

Reactor Neutrinos Overview

Thiago Bezerra (University of Sussex, Brighton, UK)

4th Conference on Science at the Sanford Underground Research Facility

11th May, 2022

US

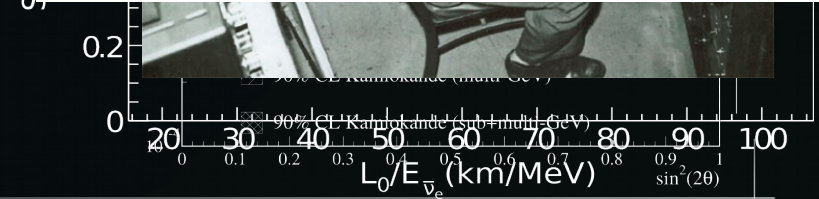
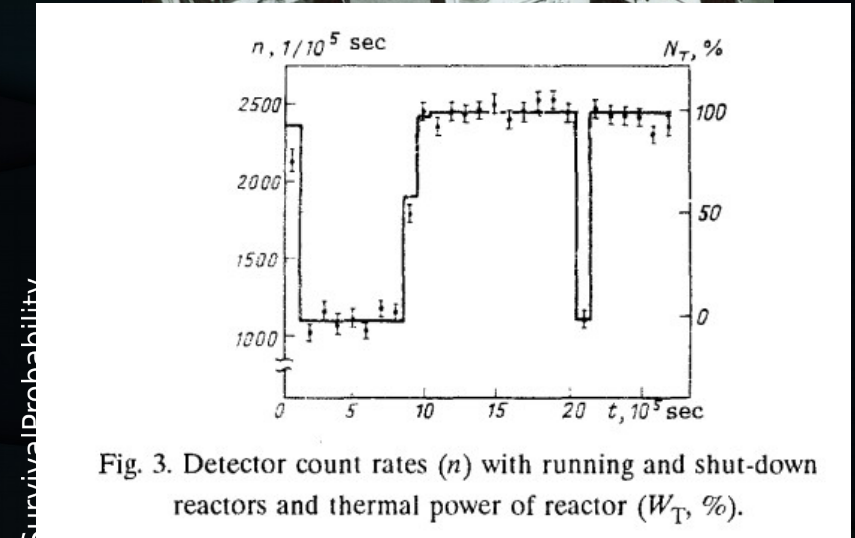
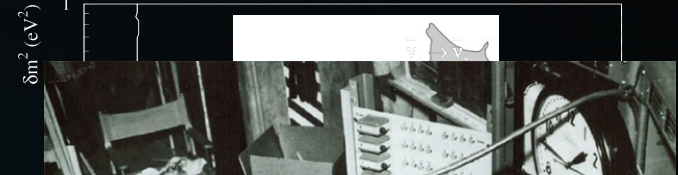
UNIVERSITY
OF SUSSEX

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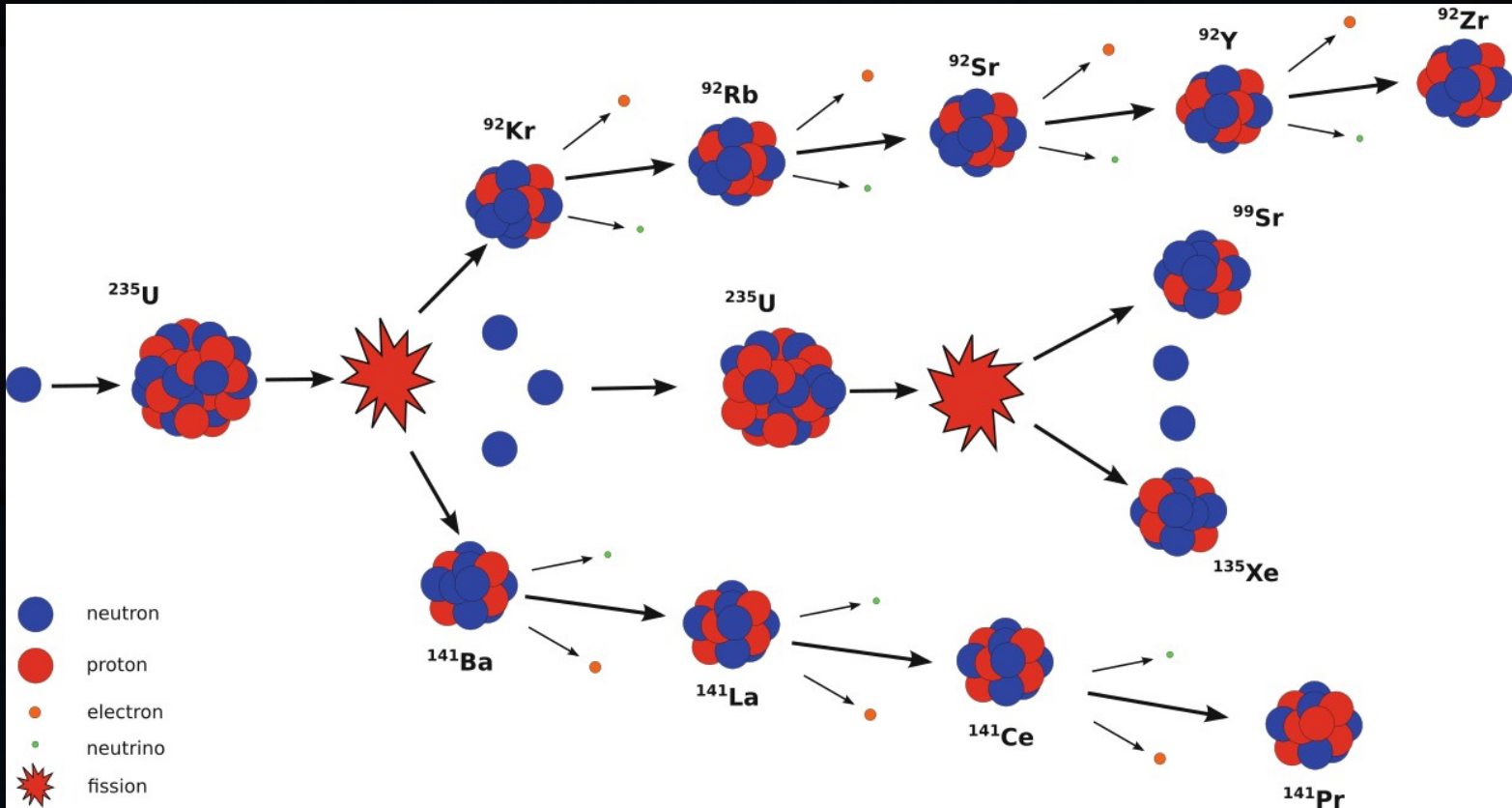
- Quick overview of the past
- Reactor neutrinos and 3 flavour oscillations status (IBD)
 - Other rich physics available: Sterile, CE ν NS, E.S.
- Future: Positron tagging

Quick overview from the past

- 50's: ν discovery
- 80's: Reactor monitoring
- 90's: ν oscillations searches
 - Bugey4: Flux measurement (1.4%)
 - CHOOZ: Upper limit on θ_{13}
- 2000's: Settling ν oscillations
 - KamLAND (Δm_{21}^2)



Nucl. reactors: “free” and copious ν source

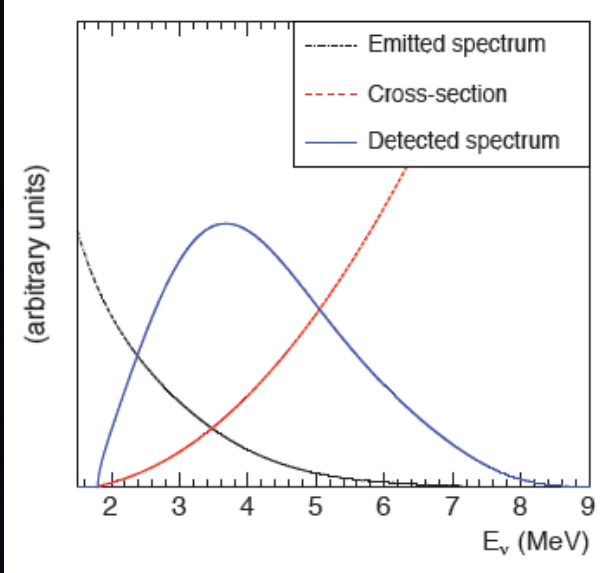


Commercial fission reactor produces $\sim 10^{20}$ ν /sec

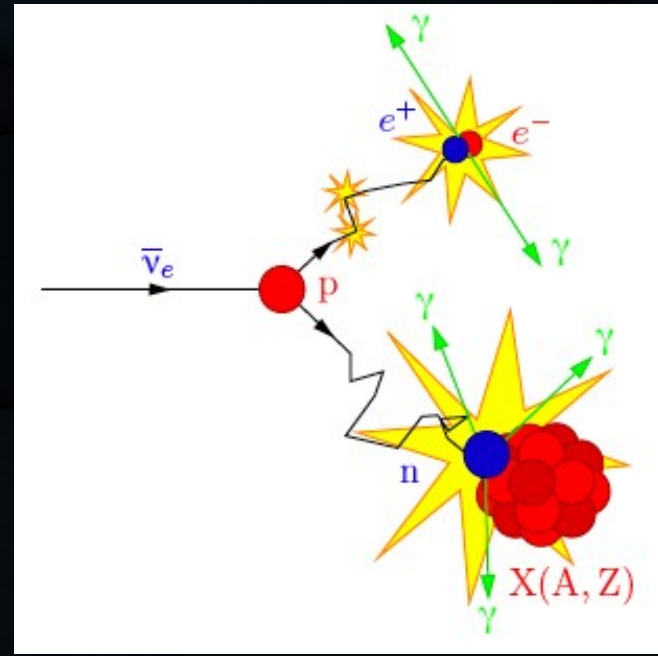
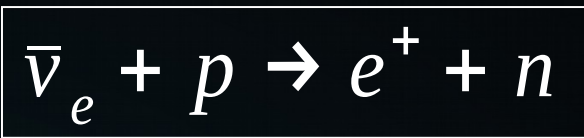
Reactor- ν detection via IBD

$$N_{\nu}^{\text{exp}}(t) = \frac{\epsilon N_p}{4\pi L^2} \times \frac{P_{th}(t)}{\langle E_f \rangle} \times \langle \sigma_f \rangle$$

