### **Reactor Antineutrino Oscillations and JUNO**

5th Conference on Science at the Sanford Underground Research Facility May 14-17 2024 Roberto Mandujano rcmanduj@uci.edu University of California, Irvine On Behalf of the JUNO Collaboration



#### **Reactors and Neutrinos**





- Intense, pure, and cost-effective source of electron antineutrinos: ~10<sup>20</sup>  $\bar{\nu}_e/s$  per  $GW_{th}$
- Produced in beta decays:  $n \rightarrow p + e^- + \bar{\nu}_e$
- Used for neutrino discovery by Reines and Cowan
- Today's talk will focus on experiments with commercial Low-Enriched Uranium reactors
  - Fissions come from <sup>235</sup>U,<sup>239</sup>Pu,<sup>241</sup>Pu,<sup>238</sup>U



### Reactor $\bar{\nu}_e$ Detection





• Reactor  $\bar{\nu}_e$  detected through Inverse Beta Decay (IBD) reaction

• 
$$\bar{\nu}_e + p \rightarrow e^+ + n$$

- Positron (prompt) signal followed by neutron capture (delayed) typically on Gd (nGd), H (nH), or C (nC)
- Temporal and spatial coincidence of prompt and delayed signals is a powerful handle to extract reactor neutrino signal





$$P_{\bar{\nu}_e \to \bar{\nu}_e}(L,E) = 1 - \sin^2 2\theta_{12} \cos^4 \theta_{13} \sin^2 \frac{\Delta m_{21}^2 L}{4E} - \sin^2 2\theta_{13} \left(\cos^2 \theta_{12} \sin^2 \frac{\Delta m_{31}^2 L}{4E} + \sin^2 \theta_{12} \sin^2 \frac{\Delta m_{32}^2 L}{4E}\right)$$

- Many avenues to explore oscillations with reactor neutrinos!
- Sensitive to  $\begin{array}{c} \theta_{13}, \theta_{12}, \|\Delta m_{21}^2\|, \|\Delta m_{31}^2\|\\ \text{and the mass ordering} \end{array}$
- Independent from  $heta_{23}, \delta_{CP}$



Note: Reactor neutrino experiments often report an ordering-independent effective mass splitting:  $|\Delta m_{32}^2| = |\Delta m_{ee}^2| - \alpha \cos^2 \theta_{12} \Delta m_{21}^2$ ,  $\alpha = +1, -1$  for NO, IO



# Measuring $\theta_{13}$





- O(10) ton liquid scintillator (LS) detectors at short (km) baselines
- Up to 8% disappearance effect
- Systematics control is key: near/far comparison cancels uncertainties in flux and correlated detection efficiencies



# Measuring $\theta_{13}$



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- Multi-zone detectors lined with photomultiplier tubes (PMTs)
- Instrumented outer shields double as muon veto
- Mineral oil as additional shielding
- Middle LS volume serves as  $\gamma$  catcher
- Innermost volume comprised of Gd-doped LS
  - Gd improves IBD signal

water, mineral oil, LS, GdLS













Phys. Rev. Lett. 130, 161802 (2023)



- Latest results using neutron capture on Gd (nGd) with 3158 days of data
- Measure  $\sin^2(2\theta_{13}), \Delta m^2_{32}$  from spectral distortion and relative rates between detectors

$$\sin^2(2\theta_{13}) = 0.0851 \pm 0.0024$$
$$\Delta m_{32}^2 = (2.466 \pm 0.060) \times 10^{-3} \,\text{eV}^2 \,(\text{NO})$$
$$\Delta m_{32}^2 = -(2.571 \pm 0.060) \times 10^{-3} \,\text{eV}^2 \,(\text{IO})$$







- Latest nGd results using 2900 day data set shown at NEUTRINO 2022
- Large reduction in statistical and systematic uncertainties from previous result

$$\frac{4000}{3000} \xrightarrow{\text{Preliminary}}_{1000} \xrightarrow{\text{P$$

$$\sin^{2}(2\theta_{13}) = 0.0892 \pm 0.0089$$
$$|\Delta m_{32}^{2}| = (2.68 \pm 0.14) \times 10^{-3} \text{ eV}^{2} \text{ (NO)}$$
$$|\Delta m_{32}^{2}| = (2.79 \pm 0.14) \times 10^{-3} \text{ eV}^{2} \text{ (IO)}$$



