :  
 $\mathcal{O}(1000)$ /s  
DSNB: few/year

**New Physics  
(Nucleon decay...)**

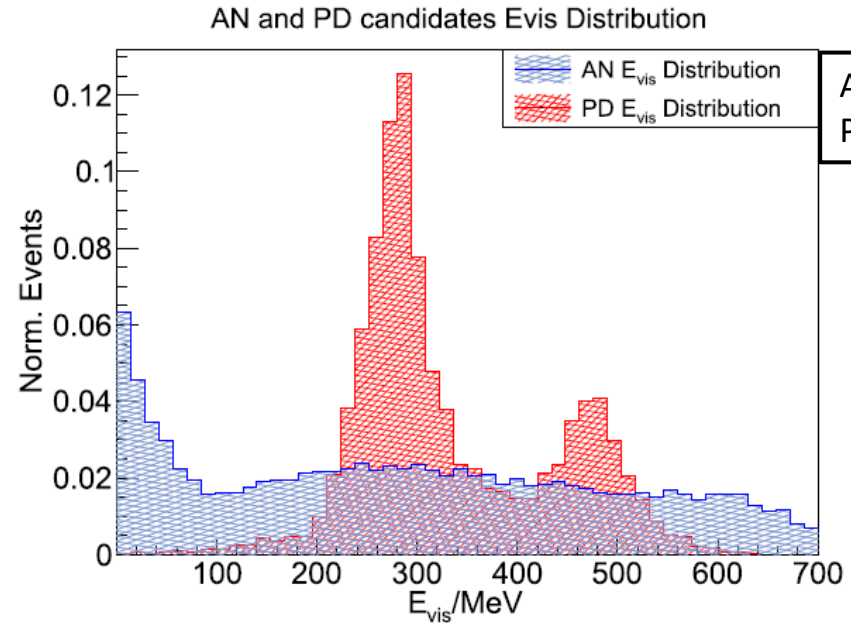
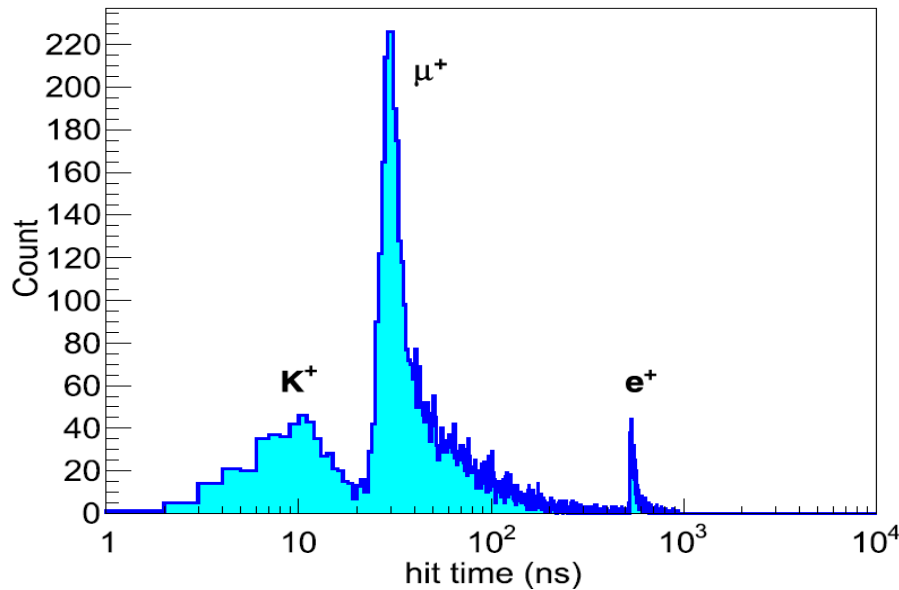
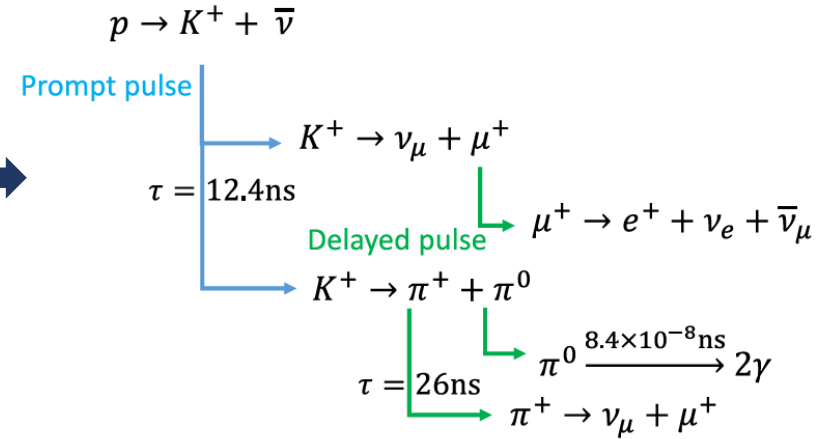


**Nucleon decay:  $p \rightarrow \bar{\nu} + K^+$**

- Form triple coincident signals

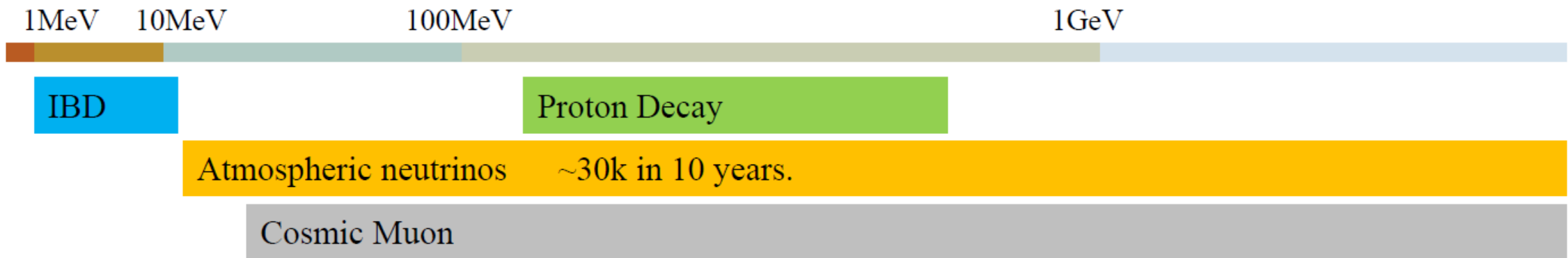
Decay mode	Branching ratio (%)	Kinetic energy sum (MeV)
$K^+ \rightarrow \mu^+ \nu_\mu$	$63.55 \pm 0.11$	152
$K^+ \rightarrow \pi^+ \pi^0$	$20.66 \pm 0.08$	354
$K^+ \rightarrow \pi^+ \pi^+ \pi^-$	$5.59 \pm 0.04$	75
$K^+ \rightarrow \pi^0 e^+ \nu_e$	$5.07 \pm 0.04$	265–493
$K^+ \rightarrow \pi^0 \mu^+ \nu_\mu$	$3.353 \pm 0.034$	200–388
$K^+ \rightarrow \pi^+ \pi^0 \pi^0$	$1.761 \pm 0.022$	354

Two most dominant channels



AN: Atmospheric neutrinos  
PD: Proton decay

➤ Most dominant background: Atmospheric neutrinos



Type	Ratio (%)	Ratio with $E_{vis}$ in [100 MeV, 600 MeV](%)	Interaction	Signal characteristics
NCES	20.2	15.8	$\nu + n \rightarrow \nu + n$ $\nu + p \rightarrow \nu + p$	Single Pulse
CCQE	45.2	64.2	$\bar{\nu}_l + p \rightarrow n + l^+$ $\nu_l + n \rightarrow p + l^-$	Single Pulse
Pion Production	33.5	19.8	$\nu_l + p \rightarrow l^- + p + \pi^+$ $\nu + p \rightarrow \nu + n + \pi^+$	Approximate Single Pulse (Second pulse too low)
Kaon Production	1.1	0.2	$\nu_l + n \rightarrow l^- + \Lambda + K^+$ $\nu_l + p \rightarrow l^- + p + K^+$	Double Pulse

## Low energy background

- Removed by energy cut
  - IBD, solar- $\nu$ , geo- $\nu$ , and low energy atm- $\nu$