Mean cross section per fission

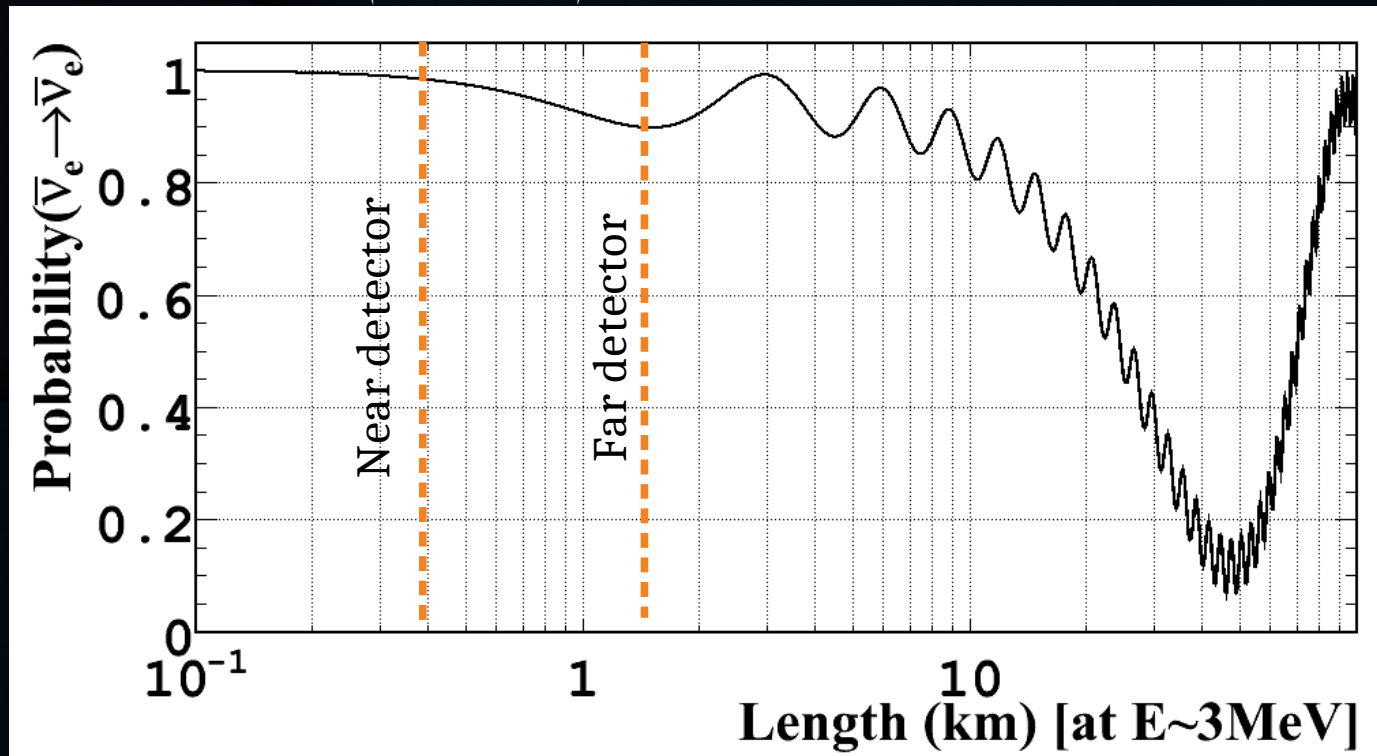


Inverse Beta Decay

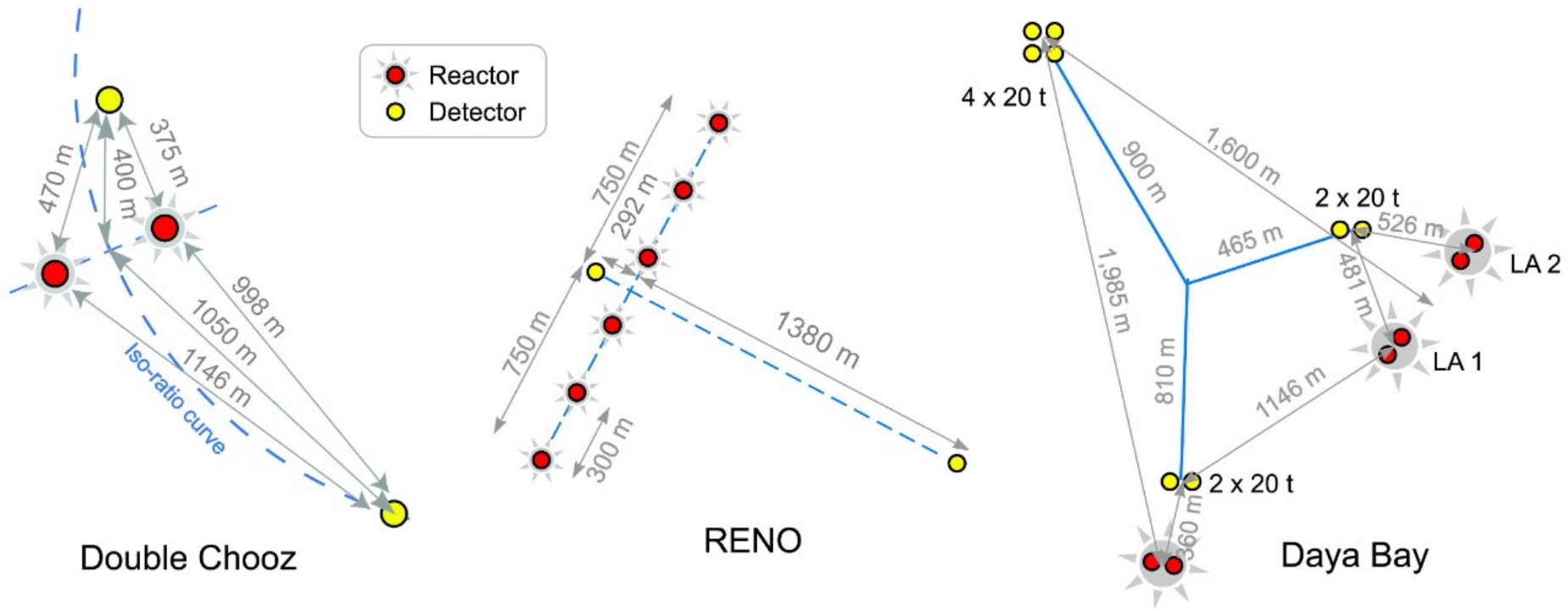


θ_{13} measurement with reactor $\bar{\nu}$'s

$$P(\bar{\nu}_e \rightarrow \bar{\nu}_e) = 1 - \sin^2 2\theta_{13} \sin^2 \left(\frac{\Delta m_{ee}^2 L}{4E} \right)$$



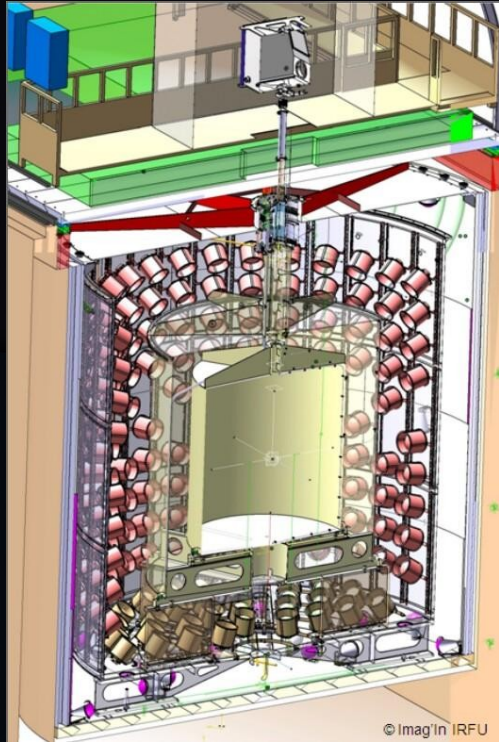
Reactor- θ_{13} experiments (2010 – present)



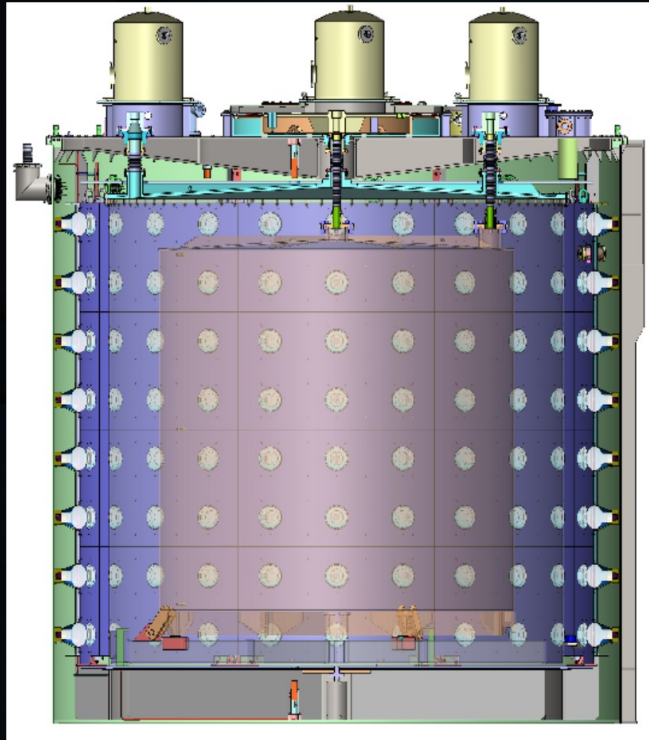
J.Phys.G37:103001,2010 (fig)

