#### **DUNE Low Energy Physics** with Solar and Supernova Neutrinos

Gleb Sinev on behalf of DUNE Collaboration CoSSURF 2024 May 16, 2024



#### Outline

- Deep Underground Neutrino Experiment
- Low-energy neutrino interactions
- Supernova neutrinos
- Solar neutrinos
- Low-Background Module
- Conclusions





#### Deep Underground Neutrino Experiment (DUNE)

- 1.2 MW (upgradeable to 2.4 MW) LBNF neutrino beam
- Near Detector (ND) complex with multiple detectors
- Far Detector (FD): 4 modules, 70 kt liquid Ar total
  - 1 mile underground at SURF, excavation completed



#### P5 report

**Recommendation 1:** As the highest priority independent of the budget scenarios, complete construction projects and support operations of ongoing experiments and research to enable maximum science. We reaffirm the previous P5 recommendations on major initiatives:

b. The first phase of DUNE and PIP-II to open an era of precision neutrino measurements that include the determination of the mass ordering among neutrinos. Knowledge of this fundamental property is a crucial input to cosmology and nuclear science (*elucidate the mysteries of neutrinos*, section 3.1).

**Recommendation 2:** Construct a portfolio of major projects that collectively study nearly all fundamental constituents of our universe and their interactions, as well as how those interactions determine both the cosmic past and future.

b. A re-envisioned second phase of DUNE with an early implementation of an enhanced 2.1 MW beam—ACE-MIRT—a third far detector, and an upgraded near-detector complex as the definitive long-baseline neutrino oscillation experiment of its kind (section 3.1).

**Recommendation 4:** Support a comprehensive effort to develop the resources—theoretical, computational, and technological—essential to our 20-year vision for the field. This includes an aggressive R&D program that, while technologically challenging, could yield revolutionary accelerator designs that chart a realistic path to a 10 TeV pCM collider.

e. Conduct R&D efforts to define and enable new projects in the next decade, including detectors for an e<sup>+</sup>e<sup>−</sup> Higgs factory and 10 TeV pCM collider, Spec-S5, DUNE FD4, Mu2e-II, Advanced Muon Facility, and line intensity mapping (sections 3.1, 3.2, 4.2, 5.1, 5.2, and 6.3).



https://www.usparticlephysics.org/ 2023-p5-report/assets/images/hero-full.jpg

#### **Strong endorsement!**

https://www.usparticlephysics.org/2023-p5-report/full-list-of-recommendations



# **Particle detection in LArTPC**

- Neutrinos interact DUNE uses Liquid-Argon Time-Projection in argon, produce charged particles Chamber (LArTPC) as detector technology
- Argon scintillates, light is quickly detected by photon detectors
- Charged particles ionize argon, electrons slowly drift to anode
- Anode is instrumented (readout wires/strips)
  - Combining with light, reconstruct 3D events





#### Low-energy v event in DUNE

Eur.Phys.J.C 81 (2021) 5, 423



Simulation with MARLEY (Comput. Phys. Commun. 269, 108123, arXiv:2101.11867 [nucl-th])

- Low-energy neutrino events typically leave several-cm track (over several channels) in TPC
- Primary track may be surrounded by deexcitation particles that increase event size to tens of cm (γs) or even m (ns)
- Radiological backgrounds make reconstruction challenging



#### Low-energy v interactions in Ar

- Charged current (CC)
  - $\nu_e + \frac{40}{18}Ar \rightarrow \frac{40}{19}K^* + e^-$ 
    - Dominant channel
  - $\bar{\nu}_e + {}^{40}_{18}Ar \rightarrow {}^{40}_{17}Cl^* + e^+$
- Neutral Current (NC)  $\nu_x + {}^{40}_{18}Ar \rightarrow {}^{40}_{18}Ar^* + \nu_x$
- Elastic scattering (ES)  $\nu_x + e \rightarrow \nu_x + e$

Eur.Phys.J.C 81 (2021) 5, 423





# **Uncertainty in interactions**

Phys. Rev. D 107, 112012 (2023)

- Models disagree by ~order of magnitude on cross sections of v<sub>e</sub>CC interactions
- May significantly affect low-energy studies in DUNE
- Need measurements

8



# Supernova neutrinos

- Neutrinos produced from core-collapse supernova burst (SNB)
- Neutrinos from only one supernova have been detected: SN 1987A
  - ~2 dozen neutrinos
- Time and energy distributions will carry information about supernova physics
  - May also help us understand neutrino physics





Eur.Phys.J.C 81 (2021) 5, 423

## **Supernova neutrinos in DUNE**

Most interactions will be v<sub>e</sub>CC

90% C.L.