## Cosmic Muon

- Removed by muon veto and FV cut



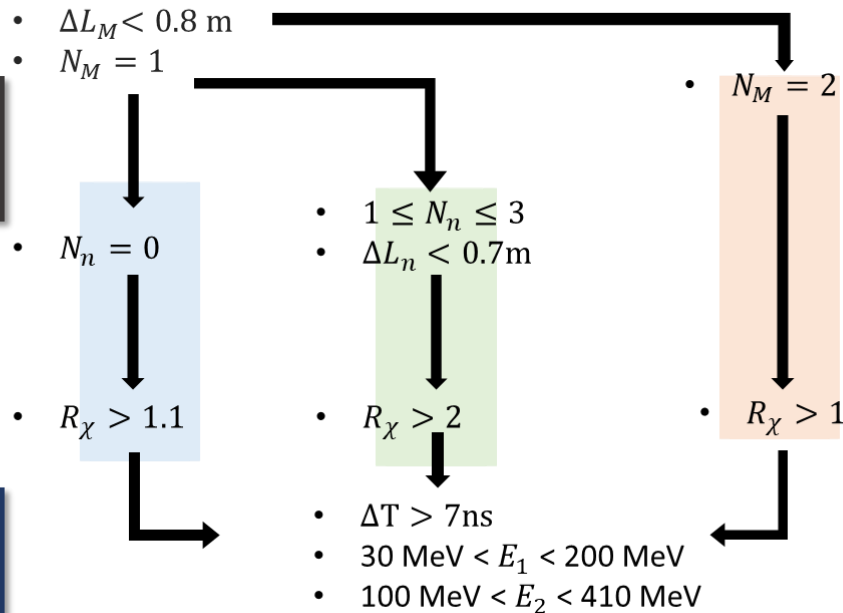
## ➤ Selection flow

### Basic selection

- $200 \text{ MeV} < E_{vis} < 600 \text{ MeV}$
- $R_v < 17.5 \text{ m}$

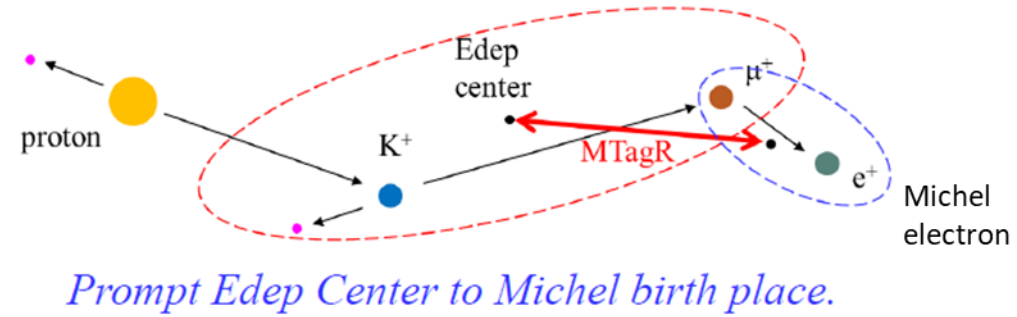


### Delayed signal selection



### Time character selection

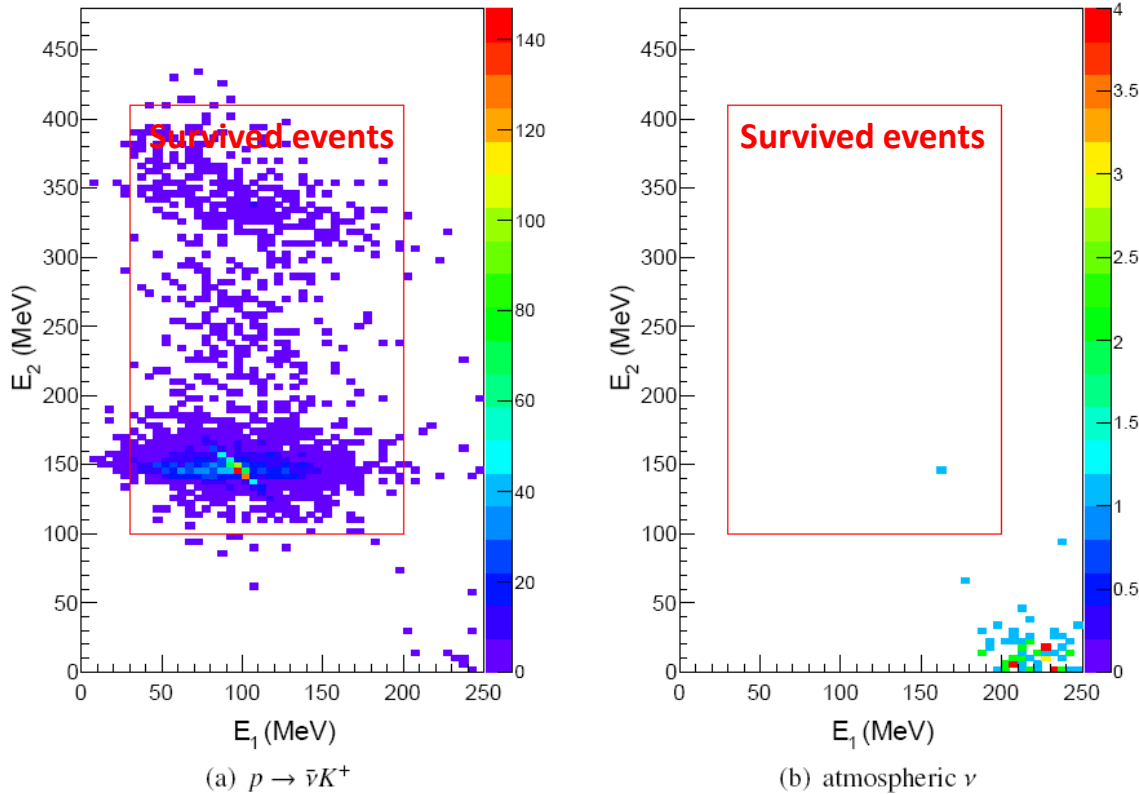
- $\Delta T > 7 \text{ ns}$
- $30 \text{ MeV} < E_1 < 200 \text{ MeV}$
- $100 \text{ MeV} < E_2 < 410 \text{ MeV}$



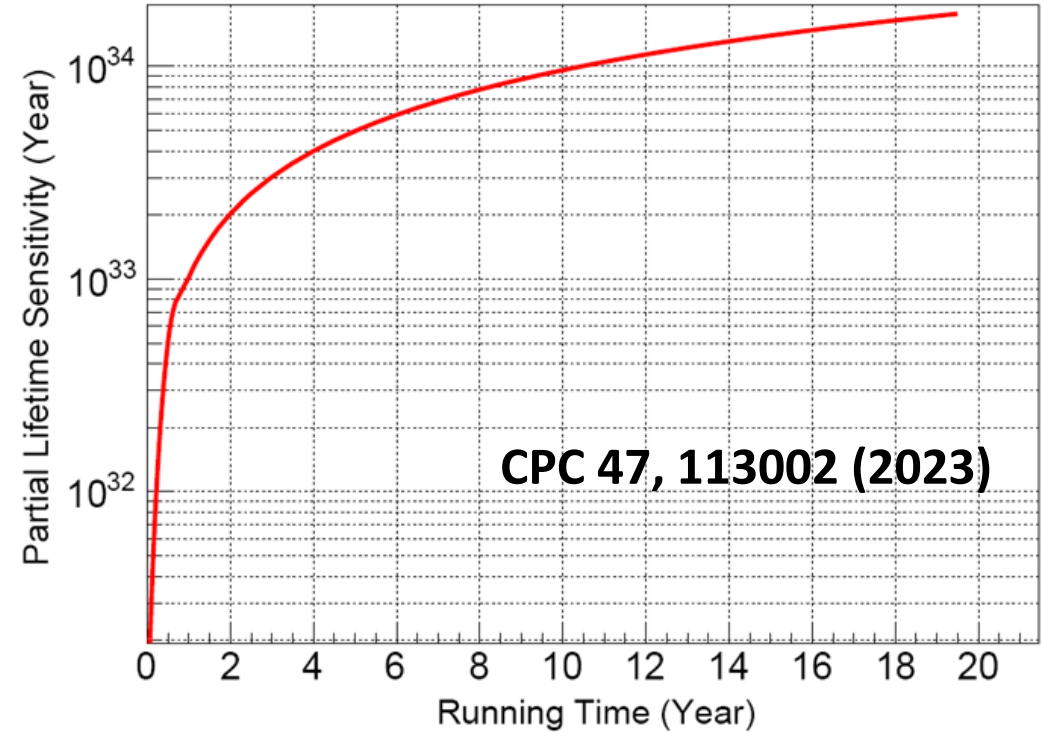
## Selection result

Criteria	Survival rate of $p \rightarrow \bar{\nu}K^+$ (%)			Survival count (fraction) of atmospheric $\nu$		
	Sample 1	Sample 2	Sample 3	Sample 1	Sample 2	Sample 3
basic selection	$E_{vis}$			51299 (32.1%)		
	$R_v$			47849 (29.9%)		
Delayed signal selection	$N_M$		4.4	20739 (13.0%)		1143 (0.7%)
	$\Delta L_M$		4.4	13796 (8.6%)		994 (0.6%)
	$N_n$	48.4	17.9	5403 (3.4%)	6857 (4.3%)	—
	$\Delta L_n$	—	16.6	—	4472 (2.8%)	—
Time character selection	$R_\chi$	45.9	9.0	3.8	4326 (2.7%)	581 (0.4%)
	$\Delta T$	28.3	7.7	2.4	121 (0.07%)	18 (0.01%)
	$E_1, E_2$	27.4	7.3	2.2	1 (0.0006%)	0
Total	36.9			1		

- $R_v$ : Fiducial volume
- $N_M$ : tagged Michel electron number
- $\Delta L_M$ : correlated distance to Michel electron
- $N_n$ : tagged neutron number
- $\Delta L_n$ : correlated distance to neutron
- $R_\chi$ :  $\chi^2$  ratio



**Background: 0.2/10 yrs**  
**Signal efficiency : 36.9%**



**$\tau/B(p \rightarrow \bar{\nu} K^+) > 9.6 \times 10^{33}$  yrs/10 yrs**  
**Best limit:  $5.9 \times 10^{33}$  yrs from Super-K**

# **Nucleon decay: invisible neutron decay**

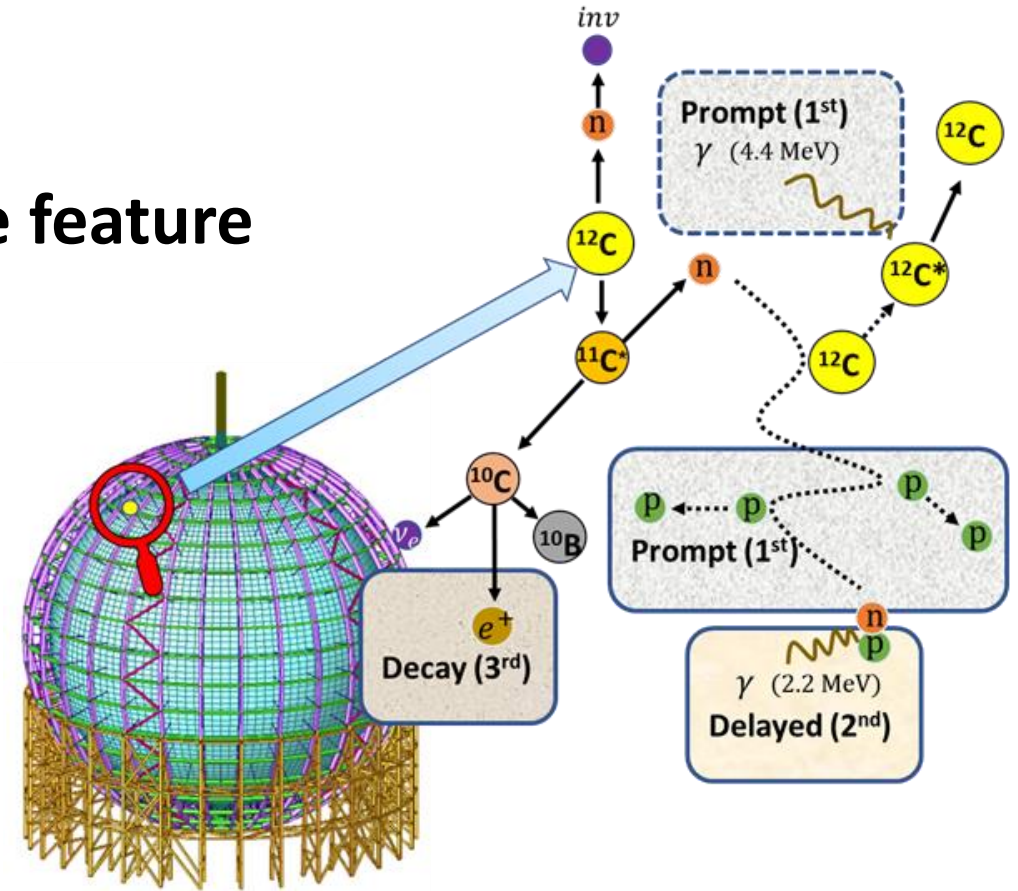
➤ Bounded neutrons in  $^{12}\text{C}$  : two invisible decay modes

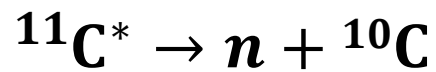
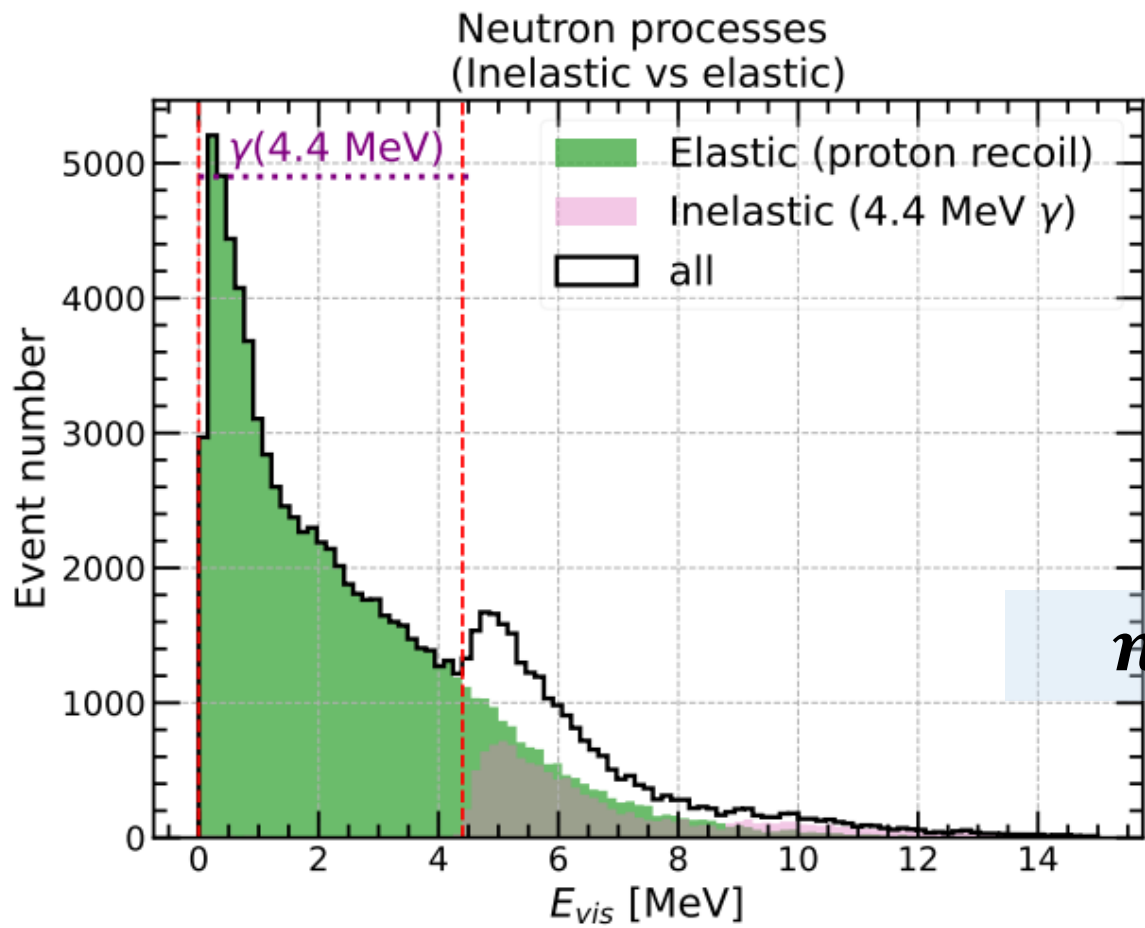
- $n \rightarrow inv$  ( $^{12}\text{C} \rightarrow ^{11}\text{C}^*$ )
- $nn \rightarrow inv$  ( $^{12}\text{C} \rightarrow ^{10}\text{C}^*$ )