Reactor- θ_{13} detectors

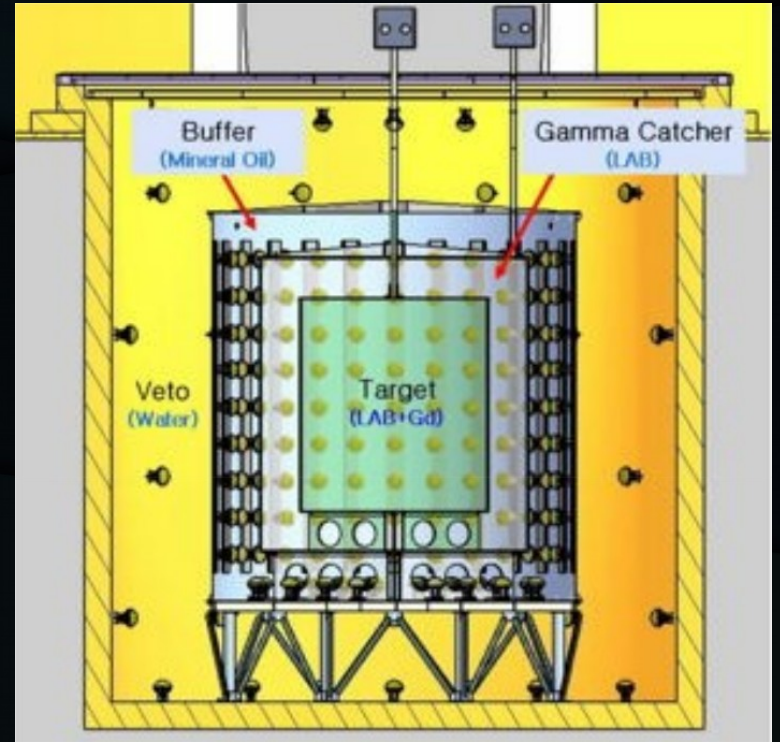
DC



DYB

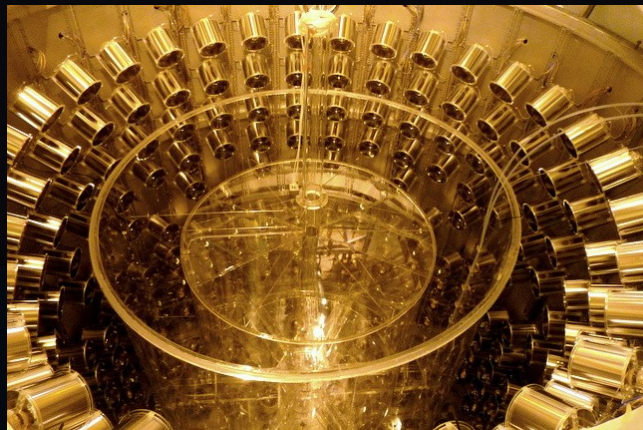


RENO

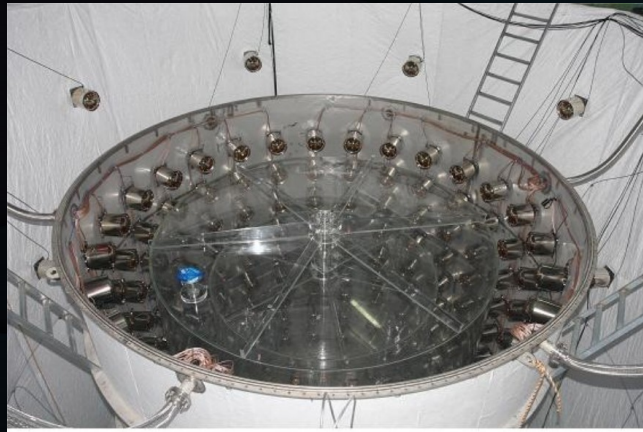


Reactor- θ_{13} detectors

DC



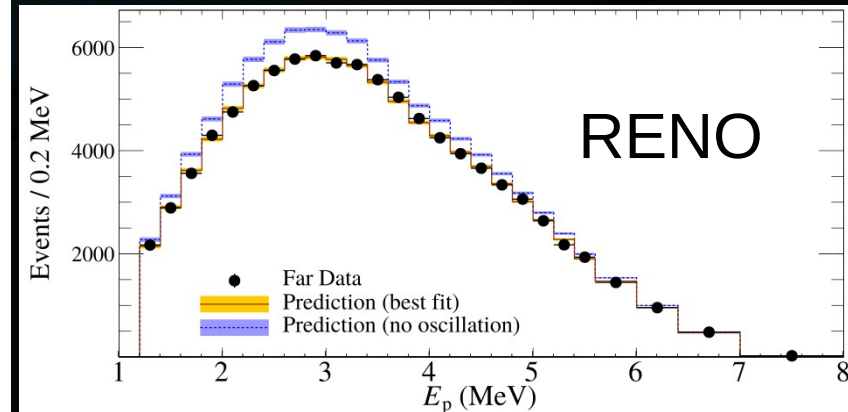
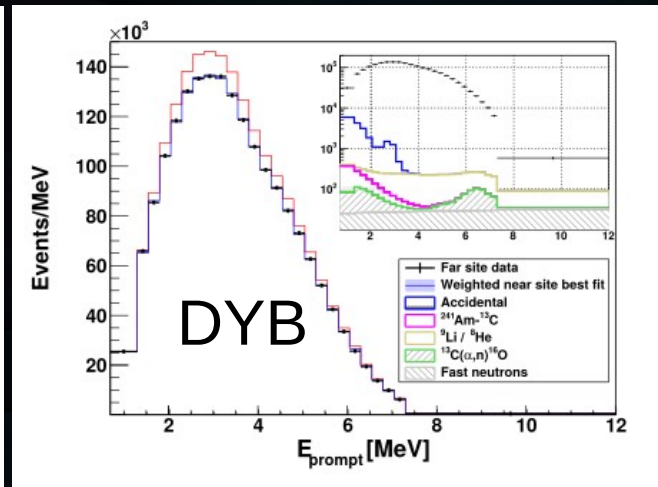
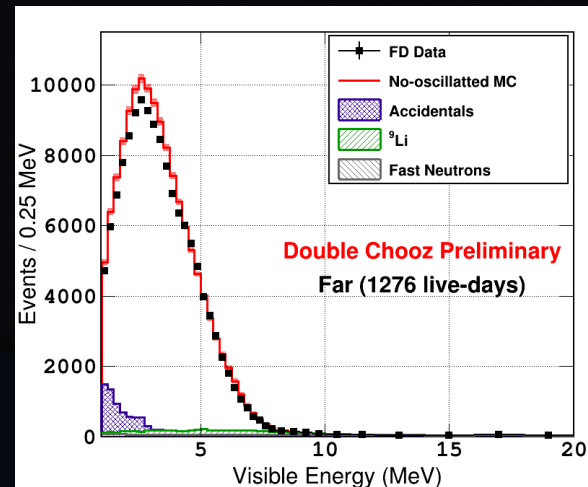
RENO



DYB

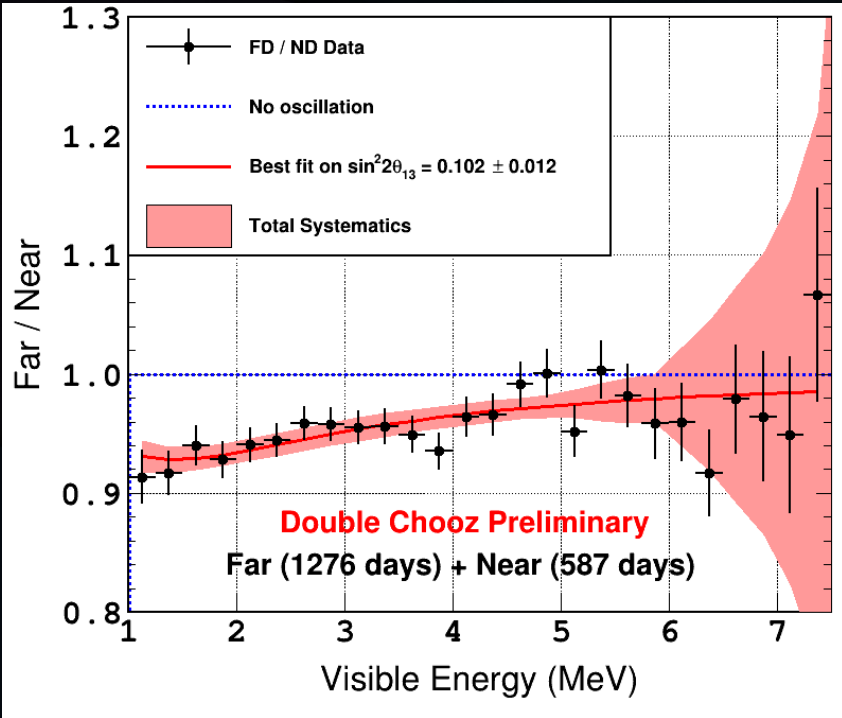


Reactor- θ_{13} spectra

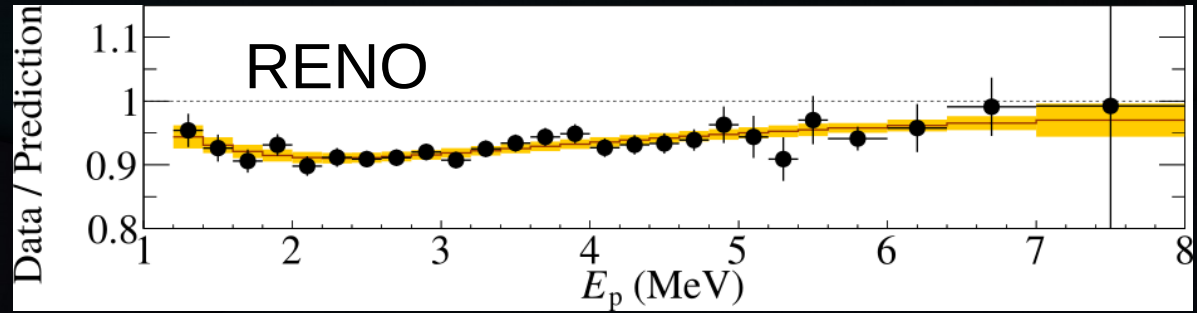


Reactor- θ_{13} far/near ratios

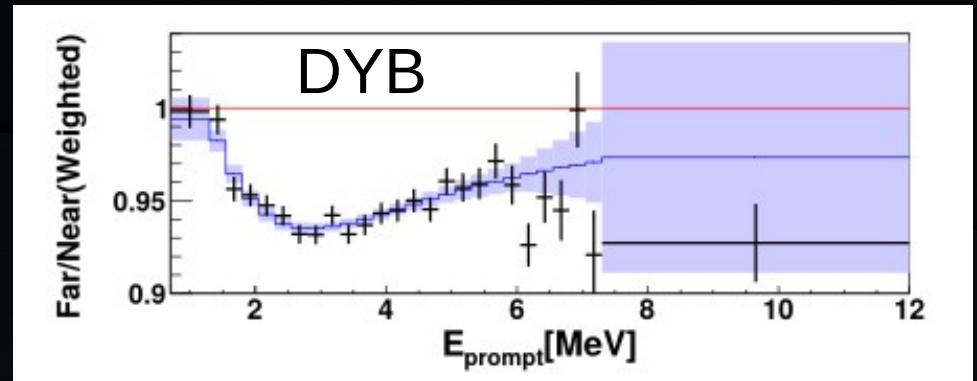
~1.0 km



~1.4 km

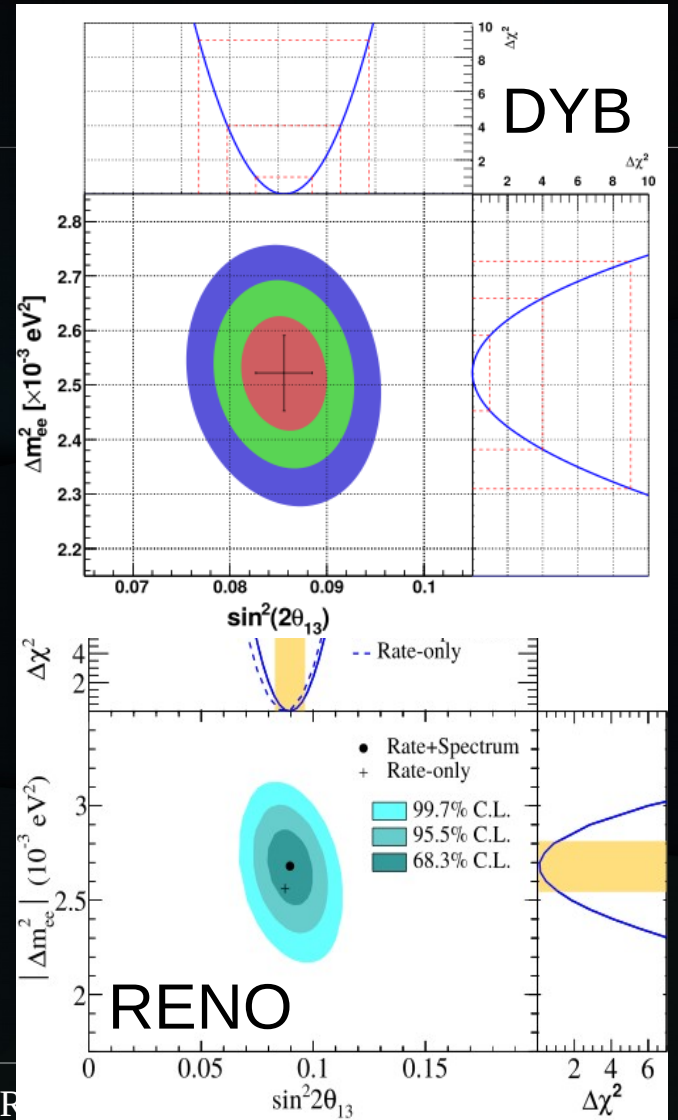
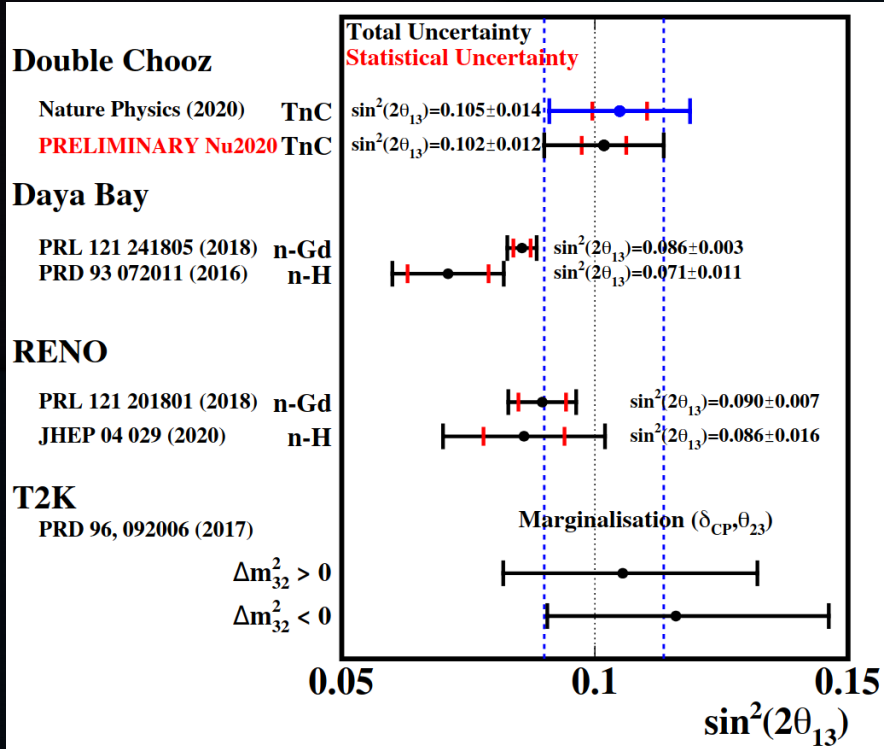