7.00kpc

4.00kpc

0.9

0.8

 $\epsilon$  (10<sup>53</sup> erg) 9.0  $\epsilon$  (20) 9.0  $\epsilon$  (10<sup>53</sup> erg)

0.4

0.3

10

- Depending on neutrino oscillations, neutronization burst may be clearly observable or not present at all
- Fitting energy spectrum may help narrow down supernova models

**×** Truth:  $\langle E_{..} \rangle$  = 9.5 MeV,  $\varepsilon$  = 5e52 erg





Events per bin

# **Supernova neutrinos in DUNE**

 Expect to detect 1,000s of neutrinos from supernova close to Milky Way center Mostly v<sub>e</sub>CC, unique to DUNE/LAr detectors!

- On order of 1 event from Andromeda
- Many more than were detected from SN 1987A
- Possibility of pointing (res. of ~5°) and early warning



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## **Solar neutrinos**

- DUNE 10<sup>13</sup> SFII-GS98 + ecCNO 10<sup>12</sup> Solar Neutrino Spectra ( $\pm 1\sigma$ ) 10<sup>11</sup> pp[±0.6%] 10<sup>10</sup> <sup>7</sup>Be[±7%] <sup>-</sup>lux [cm<sup>-2</sup> s<sup>-1</sup> (100 kev<sup>-1</sup>)] <sup>7</sup>Be[±7%] 10<sup>9</sup> pep[±1.2%] 10<sup>8</sup> **٧[±14%]** 10 10<sup>6</sup> eN[±14%] <sup>5</sup>O[±15%] eO+eF[±15%] 10<sup>5</sup> B[±14%] <sup>7</sup>F[ ±19%] 10<sup>4</sup>  $10^{3}$ hep[±30%]  $10^{2}$ 10 0.5 0.1 0.2 1.0 2.0 5.0 10.0 20.0 Neutrino Energy in MeV
- https://pdg.lbl.gov/2023/reviews/rpp2023-rev-neutrino-mixing.pdf

- v<sub>e</sub> neutrinos produced in nuclear reactions in Sun
- Different reactions produce components with different spectra and intensity



# Solar neutrinos in DUNE

- DUNE will observe thousands of solar neutrinos/year
- Challenge is to mitigate radiological backgrounds **DUNE** Preliminary to not drown signal

10

10

Events / 400 kton-years

- Currently expect threshold ~9 MeV visible energy
- <sup>8</sup>B and hep neutrinos will be detectable
  - hep neutrinos have not been observed
  - No rad background for hep v
- Possible to study oscillations



## **Solar neutrino oscillations**

 DUNE has potential to use solar neutrinos to improve knowledge of oscillations parameters, as well as to probe beyond-standard-model physics



F. Capozzi, S.W. Li, G. Zhu, J.F. Beacom, Phys. Rev. Lett. 123, 131803 (2019)



# Low-background module (SLoMo)

- SURF Low Background Module
  - One concept for Module of Opportunity
  - Improved capabilities for low-energy studies
- Significant reduction of radiological backgrounds
  - 3 kt of radiopure underground argon
  - constructed out of radiopure materials
  - water shielding
- better light coverage





T. Bezerra et al.

#### T. Bezerra et al. JPhysG Vol 50, 6, 9-May-2023

# Physics with SLoMo

- Improved measurement of solar neutrino-oscillation parameters
- Studying non-standard neutrino oscillations with solar neutrinos
- Measuring CNO solar neutrinos
- Detecting CEvNS from SNB neutrinos
- Other topics
  - Dark matter including seasonal variations
  - Neutrinoless double-beta decay
- A. Mastbaum, F. Psihas, J. Zennamo PhysRevD 106, 092002 (2022)

108

106

104

10<sup>2</sup>

100

0.00

0.25

Events / 3.0 kTonne-yr / 0.05 MeV

CNOhiN

CNOIoM 7Be

Ar42 Ba/L/1500\*9.2E-5/1E5

0.50

0.75

1.00

Deposited Energy [MeV]

1.25

**B** 

per Ar39 Bq/L/1500 1 25-1.45 MeV regio

1.35

1.50

1.75

2.00







#### Conclusions

- DUNE has potential to become powerful tool to study low-energy neutrino physics
  - Large-size LArTPC is suited for detection of supernova and solar neutrinos
  - Radiological-background mitigation is important for reaching this goal
  - SLoMo as Module of Opportunity may significantly improve its capabilities



# **DUNE Collaboration**

1,400+ people from
200+ institutions in 30+ countries





DUNE Collaboration meeting at CERN, 2024

18 05/16/2024 Gleb Sinev | DUNE Low Energy Physics with Solar and Supernova Neutrinos



#### JPhysG Vol 50, 6, 9-May-2023 SNB CEVNS glow in SLOMO



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#### SLoMo supernova reach



DUNE