➤ De-excitation modes have **triple coincidence feature**

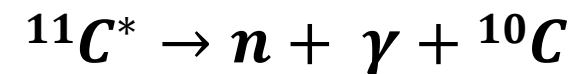
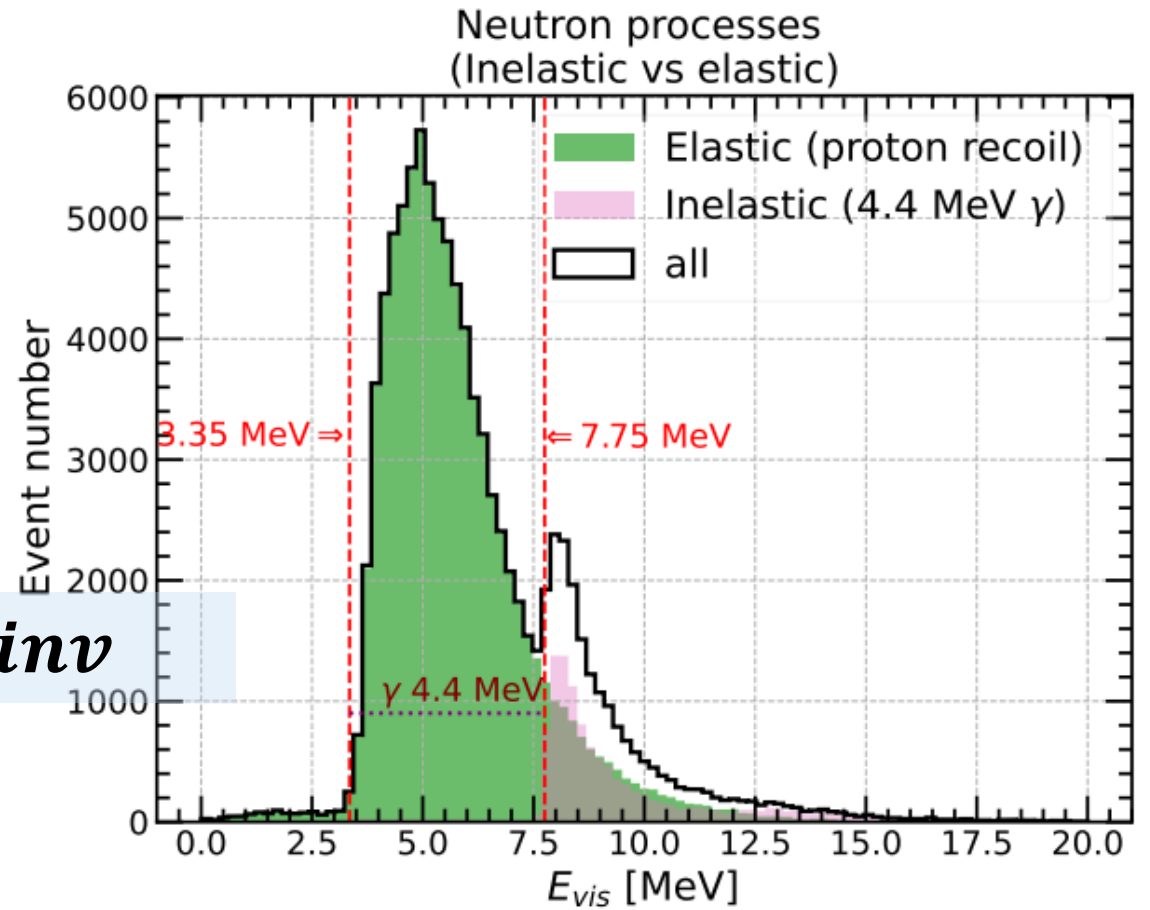
$^{11}\text{C}^* \rightarrow n +$	$^{10}\text{C}$	$(Br_{n1} = 3.0\%)$
$^{11}\text{C}^* \rightarrow n + \gamma +$	$^{10}\text{C}$	$(Br_{n2} = 2.8\%)$
$^{10}\text{C}^* \rightarrow n +$	$^9\text{C}$	$(Br_{nn1} = 6.2\%)$
$^{10}\text{C}^* \rightarrow n + p +$	$^8\text{B}$	$(Br_{nn2} = 6.0\%)$

Yuri Kamyskov, Edwin Kolbe PRD 67, 076007 (2003)



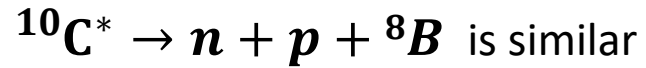
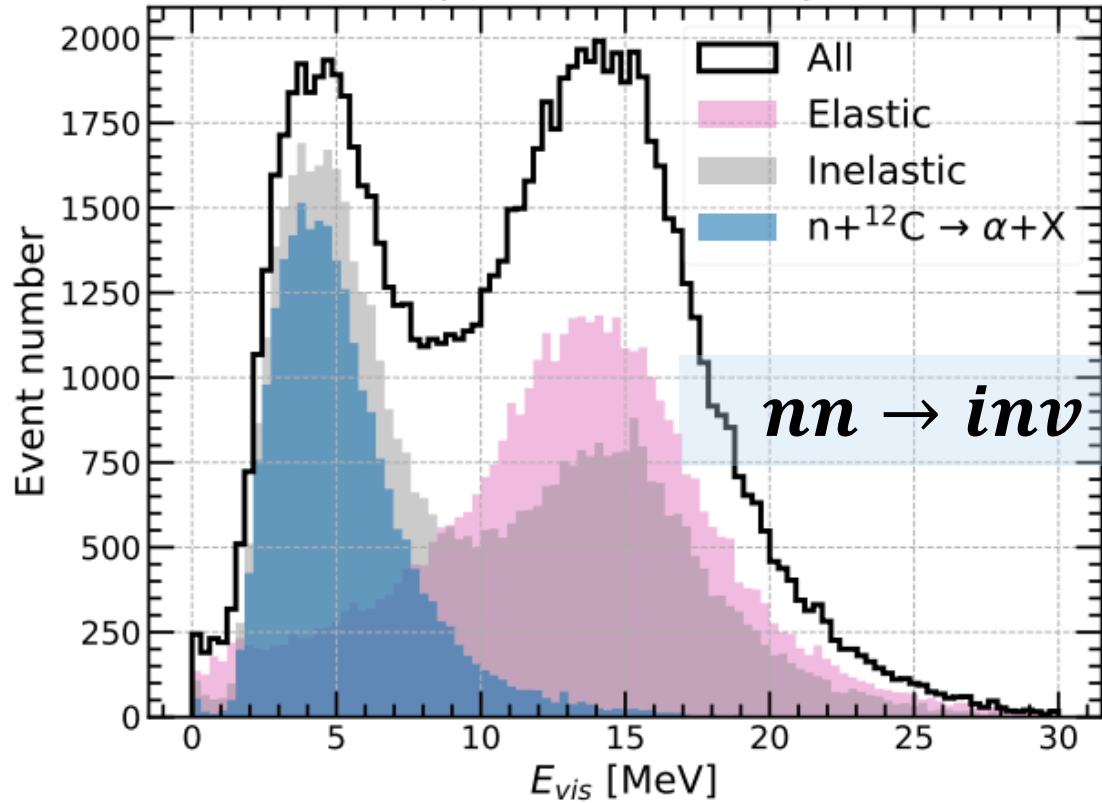


$n \rightarrow inv$

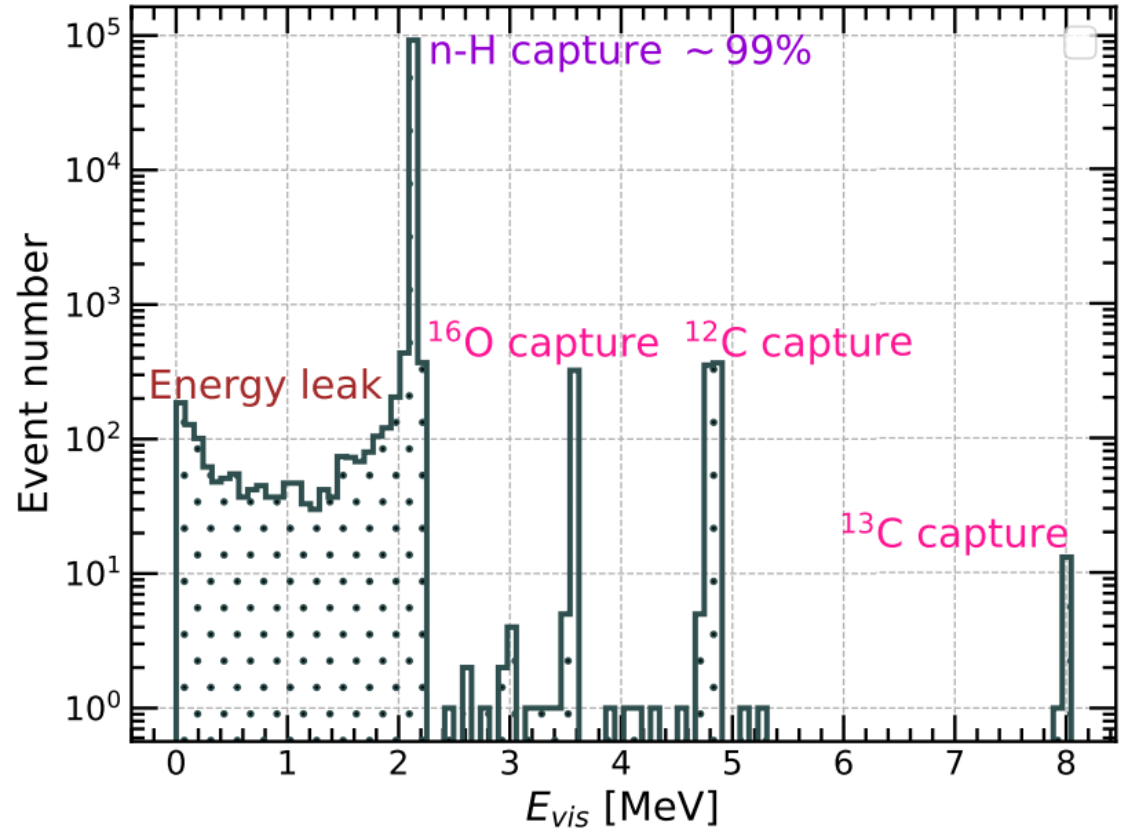


## Prompt signal

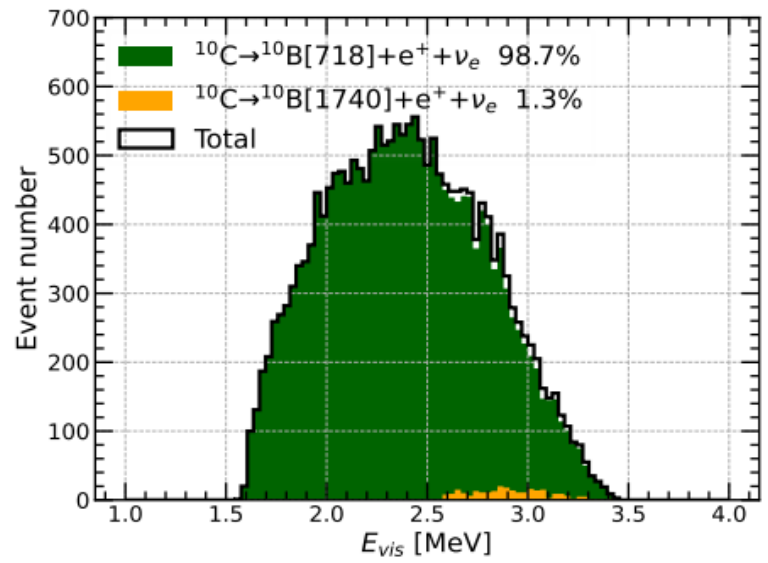
Neutron processes  
(Inelastic vs elastic)