~1.7 km

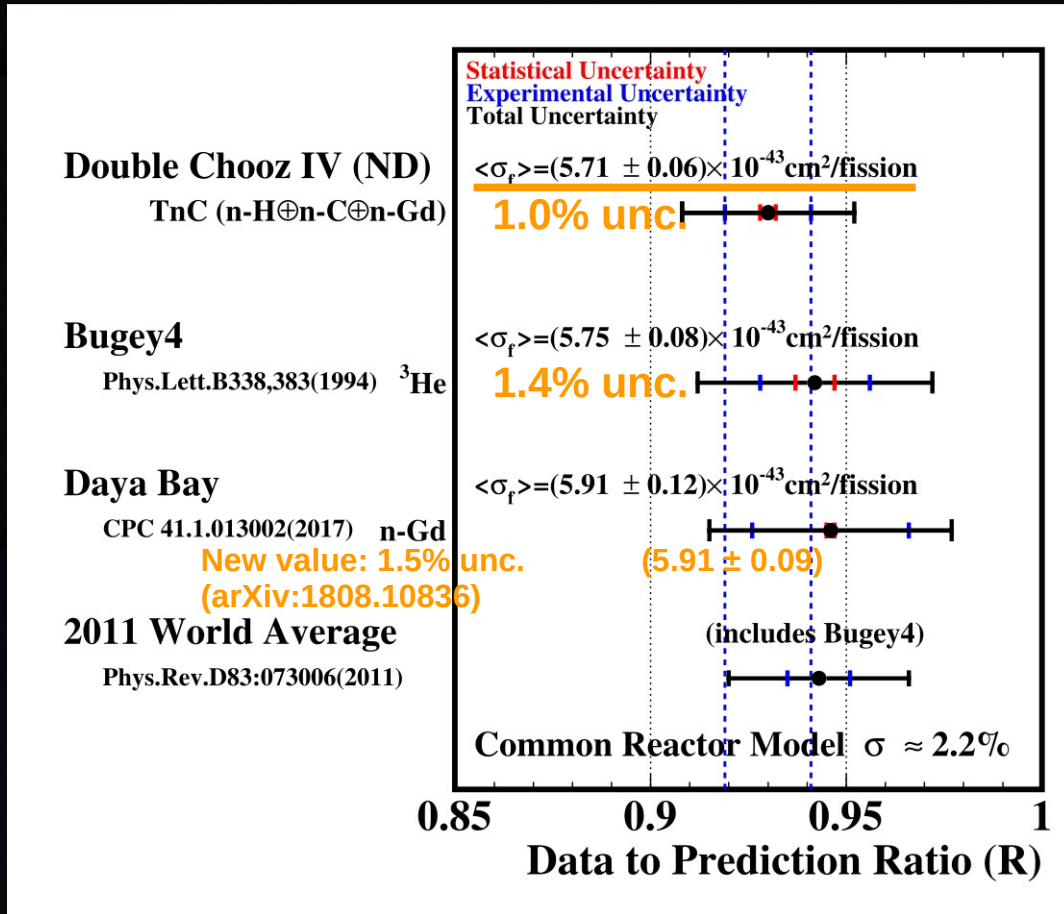


Reactor- θ_{13} summary

$$P(\bar{\nu}_e \rightarrow \bar{\nu}_e) = 1 - \sin^2 2\theta_{13} \sin^2 \left(\frac{\Delta m_{ee}^2 L}{4E} \right)$$

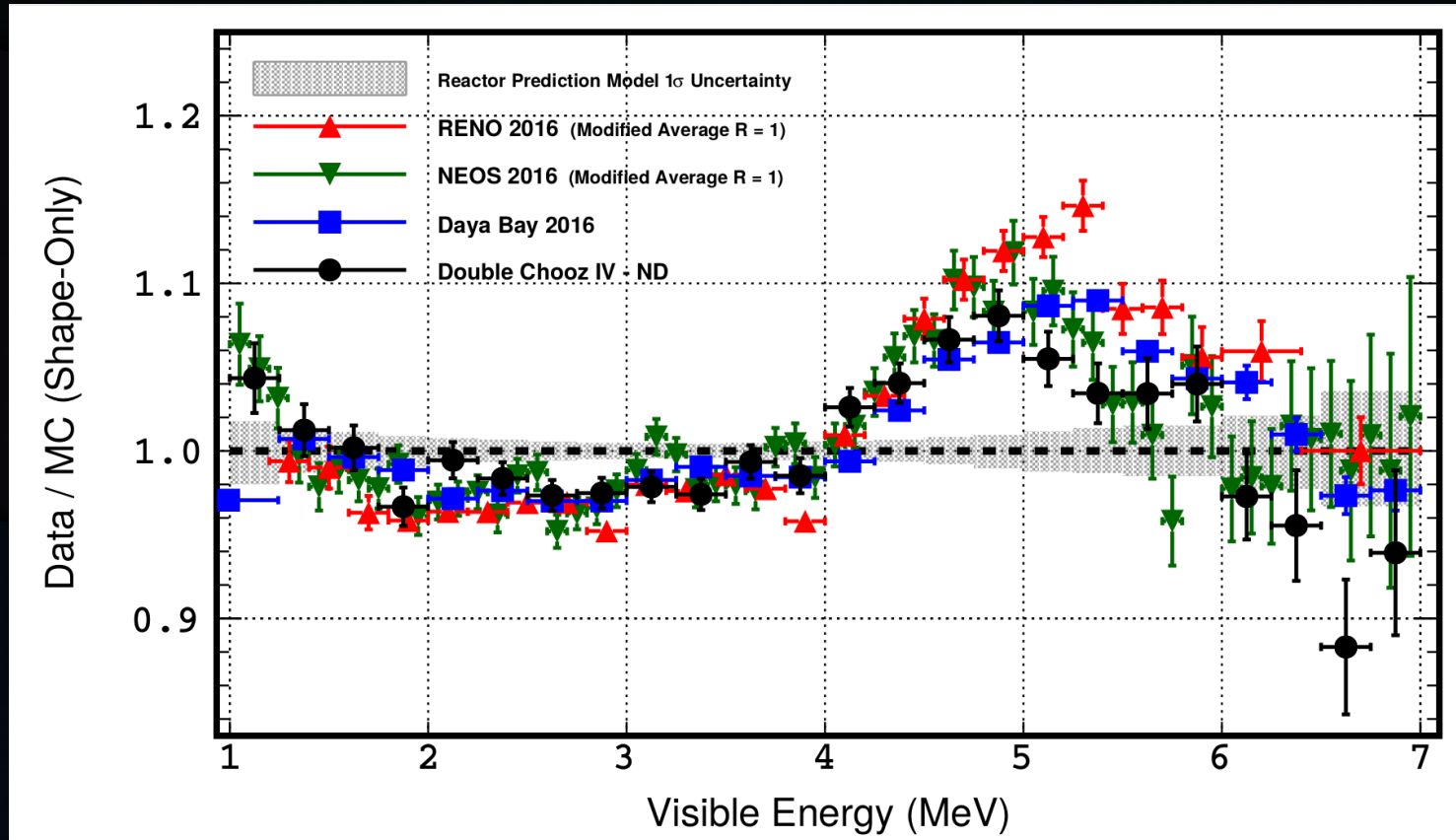


Reactor- θ_{13} : flux normalization



Many experiments exploring this effect (VSBL)

Reactor- θ_{13} : spectral shape distortion



Nature Physics, volume 16, pages 558–564 (2020)

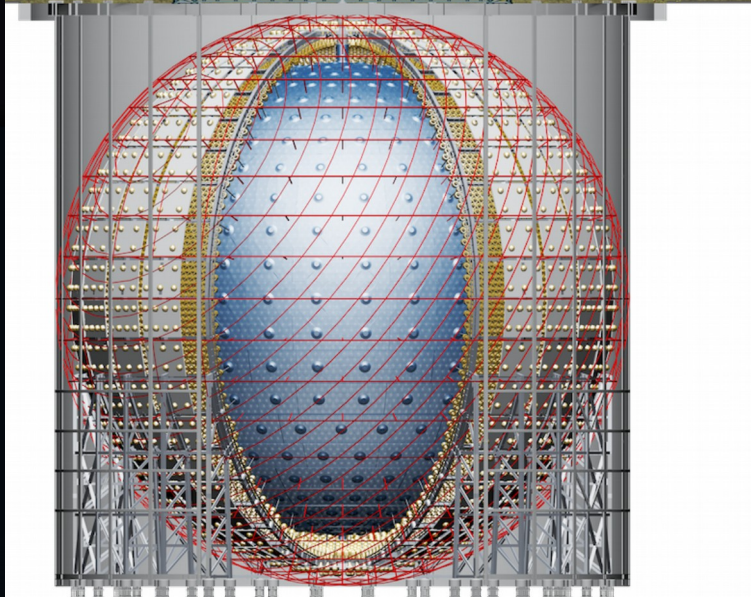
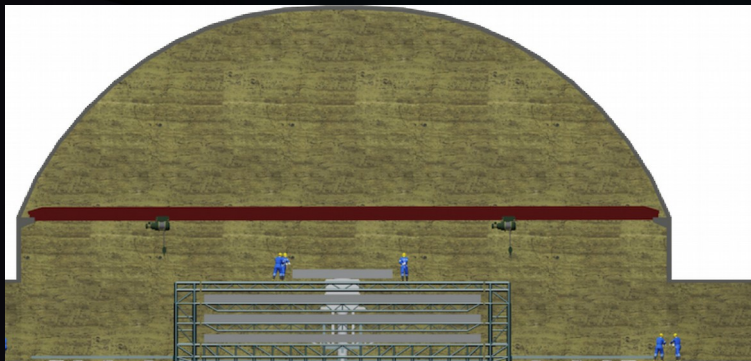
C.Giunti *et al.*, Flux prediction review: arXiv:2110.06820

Reactor IBD Future

JUNO

JUNO

The Jiangmen Underground Neutrino Observatory



-> Huge liquid scintillator

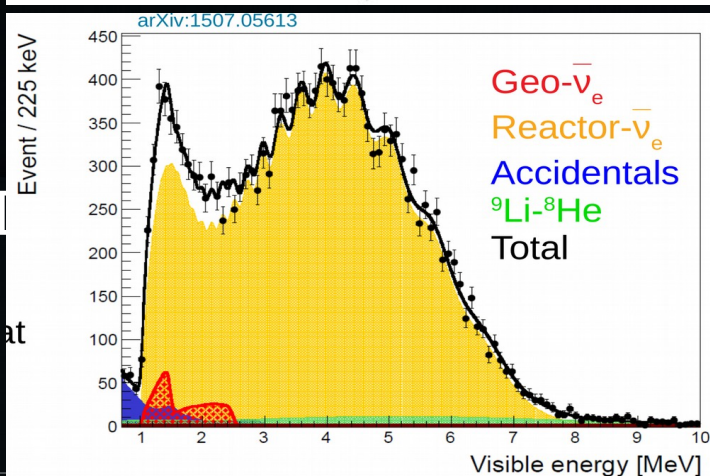
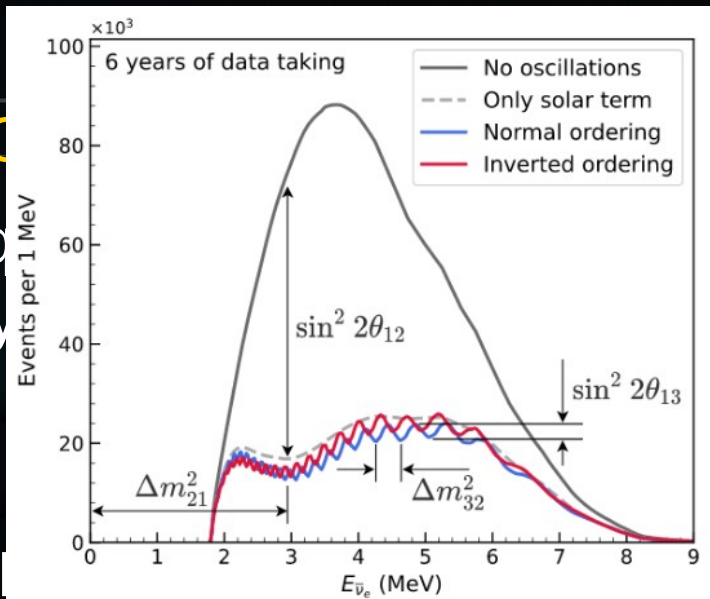
-> Two key parameters