### **Double Chooz**



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Nature Phys. 16 (2020) 5, 558-564

- 818 live day data-set using total neutron capture
  - nGd +nH+nC
  - Control of detection systematics and increased volume (γ catcher)
- Slightly higher  $\sin^2(2\theta_{13})$  w.r.t similar experiments
  - <  $2\sigma$  difference

 $\sin^2(2\theta_{13}) = 0.105 \pm 0.014$ 





### **Present Picture**



- $\theta_{13}$  measurement with reactor  $\bar{\nu}_e$  will remain the most precise for years to come
  - Best known angle in PMNS matrix! (for now...)



Good sensitivity to  $\Delta m_{32}^2$ , in agreement with accelerator experiments







### **JUNO Overview**



- The Jiangmen Underground Neutrino Observatory (JUNO) is a large liquid scintillator neutrino detector, ~650 m underground in southern China
- 20 kton mass: the largest detector of its kind ever built
- Multi-purpose experiment with rich physics portfolio
- Main physics goals are neutrino mass ordering (NMO) determination and precision measurement of neutrino oscillation parameters







### **Detector Design**



- 35.4 m diameter acrylic sphere filled with 20 kton of liquid scintillator (LS)
  - LS designed for high light yield and low attenuation
- 17,612 20" PMTs (LPMTs) and 25,600 3" PMTs (SPMTs)
  - ~78% photo-coverage
  - ~30% detection efficiency (LPMT)
- Instrumented outer water tank and top scintillator panels
- Unprecedented 3% energy resolution at 1 MeV









- JUNO-TAO (Taishan Antineutrino Detector) will be a satellite detector
  - 44 m from 4.6 GW<sub>th</sub> reactor
  - ~1 ton GdLS fiducial volume
  - Instrumented with SiPM providing <2% at 1 MeV energy resolution and >90% photocoverage
  - Operates at -50°C
- Measure reactor antineutrino energy spectrum with excellent resolution
  - Remove possible model dependence from JUNO NMO measurement
  - Search for sterile neutrinos
  - Isotopic  $\bar{\nu}_e$  rate and shape
  - Important inputs for experiments and nuclear databases



arXiv: 2005.08745



# JUNO $\bar{\nu}_e$ Oscillations



- Observation of  $\theta_{12}, \Delta m^2_{21}, \Delta m^2_{31}$  and  $\theta_{13}$  driven oscillations
- Determination of NMO through interference effects in fine structure of oscillated spectrum (allowed by large  $\theta_{13}$ )
  - Precise energy spectrum reference from JUNO-TAO
  - Independent of  $\delta_{CP}$ , octant of  $\theta_{23}$
  - Complementary to accelerator and atmospheric measurement (different baseline, energy, and technology)
  - Reach ~5σ in combination with other experiments (*PRD 101, 032006 (2019), Sci Rep 12, 5393 (2022)*)







### **Updated Sensitivities**





- Recent sensitivity study with full treatment of systematics using best knowledge of detector response to date:
  - Updated number of reactors
  - Realistic simulation and veto efficiencies
  - Final detector overburden and location information
  - Spectral shape constraints from JUNO TAO included



# **Oscillation Parameters**





	$\Delta m_{21}^2$	$\Delta m_{31}^2$	$\sin^2\theta_{12}$	$\sin^2\theta_{13}$
JUNO 100 days	0.8%	1%	1.9%	47.9%
JUNO 6 years	0.3%	0.2%	0.5%	12.1%
KamLAND	2.4%	-	-	_
T2K	-	2.6%	-	_
SNO+SK	-	_	4.5%	_
Daya Bay	_	_	_	3.4%

JUNO Relative Uncertainty vs. Leading Experiments

• Measurement of  $\sin^2\theta_{12}$ ,  $\Delta m^2_{21}$  and  $\Delta m^2_{31}$  to ~1% precision with O(100 days) data

- 6 years for order of magnitude improvement over existing constraints
- Precise tests of neutrino oscillations and  $U_{PMNS}$  unitarity (1%)



#### **Construction Progress**



#### Steel Support Structure finished







PMT and electronics installation ongoing!