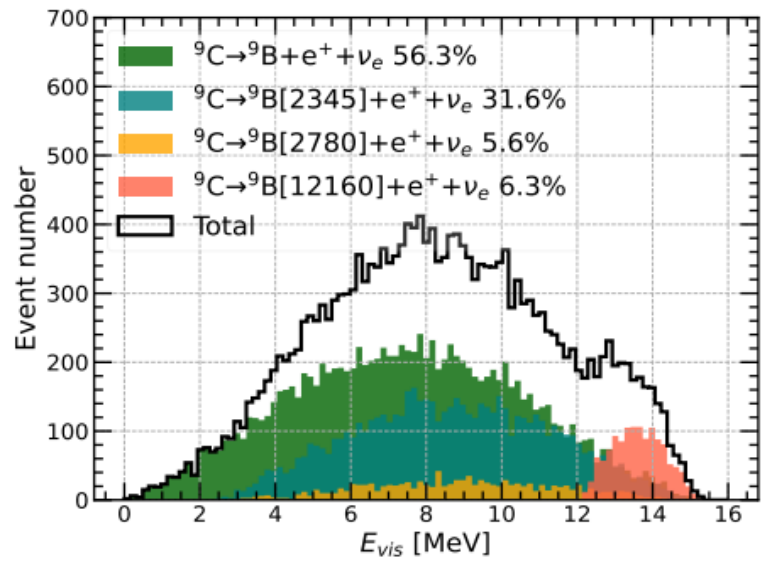
## Delayed signal



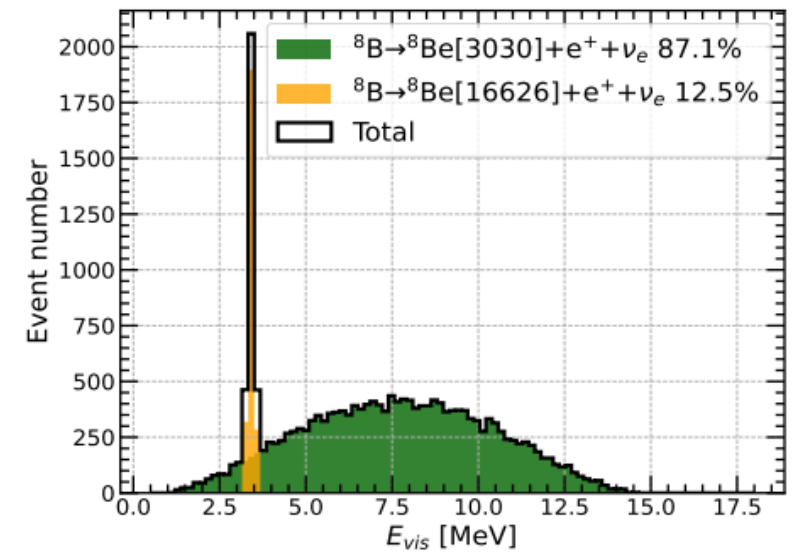




$n \rightarrow inv$



$nn \rightarrow inv$



➤ Six background sources

➤ Single

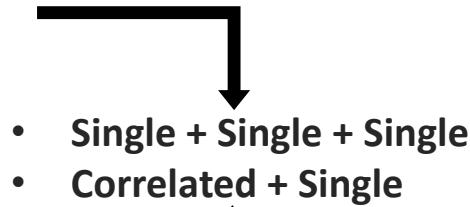
- Radioactivity
- Long-lived isotopes

➤ Correlated (Prompt-Delayed)

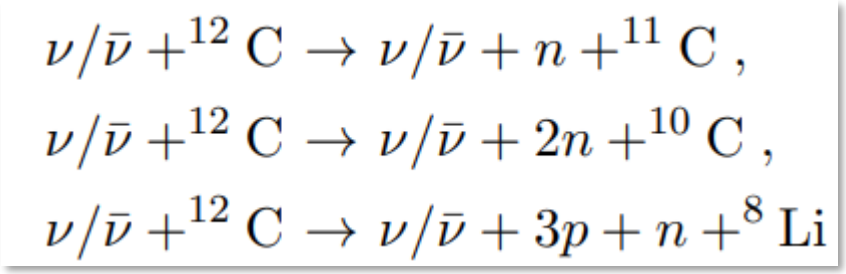
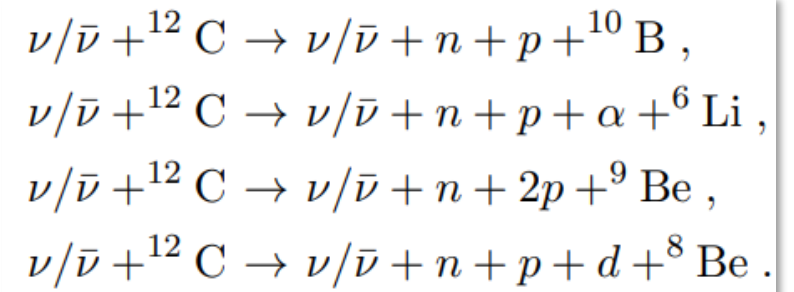
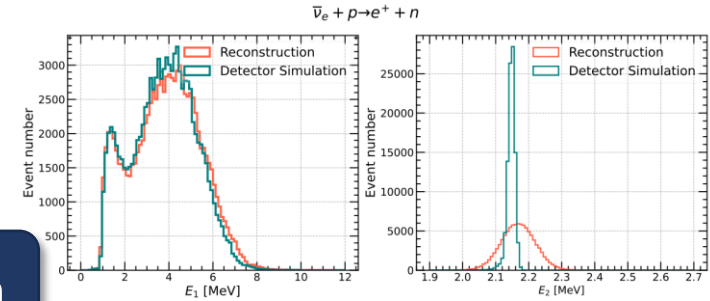
- IBD (Inverse Beta decay)
- **Atm - ν NC** (atmospheric neutrino neutral current)
- Long-lived isotopes (Li9/He8) from cosmic muons
- Fast neutron
- Alpha-n

➤ Triple (Prompt-Delayed-Decay)

- **Atm - ν NC**



**Combination**



➤ Consider **10 years** data taking

- Signal rate
  - from the final sensitivity result

## Selection criteria

Quantity	$n \rightarrow inv$	$nn \rightarrow inv$
$R_{1,2,3}$ [m]	$< 16.7$	$< 16.7$
$E_1$ [MeV]	0.7-12	0.7-30
$E_2$ [MeV]	1.9-2.5	1.9-2.5
$E_3$ [MeV]	1.5-3.5	3.0-16.0
$\Delta T_{12}$ [ms]	$< 1$	$< 1$
$\Delta T_{23}$ [s]	0.002-100	0.002-3.0
$\Delta R_{12}$ [m]	$< 1.5$	$< 1.5$
$\Delta R_{23}$ [m]	$< 1.5$	$< 1.5$
$\Delta R_{13}$ [m]	$< 1.0$	$< 1.0$

$i = 1, 2, 3$  represents the Prompt, Delayed, Decay signal

Notations of the interested parameters

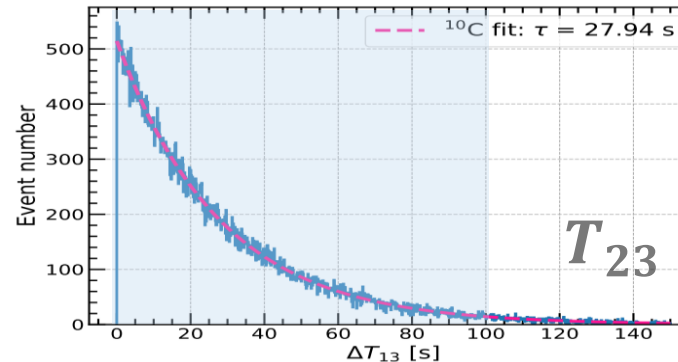
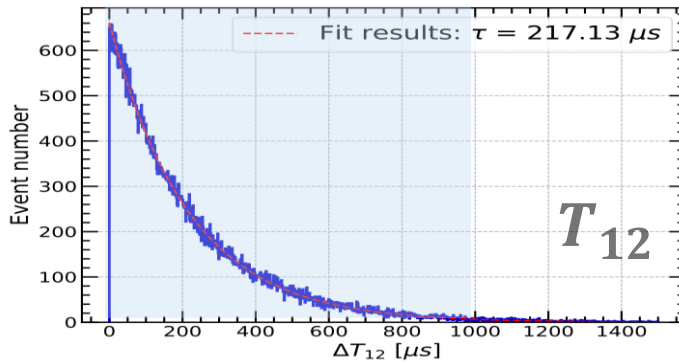
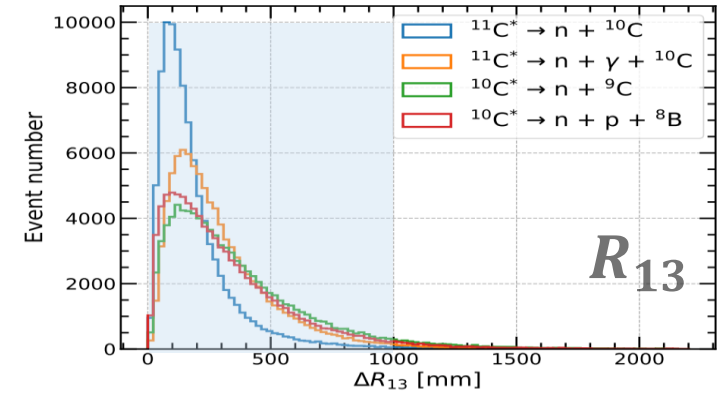
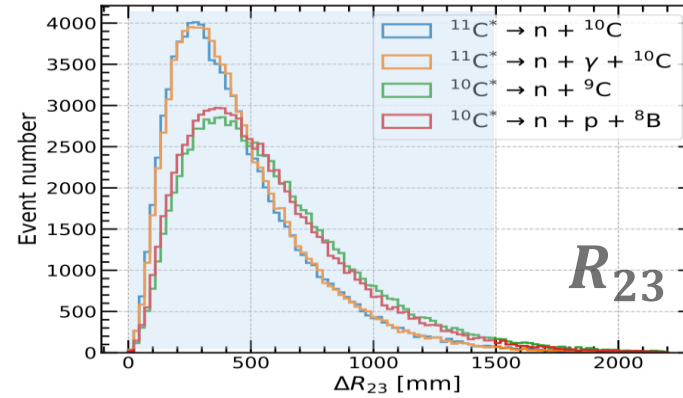
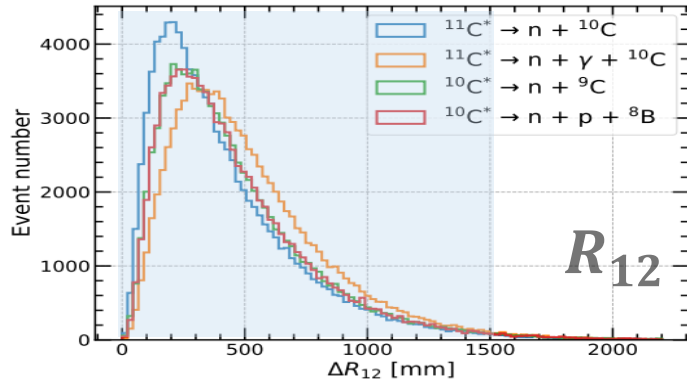
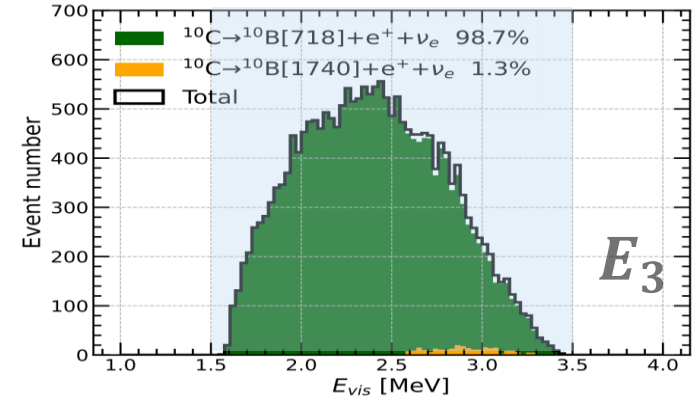
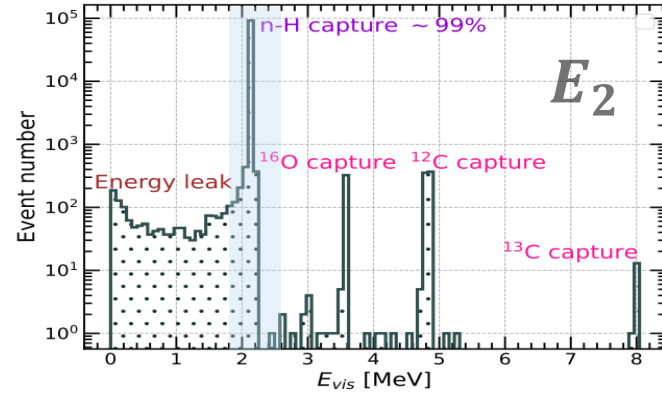
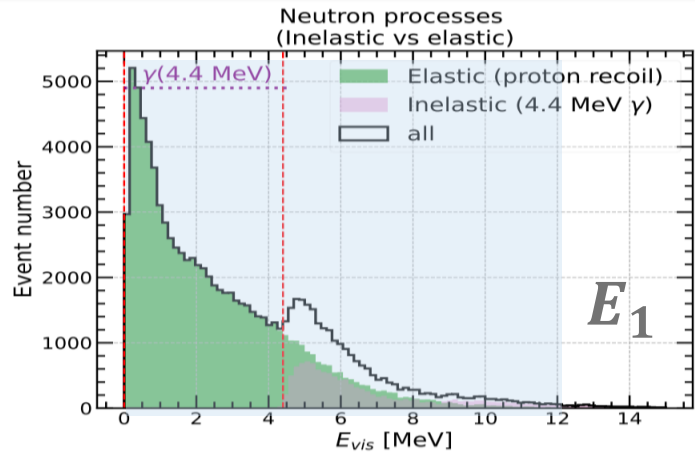
$E_i$  ( $i = 1, 2, 3$ ): Reconstruct energy