DC/DYB/

Borexino

KamLAND

JUNO at



precise

solution

@ 1 MeV)

8%

5%

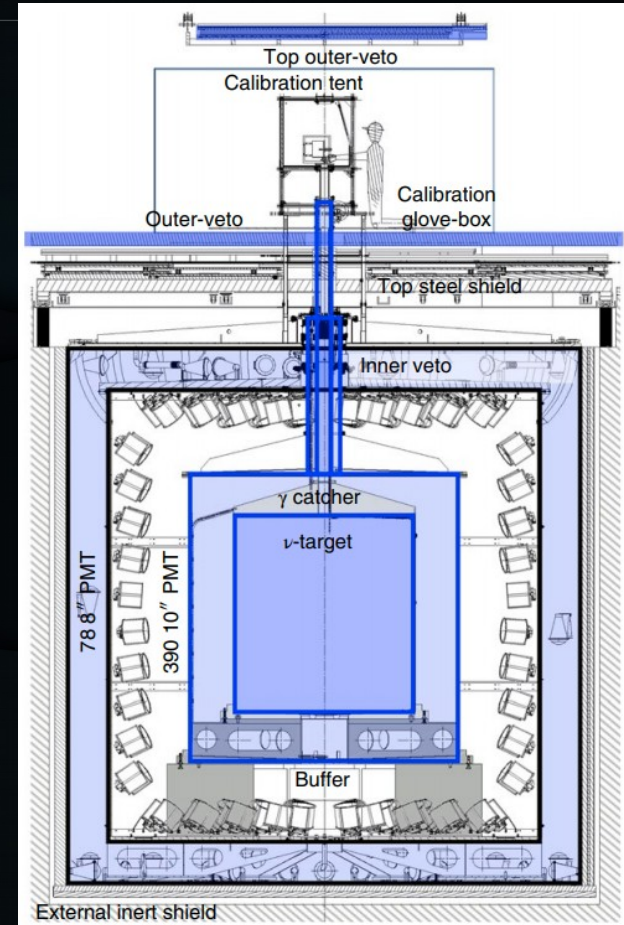
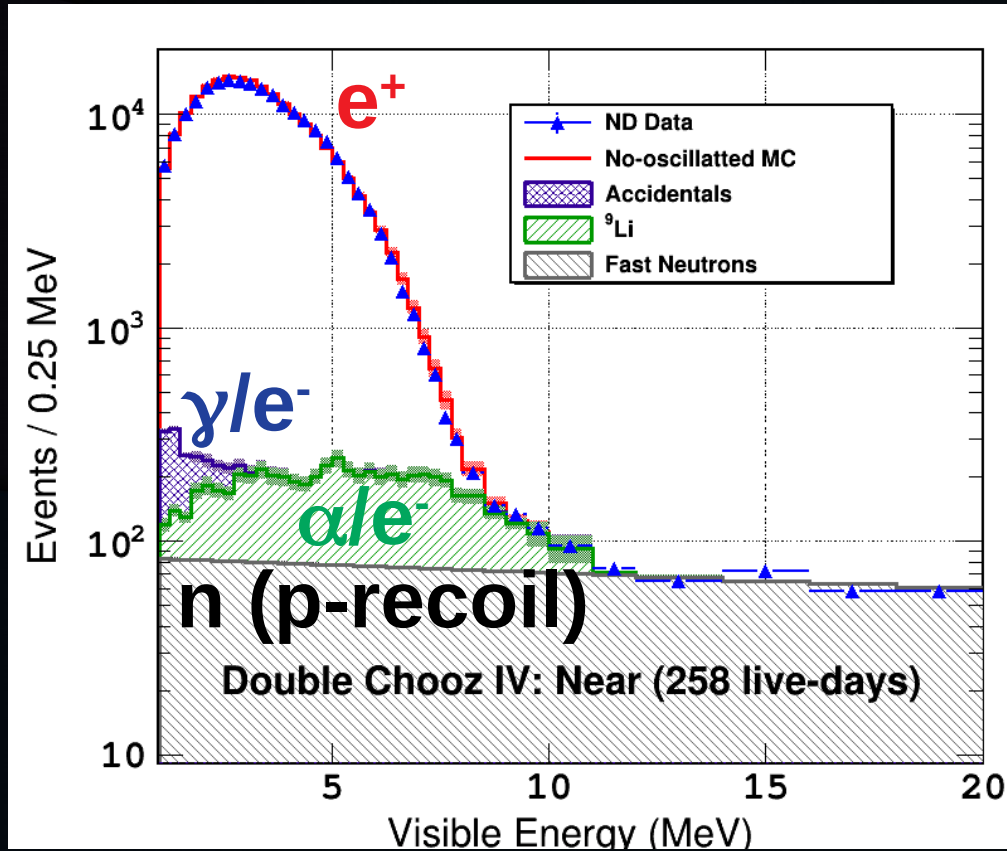
6%

3%

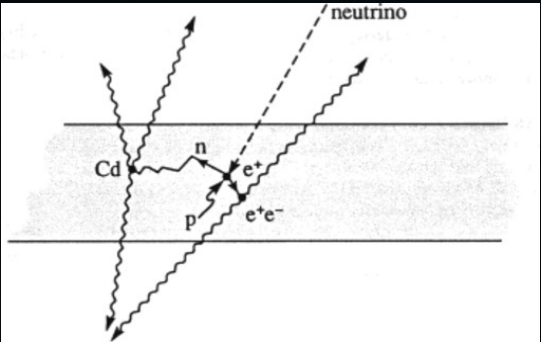
Reactor IBD Future (beyond JUNO)

Positron Tagging

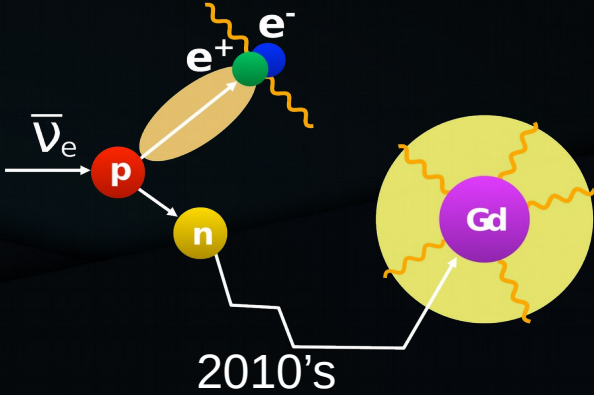
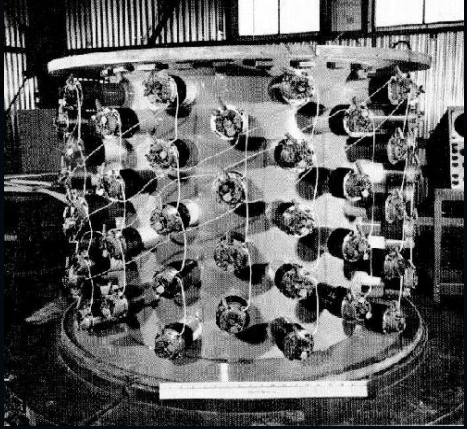
Expensive way to remove backgrounds



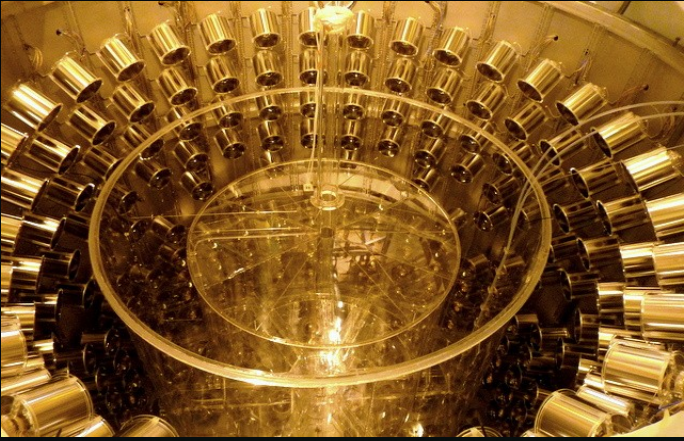
60 years old strategy!