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#### **Construction Progress**



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• Filling set to start this year!



### **Neutrinos from Near and Far**











- Reactor antineutrinos have been key to neutrino science: and there are exciting prospects in the near future!
  - Discovery of non-zero  $\theta_{13}$  performed with reactor  $ar{
    u}_e$
  - Will be its most precise measurement for the foreseeable future
- JUNO is a next-generation, precision, multi-purpose experiment with a rich program in neutrino, astroparticle, and new physics searches
- Using its reactor  $\bar{\nu}_e$  dataset:
  - NMO measurement to  $3\sigma$  with about 6 years of data-taking
  - Sub-percent precision for  $\sin^2\theta_{12}$ ,  $\Delta m^2_{21}$  and  $\Delta m^2_{31}$  with as little as O(100) days of data
- Much progress has been done in construction and commissioning, with LS data-taking starting next year!



# Thank you!













- Comprehensive calibration strategy
  - Gamma/neutron sources, cosmogenic <sup>12</sup>B and UV laser
  - Multi-positional source deployment
- SPMTs serve as linear reference for LPMT nonlinearity
  - Operate in photon-counting mode for ~1-10 MeV
- Dual Calorimetry Calibration compares LPMT charge to SPMT charge under same source
  - Channel-wise LPMT charge vs. total SPMT charge
  - UV laser energies span region of interest
  - Gamma sources match time profile of neutrino (positron) signal
  - Absolute energy scale uncertainty <1%</li>





### **JUNO-TAO Physics Reach**



- With its short baseline, TAO has great potential in sterile oscillation searches
  - Sensitivity improved by virtual segmentation of detector







### **Systematics**



- Updated treatment of systematics
  - Values for 6 year exposure shown
- Rate systematics mitigated by spectral shape constraint on normalization

$\Delta m_{31}^2$	1σ (%)	
Statistics	0.17	
Reactor:		
- Uncorrelated	< 0.01	
- Correlated	0.01	
- Reference spectrum	0.05	
- Spent Nuclear Fuel	< 0.01	
- Non-equilibrium	< 0.01	
Detection:		
- Efficiency	0.01	
- Energy resolution	< 0.01	
- Nonlinearity	0.04	
- Backgrounds	0.04	
Matter density	0.01	
All systematics	0.08	
Total	0.19	
	0	.0 0.1
		%

$\sin^2 \theta_{12}$	lσ (%)			
Statistics	0.34		1	]
Reactor:			1	
- Uncorrelated	0.10			
- Correlated	0.27			
- Reference spectrum	0.09			
- Spent Nuclear Fuel	0.05			
- Non-equilibrium	0.10			
Detection:				
- Efficiency	0.23			
- Energy resolution	0.01			
- Nonlinearity	0.09			
- Backgrounds	0.20		(	
Matter density	0.07			
All systematics	0.40			
Total	0.52			
	0	.0 0	.2 %	0.4

2		7
$\Delta m_{21}^2$	1σ (%)	
Statistics	0.16	
Reactor:		
- Uncorrelated	0.01	
- Correlated	0.03	
- Reference spectrum	0.07	
- Spent Nuclear Fuel	0.07	
- Non-equilibrium	0.14	
Detection:		
- Efficiency	0.02	
- Energy resolution	0.01	
- Nonlinearity	0.05	
- Backgrounds	0.18	
Matter density	0.01	
All systematics	0.27	
Total	0.32	
	0	0 02

$\sin^2 \theta_{13}$	1σ (%)	
Statistics	8.94	
Reactor:		
- Uncorrelated	2.53	
- Correlated	6.83	
- Reference spectrum	3.48	
- Spent Nuclear Fuel	1.55	
- Non-equilibrium	2.65	
Detection:		
- Efficiency	5.81	
- Energy resolution	0.39	
- Nonlinearity	2.09	
- Backgrounds	4.89	
Matter density	0.98	
All systematics	8.16	
Total	12.11	
		0 5 10

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