$R_i$  ( $i = 1, 2, 3$ ): Radial position

$\Delta R_{ij}$  ( $i, j = 1, 2, 3, i < j$ ): Distance between signal pair ( $i^{th}, j^{th}$ )

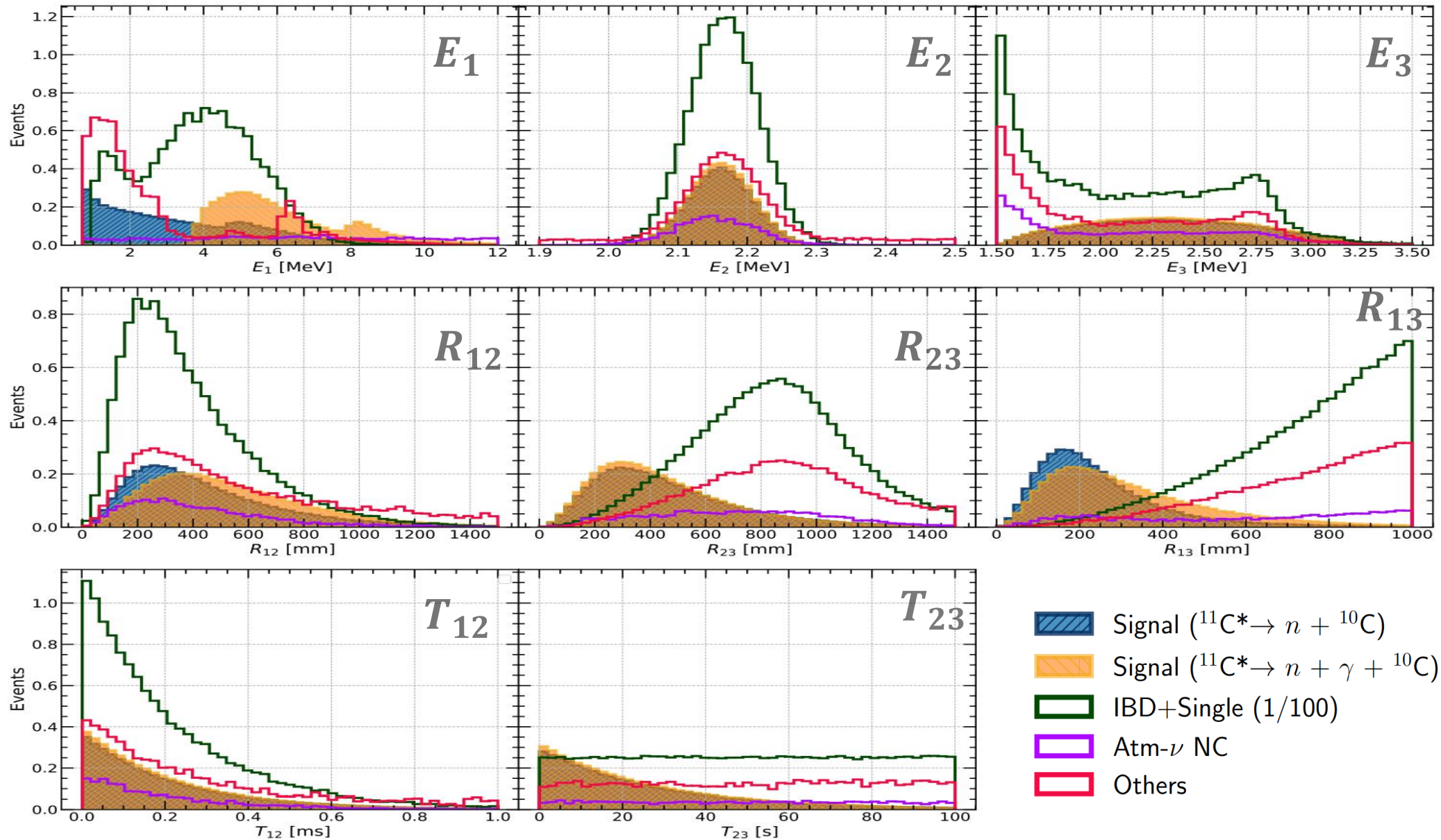
$\Delta T_{ij}$  ( $i, j = 1, 2, 3, i < j$ ): Time interval between signal pair ( $i^{th}, j^{th}$ )

# Signal characteristic of ( $n \rightarrow inv$ )

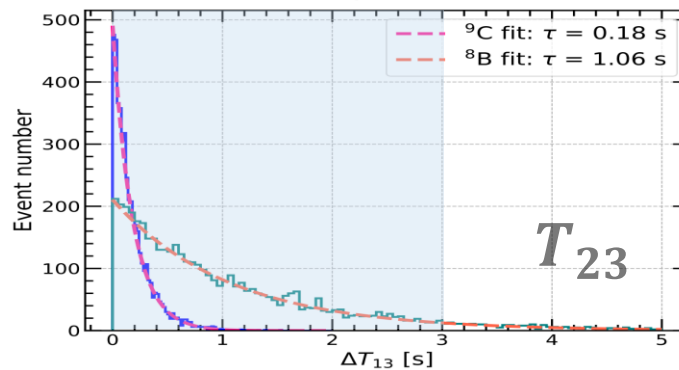
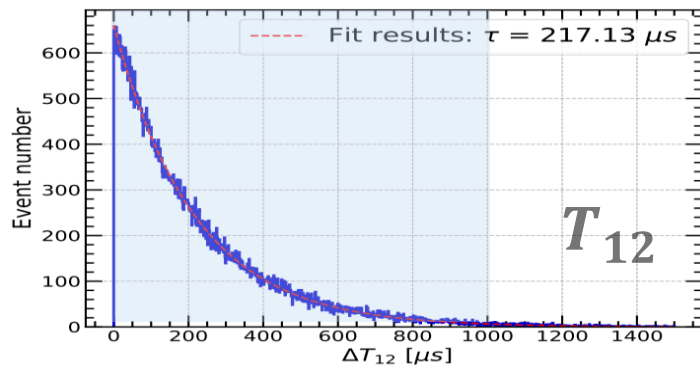
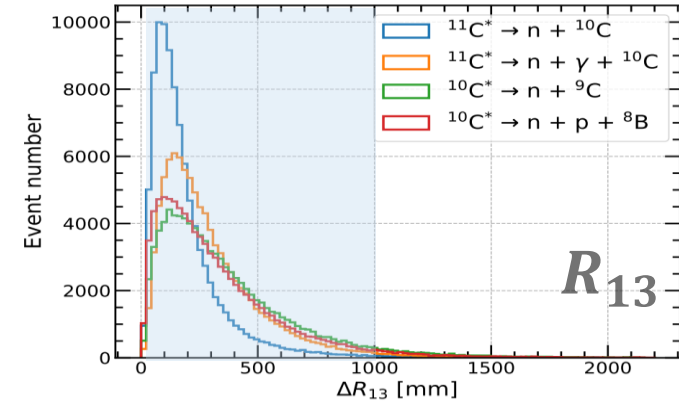
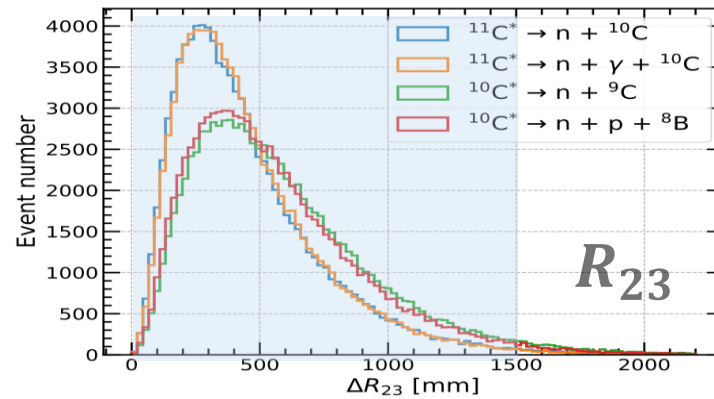
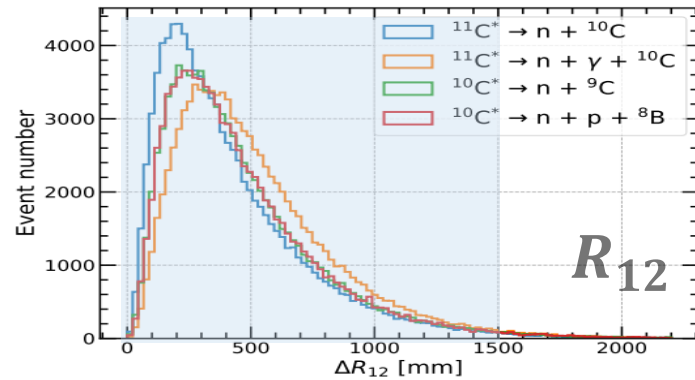
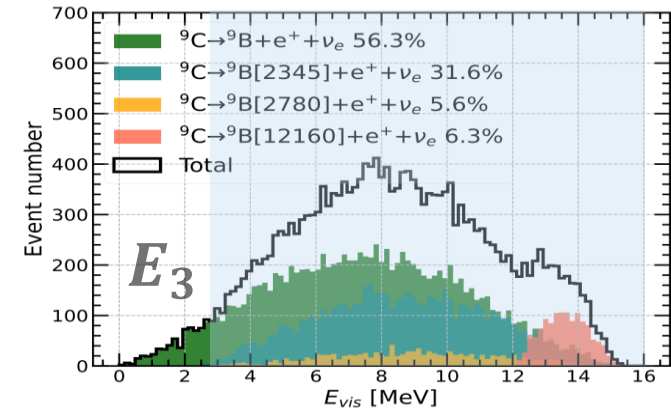
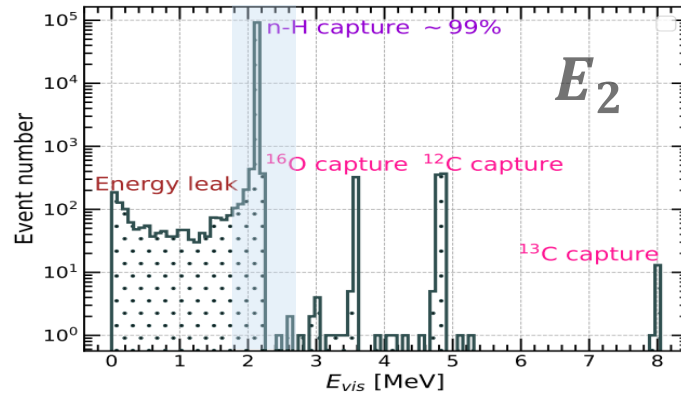
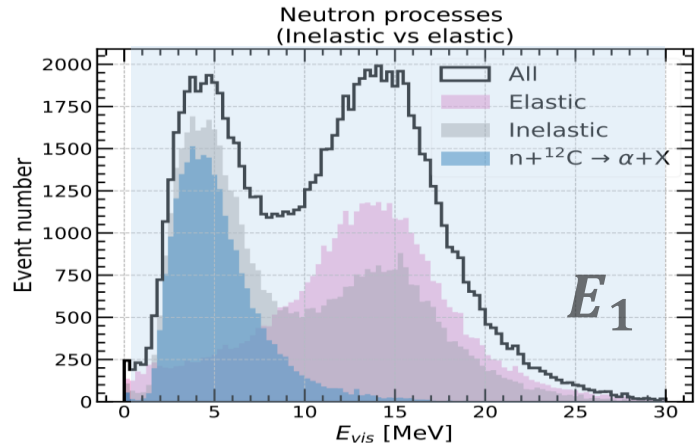