1950's

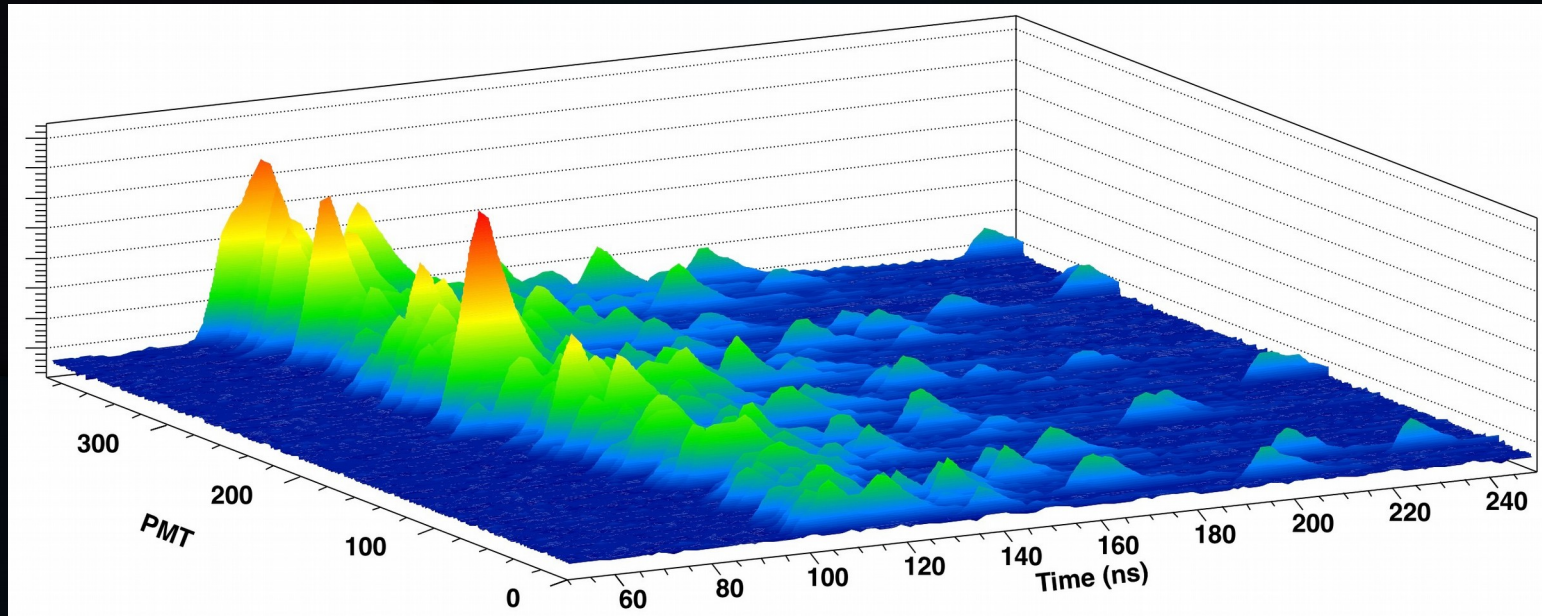


2010's

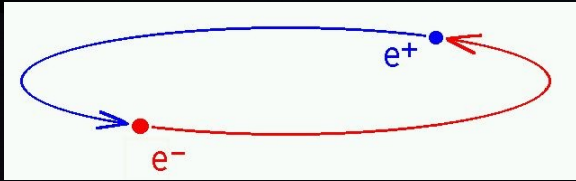


Efforts on e^+ identification

Double Chooz case: FADC, a powerful tool

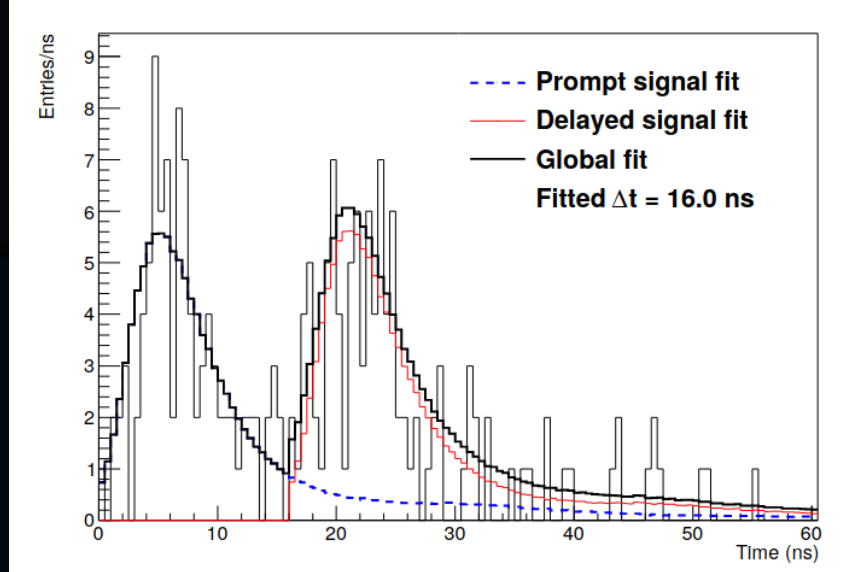


Efforts on e^+ identification



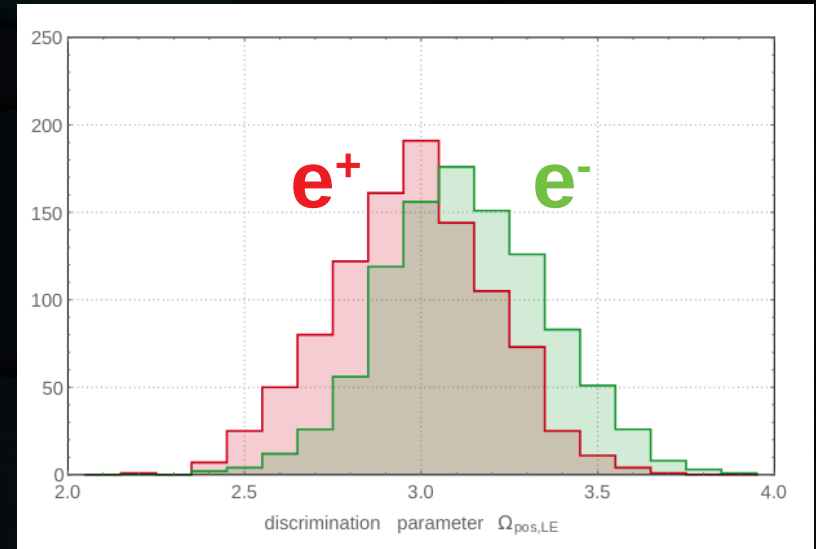
$$\varepsilon < 50\%$$

Ortho-positronium observation



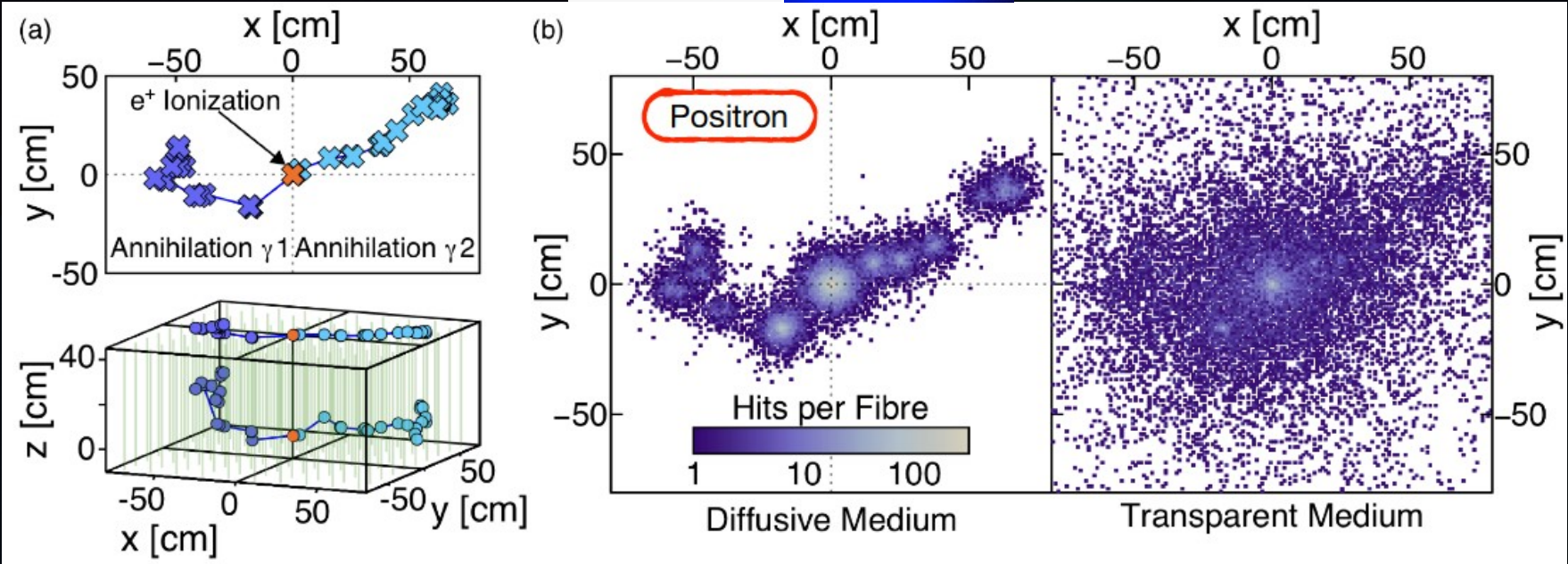
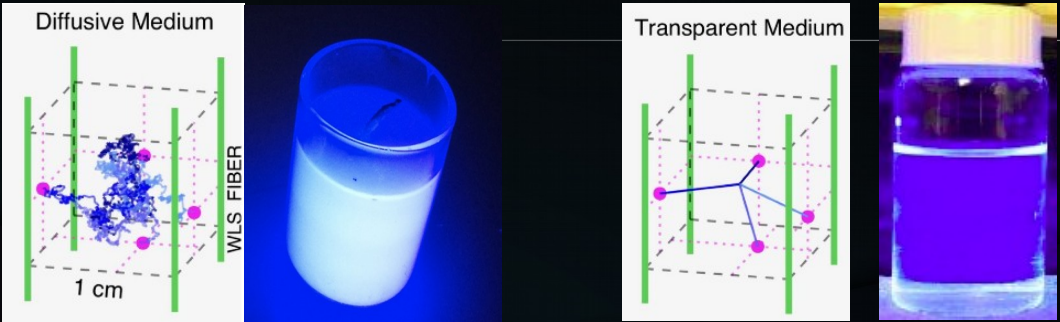
JHEP 1410 (2014) 032

Pulse shape discrimination
Fourier power spectra of scintillator pulses

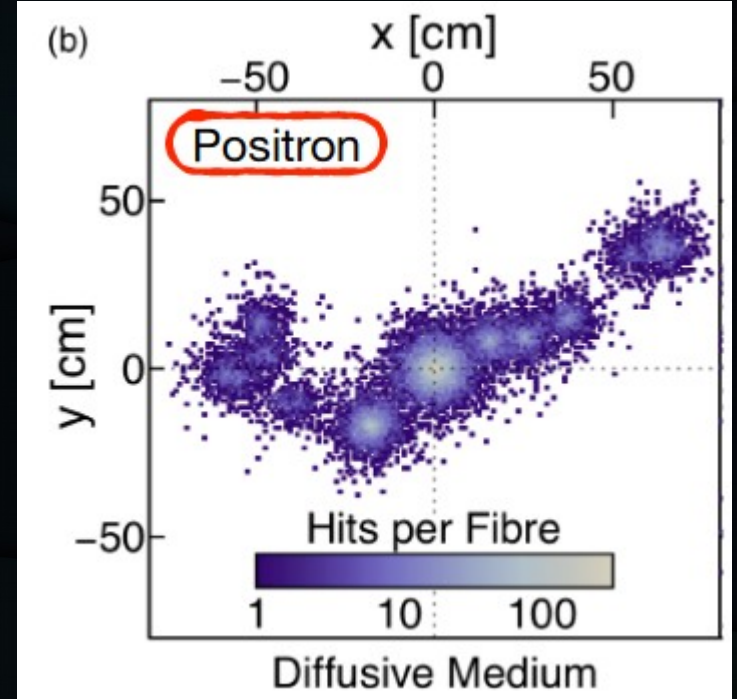
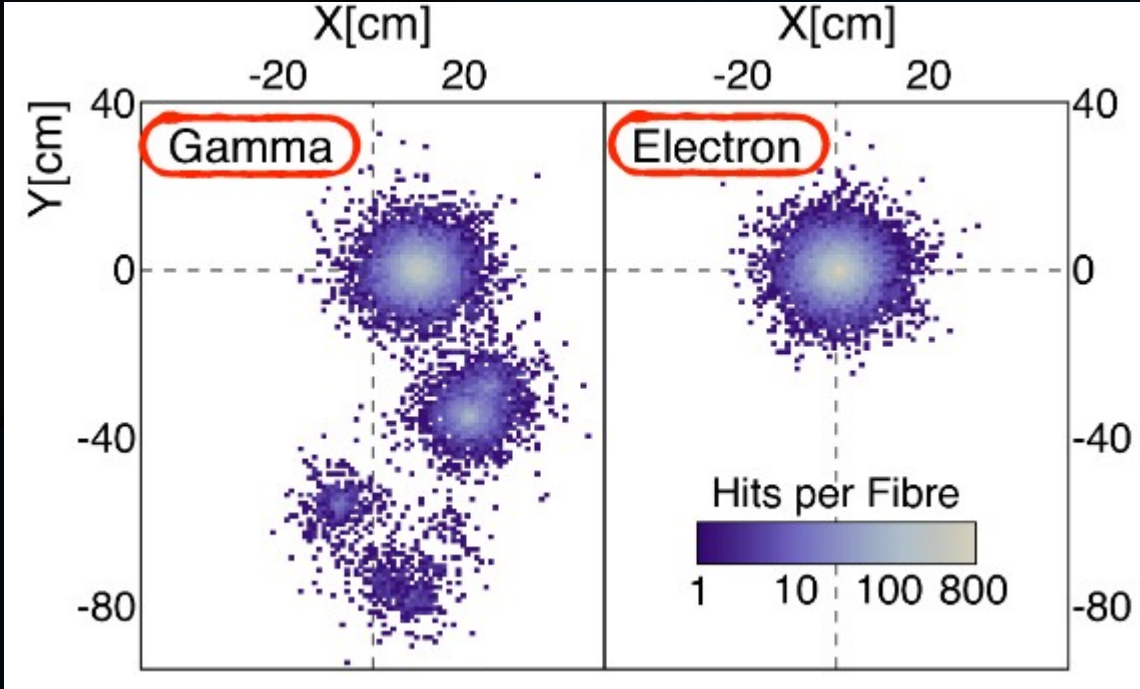


2018 JINST 13 P01031

A new approach!



Powerful particle identification!



Cannot separate these 3 on an event-by-event basis with conventional LS detectors!

First papers

arXiv:1908.02859

communications physics

ARTICLE

<https://doi.org/10.1038/s42005-021-00763-5>

OPEN

Neutrino physics with an opaque detector

LiquidO Consortium*

In 1956 Reines & Cowan discovered the neutrino using a liquid scintillator detector. The neutrinos interacted with the scintillator, producing light that propagated across transparent volumes to surrounding photo-sensors. This approach has remained one of the most widespread and successful neutrino detection technologies used since. This article introduces a concept that breaks with the conventional paradigm of transparency by confining and collecting light near its creation point with an opaque scintillator and a dense array of optical fibres. This technique, called LiquidO, can provide high-resolution imaging to enable efficient identification of individual particles event-by-event. A natural affinity for adding dopants at high concentrations is provided by the use of an opaque medium. With these and other capabilities, the potential of our detector concept to unlock opportunities in neutrino physics is presented here, alongside the results of the first experimental validation.

arXiv:1908.03334

Jinst

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Published: November 3, 2019

Novel opaque scintillator for neutrino detection

C. Buck,¹ B. Gramlich and S. Schoppmann

*Max-Planck-Institut für Kernphysik,
Saupfercheckweg 1, 69117 Heidelberg, Germany
E-mail: christian.buck@mpi-hd.mpg.de*

ABSTRACT: There is rising interest in organic scintillators with low scattering length for future neutrino detectors. Therefore, a new scintillator system was developed based on admixtures of paraffin wax in linear alkyl benzenes. The transparency and viscosity of this gel-like material can be tuned by temperature adjustment. Whereas it is a colorless transparent liquid at temperatures around 40°C, it has a milky wax structure below 20°C. The production and properties of such a scintillator as well as its advantages compared to transparent liquids are described.

KEYWORDS: Detector design and construction technologies and materials; Neutrino detectors; Scintillators, scintillation and light emission processes (solid, gas and liquid scintillators)

ArXiv ePrint: 1908.03334

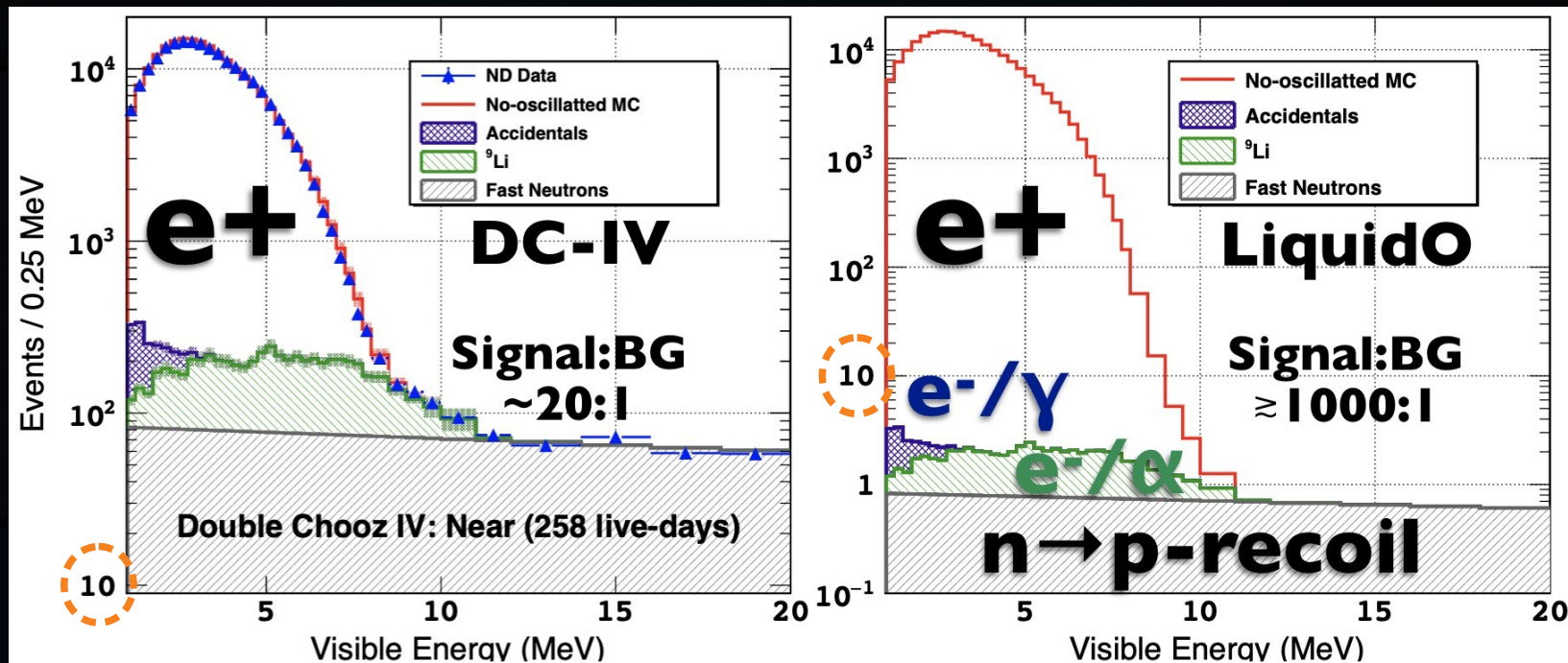
*Corresponding author:

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<https://doi.org/10.1088/1748-0221/14/11/P11007>

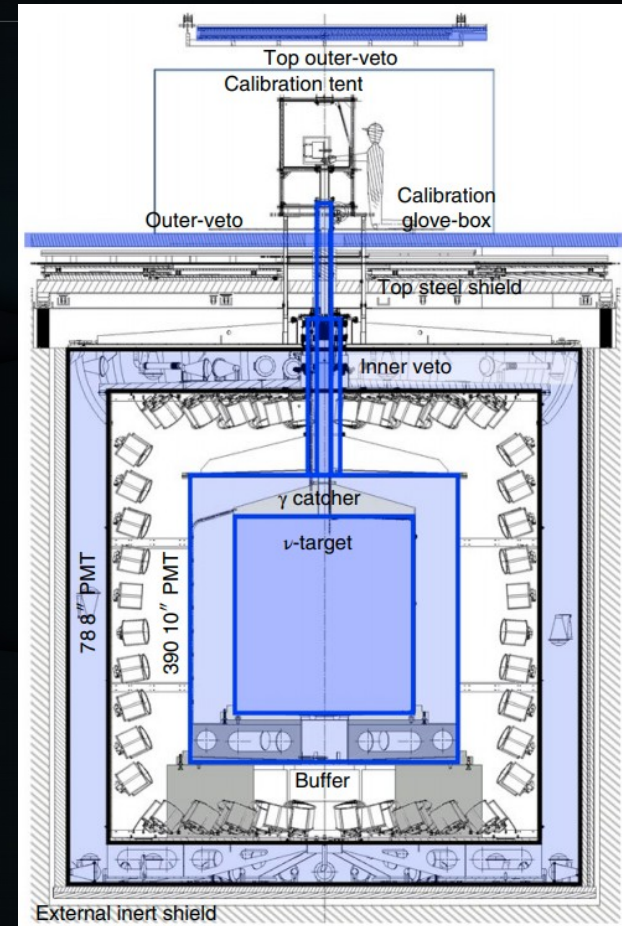
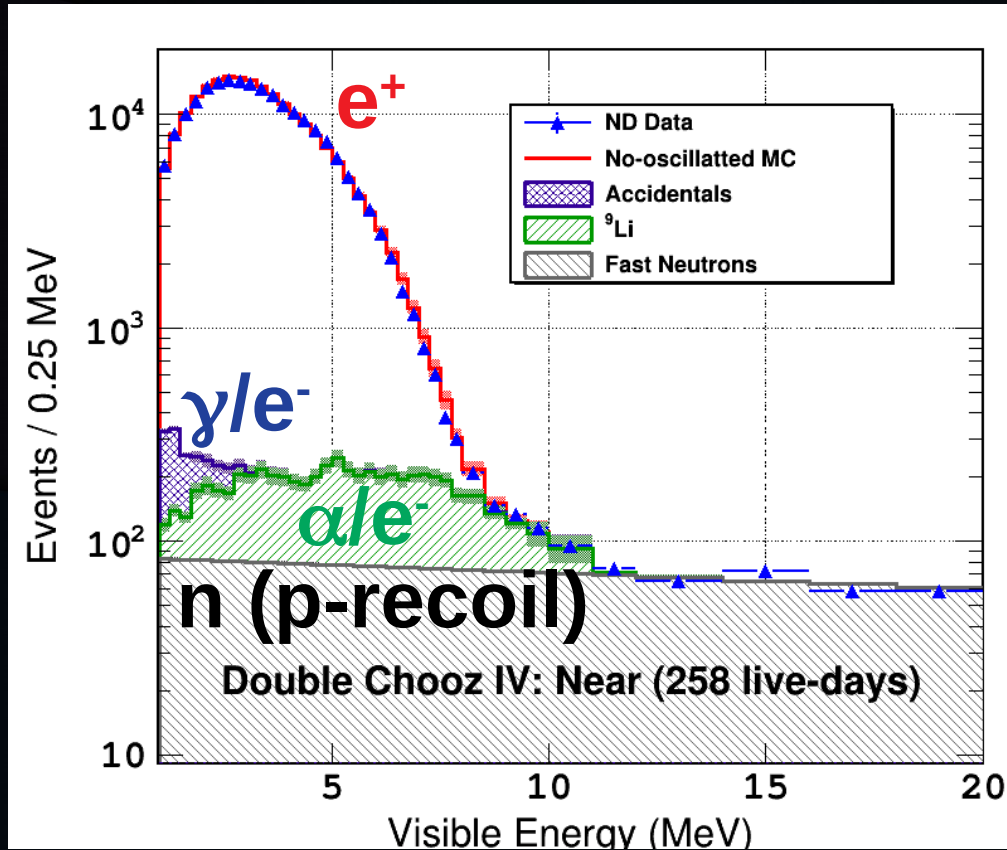
2019 JINST 14 P11007