Select region

# Event selection ( $n \rightarrow inv$ )



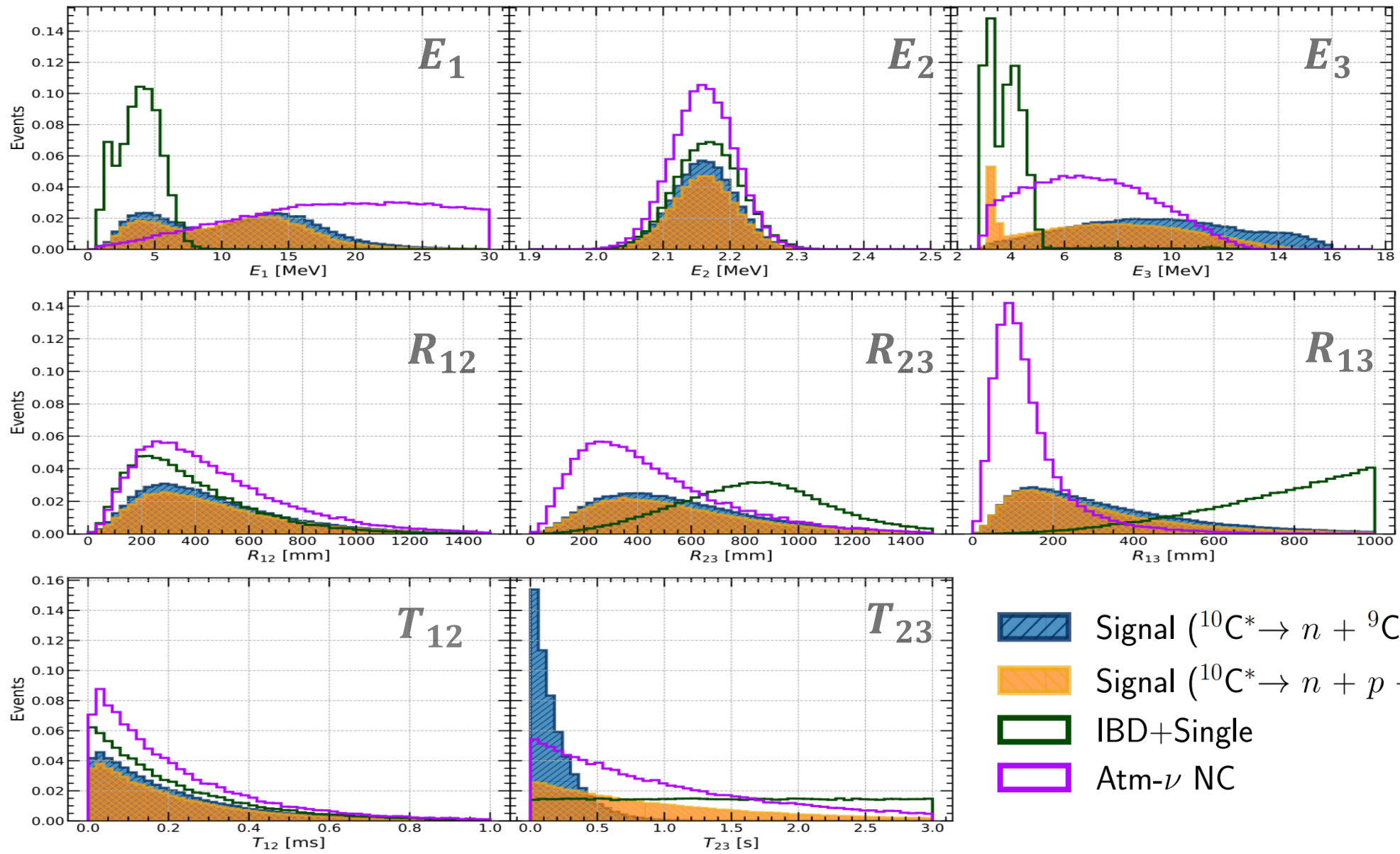




Select region



# Event selection ( $nn \rightarrow inv$ )



- Consider **10 years** data taking
  - Signal rate
    - from the final sensitivity result

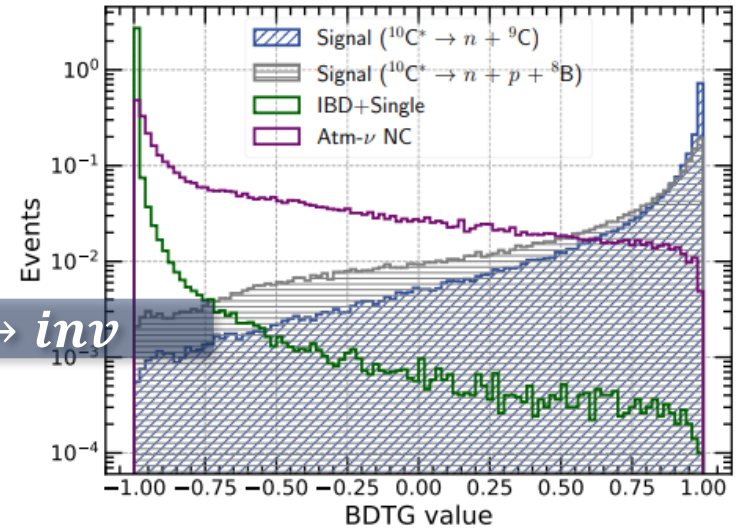
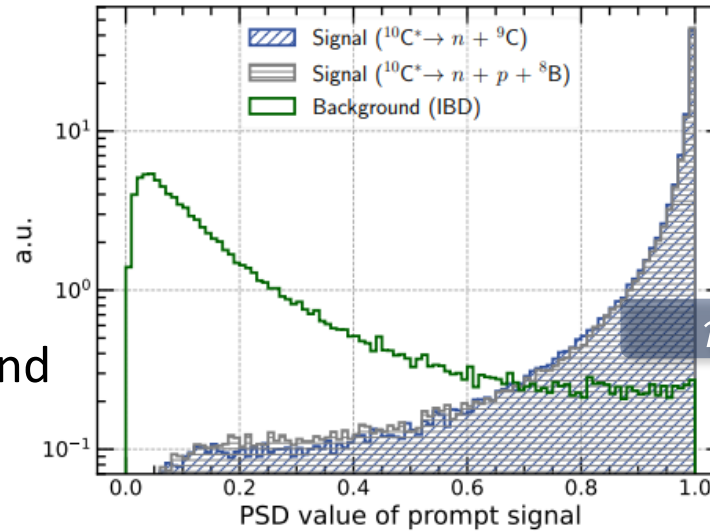
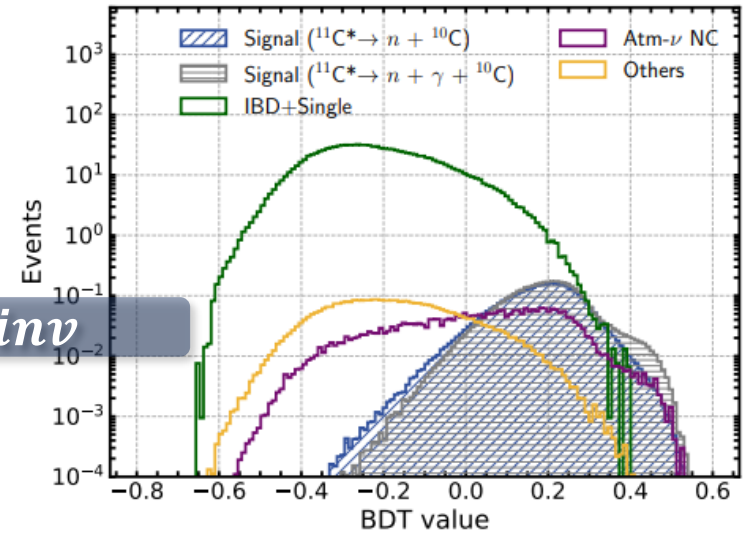
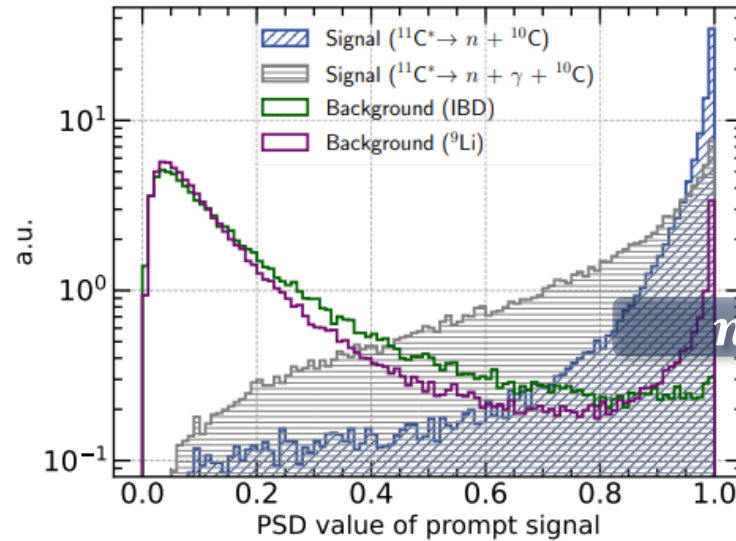
## Signal efficiencies

Selection Criterion	$n \rightarrow inv$		$nn \rightarrow inv$	
	$^{11}\text{C}^* \rightarrow n + ^{10}\text{C}$	$^{11}\text{C}^* \rightarrow n + \gamma + ^{10}\text{C}$	$^{10}\text{C}^* \rightarrow n + ^9\text{C}$	$^{10}\text{C}^* \rightarrow n + p + ^8\text{B}$
All triple signals	100	100	100	100
Muon Veto	$65.7 \pm 0.2$	$65.5 \pm 0.2$	$80.8 \pm 0.2$	$78.3 \pm 0.2$
Fiducial Volume	$83.5 \pm 0.4$	$82.7 \pm 0.4$	$82.9 \pm 0.4$	$83.1 \pm 0.4$
Event Selection	$75.4 \pm 0.9$	$89.7 \pm 0.3$	$89.2 \pm 0.3$	$83.5 \pm 0.3$
Multiplicity Cut	$93.8 \pm 0.1$	$93.8 \pm 0.1$	$99.9 \pm \mathcal{O}(10^{-4})$	$99.9 \pm \mathcal{O}(10^{-4})$
<b>Combined Selection</b>	$38.8 \pm 0.5$	$45.6 \pm 0.3$	$59.7 \pm 0.4$	$54.3 \pm 0.4$