Impact on IBD detection

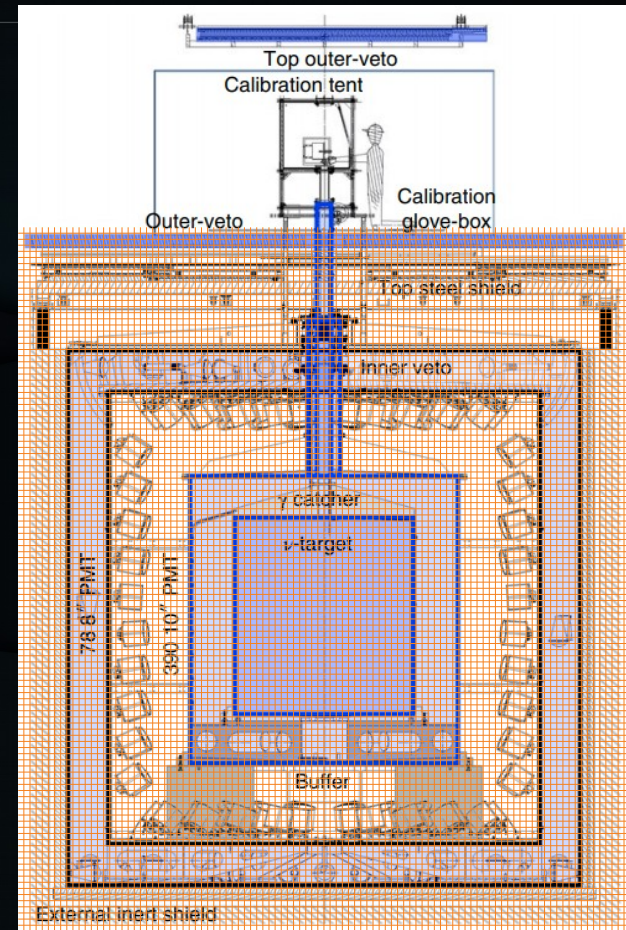
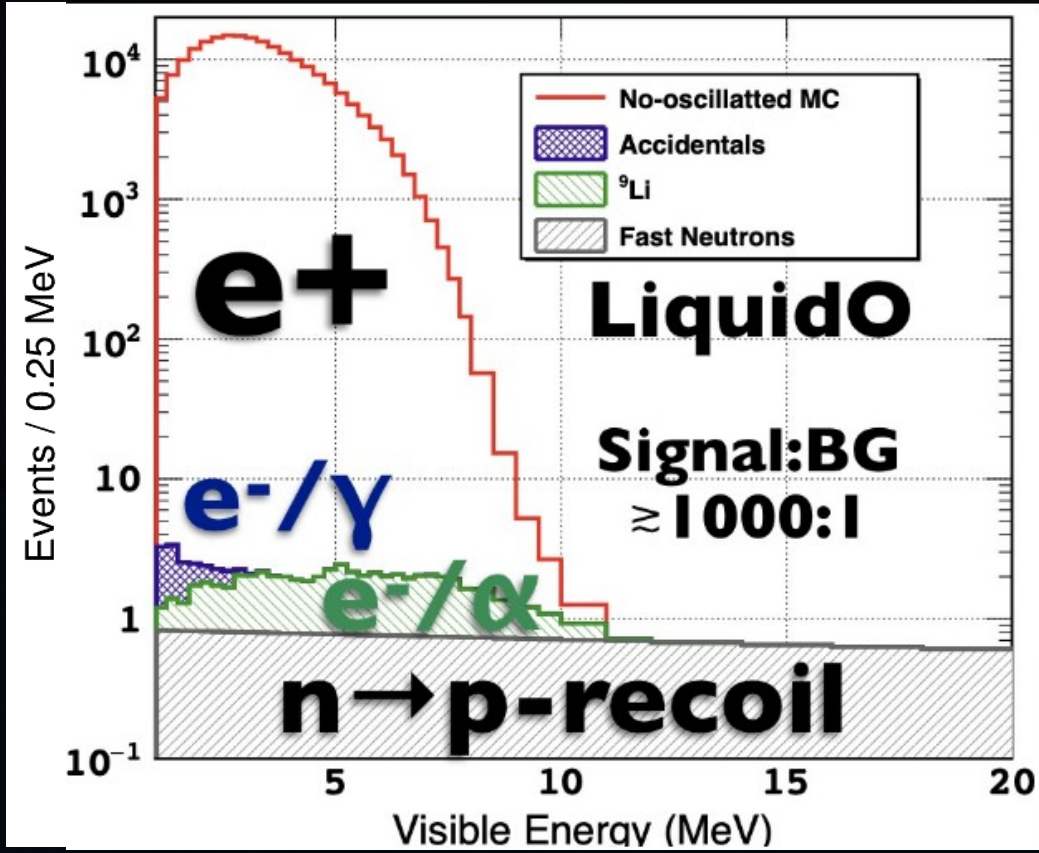


Applying expected PID and BG rejection

Expensive way to remove backgrounds



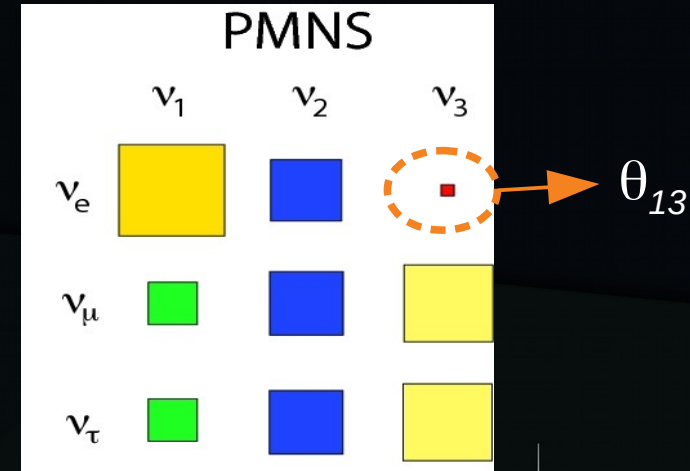
Smart way to remove backgrounds



10 fold increase!

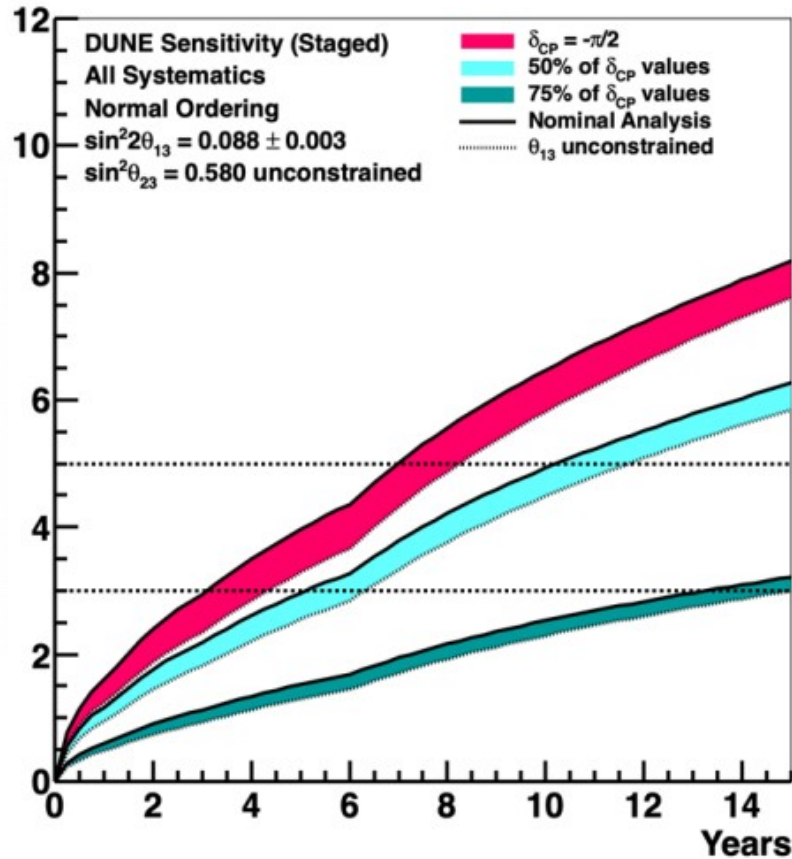
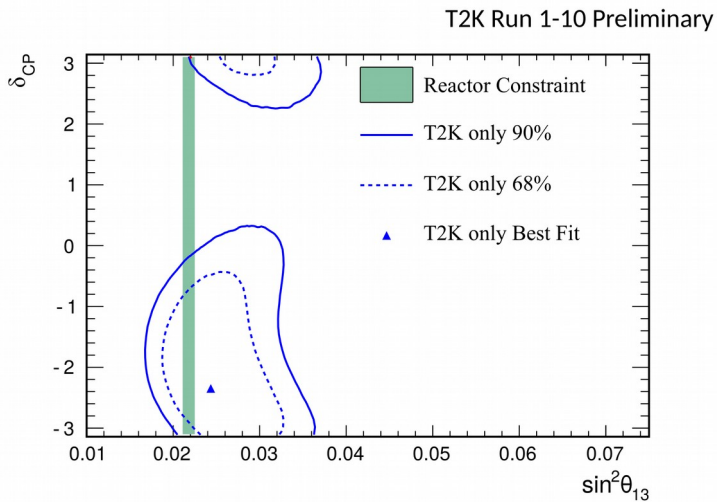
Impact on θ_{13}

- Not “doomed” to 2~3% unc. any more
- Possible to get sub-percent unc. with current technology (preliminary, energy scale main syst.)



$$J_r = \sin \theta_{12} \cos \theta_{12} \sin \theta_{23} \cos \theta_{23} \sin \theta_{13} \cos^2 \theta_{13} \sin \delta_{CP}$$

Impact of sub-percent θ_{13}



Impact of
 constraining $\sin^2 2\theta_{13}$
 within 3.4%

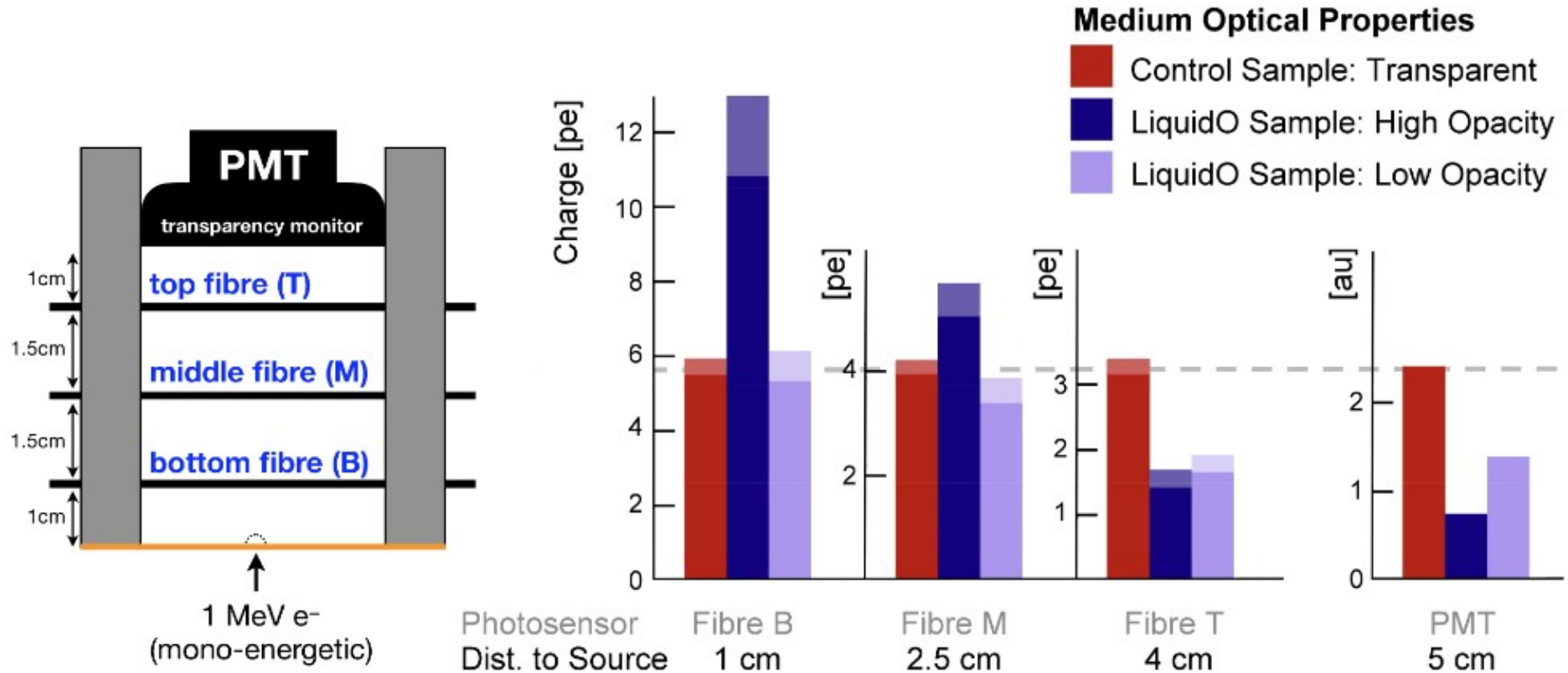
How many years we
 gain if < 1%?

Summary

- Nuclear reactors are a key source for neutrino physics (no doubt!);
- IBD channel driving θ_{13}
 - Best known oscillation parameter (others soon too, with JUNO)
- Positron tagging with opaque scintillator has a promising future
 - Counter intuitive technique
 - Proof-of-principle already obtained. Working towards multi-ton demonstrator
 - Possible way to do a sub-% measurement of θ_{13}

Back up slides

Does it work?



Commission a 7.5 litre prototype called "Mini-LiquidO"