## ➤ Two suppression method

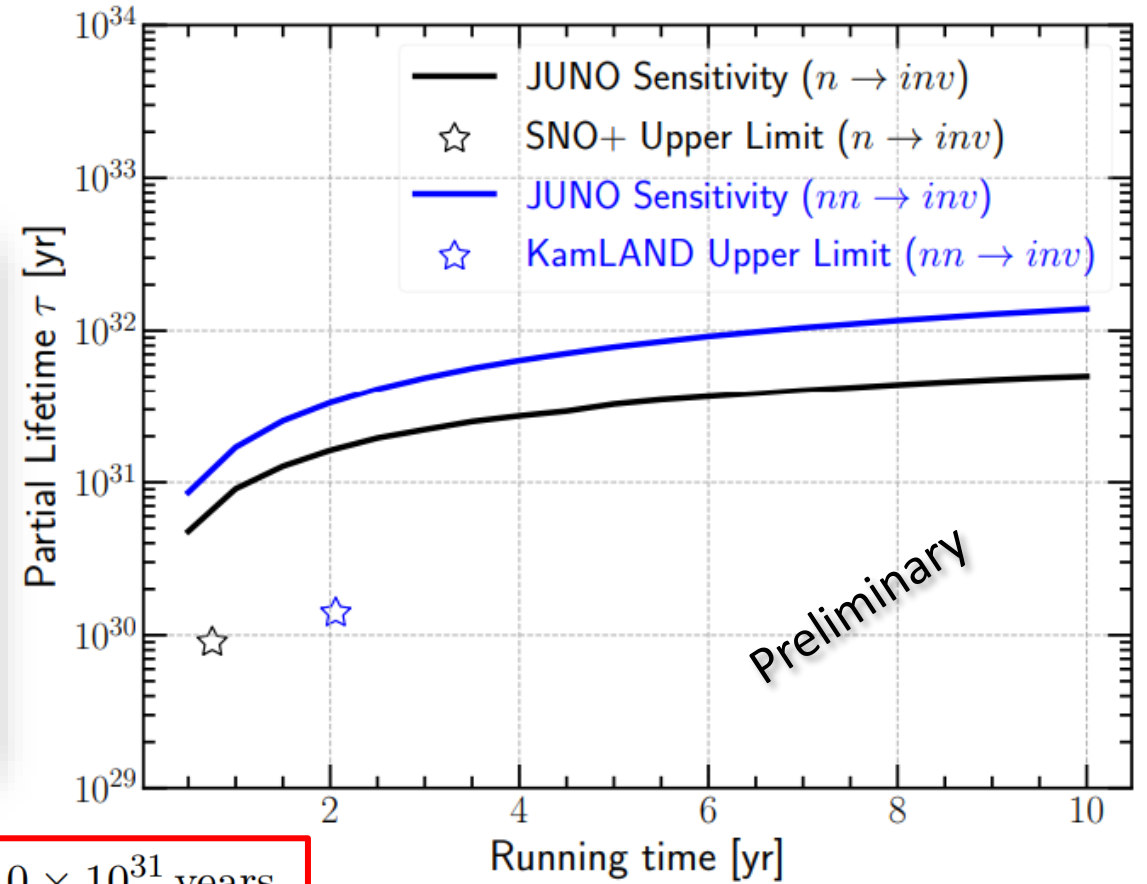
- **Pulse Shape Discrimination**
  - Particle's emission photon time are different
- **Multi Variate Analysis**
  - Combine multidimensional features

- Both PSD and MVA have **good performance**
  - Effectively suppress background



## Summary of Signal and background

Backgrounds (10 years)	$n \rightarrow inv$		$nn \rightarrow inv$	
	Basic selection	PSD + MVA	Basic selection	PSD + MVA
IBD + Single	$1235 \pm 50$	$2.72 \pm 0.10$	$3.01 \pm 0.09$	$0.0110 \pm 0.0003$
Atm- $\nu$ NC	$3.0 \pm 1.1$	$0.93 \pm 0.67$	$4.3 \pm 3.5$	$0.55 \pm 0.63$
$^{13}\text{C}(\alpha,n)^{16}\text{O}$ + Single	$3.4 \pm 1.4$	$0.036 \pm 0.013$	–	–
$^9\text{Li}/^8\text{He}$ + Single	$1.55 \pm 0.39$	$0.29 \pm 0.17$	$0.13 \pm 0.13$	$0.13 \pm 0.13$
Accidental	$1.46 \pm 0.05$	$0.095 \pm 0.004$	–	–
Total	$1244 \pm 50$	$4.07 \pm 0.68$	$7.4 \pm 3.5$	$0.69 \pm 0.64$
Signal efficiency (%)	$n \rightarrow inv$		$nn \rightarrow inv$	
	Basic selection	PSD + MVA	Basic selection	PSD + MVA
$\epsilon_{n(nn)1}$	$35.6 \pm 0.2$	$23.5 \pm 0.2$	$54.0 \pm 0.3$	$48.2 \pm 0.3$
$\epsilon_{n(nn)2}$	$43.6 \pm 0.3$	$30.3 \pm 0.3$	$49.2 \pm 0.3$	$36.3 \pm 0.3$



$$\tau/B(n \rightarrow inv) > 5.0 \times 10^{31} \text{ years,}$$

$$\tau/B(nn \rightarrow inv) > 1.4 \times 10^{32} \text{ years.} \quad \mathbf{10 \text{ years}}$$

**An order of magnitude improvement to the current best limits in 2 years data taking**

- JUNO is a large LS detector
  - 20 kton LS
    - $1.45 \times 10^{33}$  free protons,  $5.30 \times 10^{33}$  bound protons/neutrons
- Competitive sensitivities for nucleon decay (some channels)
  - **Nucleon decay (JUNO 10-year sensitivity)**
    - $\tau/B(p \rightarrow \bar{\nu} K^+) > 9.6 \times 10^{33}$  year at 90% C.L.
    - $\tau/B(n \rightarrow inv) > 5.0 \times 10^{31}$  year at 90% C.L.
    - $\tau/B(nn \rightarrow inv) > 1.4 \times 10^{31}$  year at 90% C.L.
- JUNO construction near completion, overcoming challenges
- JUNO has the potential to study nucleon decay and test new physics





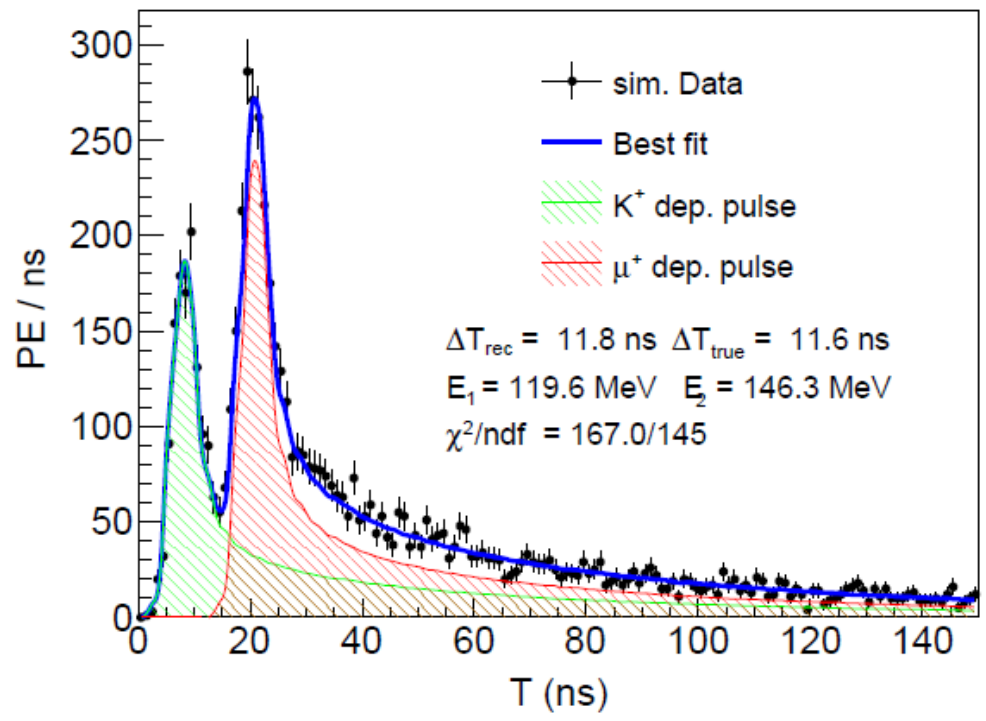
# Thank you for your attention!



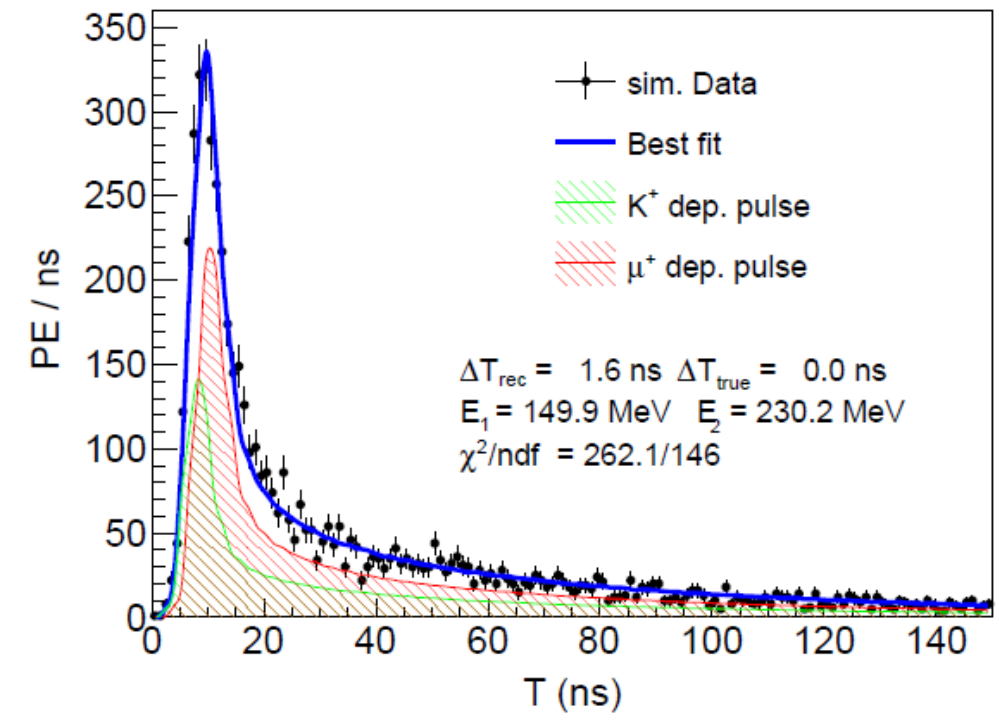


# Backup

# ➤ Multi-pulse Fitting

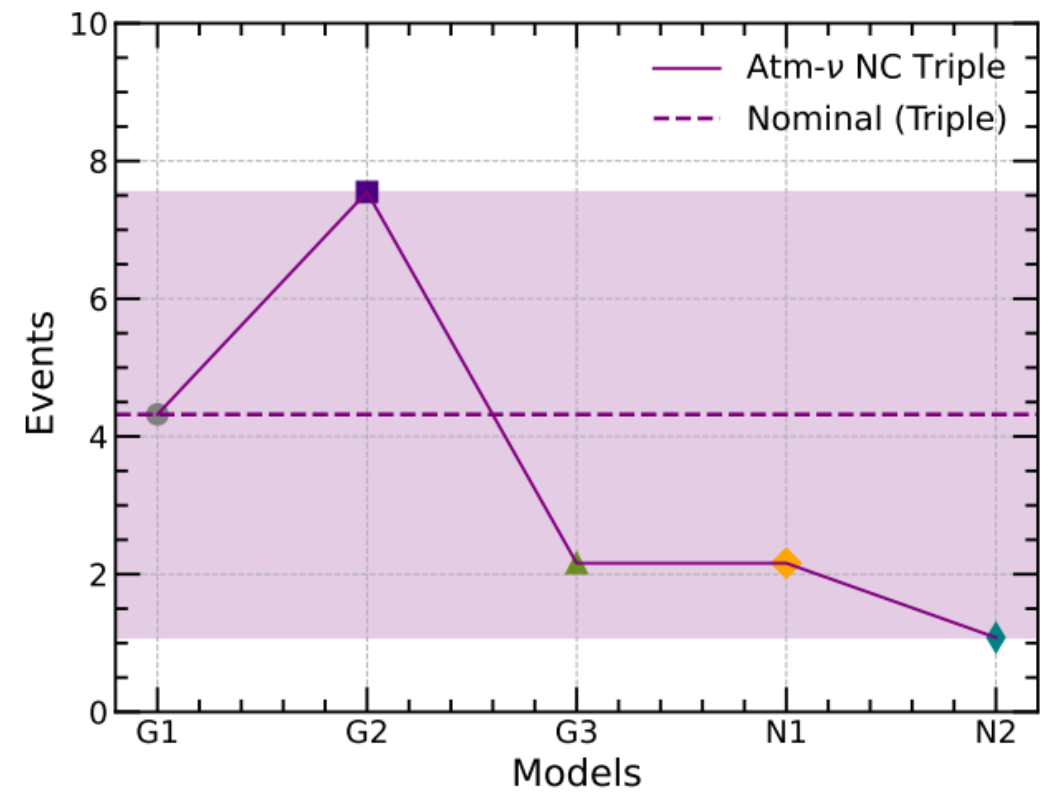
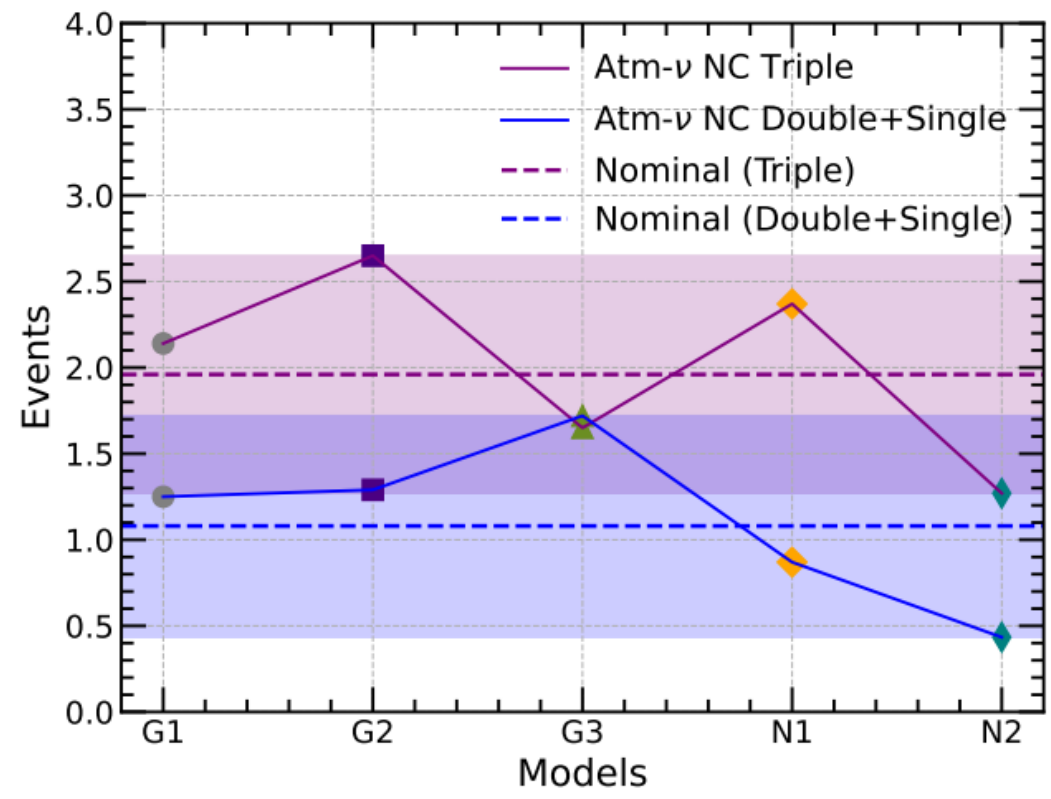


(a) Hit time spectrum of a proton decay event



(b) Hit time spectrum of an atmospheric  $\nu$  event

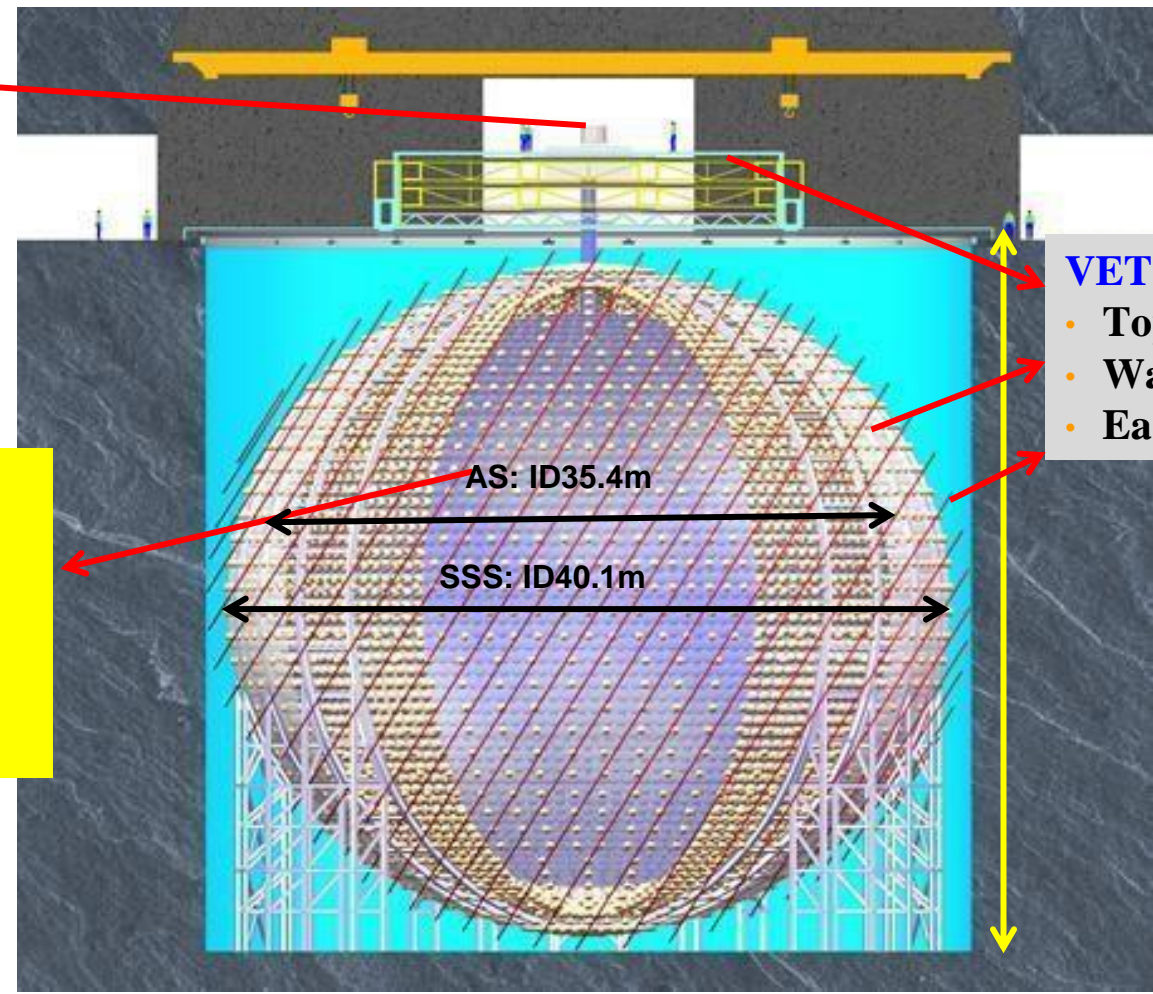
# ➤ NC background



Calibration

**Central detector**

- Steel structure (SSS)
- Acrylic sphere(AS) + 20kt Liquid scintillator
- 17612 20" PMT
- 25600 3" PMT



**VETO system**

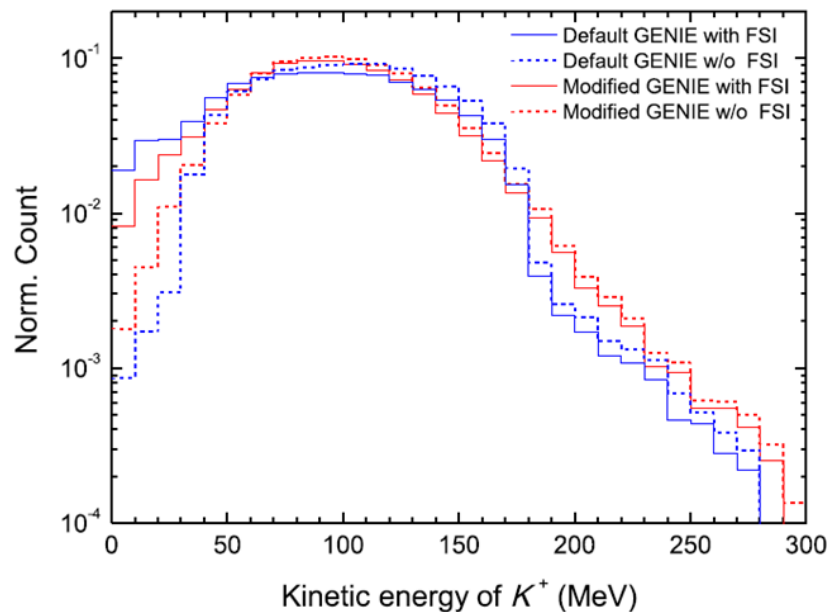
- Top Tracker
- Water VETO with 2400 20" PMT
- Earth Magnetic Field shielding coils

# ➤ Search for $p \rightarrow \bar{\nu} K^+$ in JUNO

**20 kton LS:** Free proton:  $1.45 \times 10^{33}$   
 Bound proton:  $5.30 \times 10^{33}$

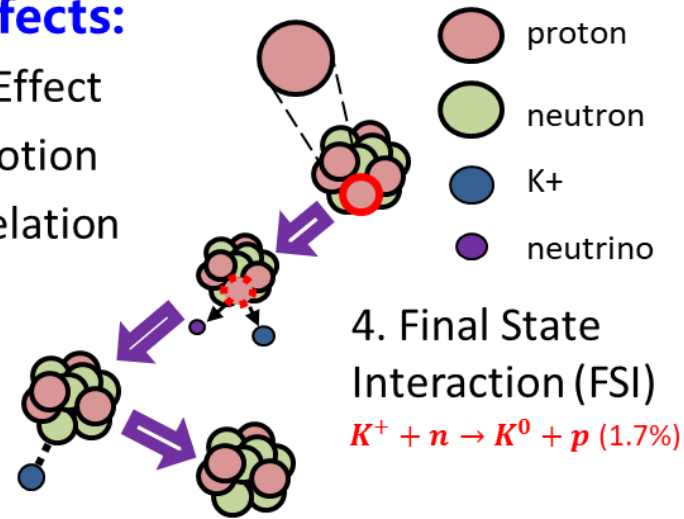
## Kinetic energy of $K^+$

Free proton  $\rightarrow$  105 MeV  
 Bound proton:  $\downarrow$



## Nuclear Effects:

1. Binding Effect
2. Fermi Motion
3. NN correlation



4. Final State Interaction (FSI)  
 $K^+ + n \rightarrow K^0 + p$  (1.7%)

5. De-excitation of remaining nuclear:  
 could emit  $\gamma/p/n$ .

- **Modify GENIE generator**
- **Implement de-excitation with TALYS**

H. Hu, W.L. Guo et al, PLB 831, 137183(2022)