

# Search for $0\nu\beta\beta$ beyond $10^{28}$ yr half-life sensitivity with nEXO

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Ako Jamil for the nEXO Collaboration

Yale University

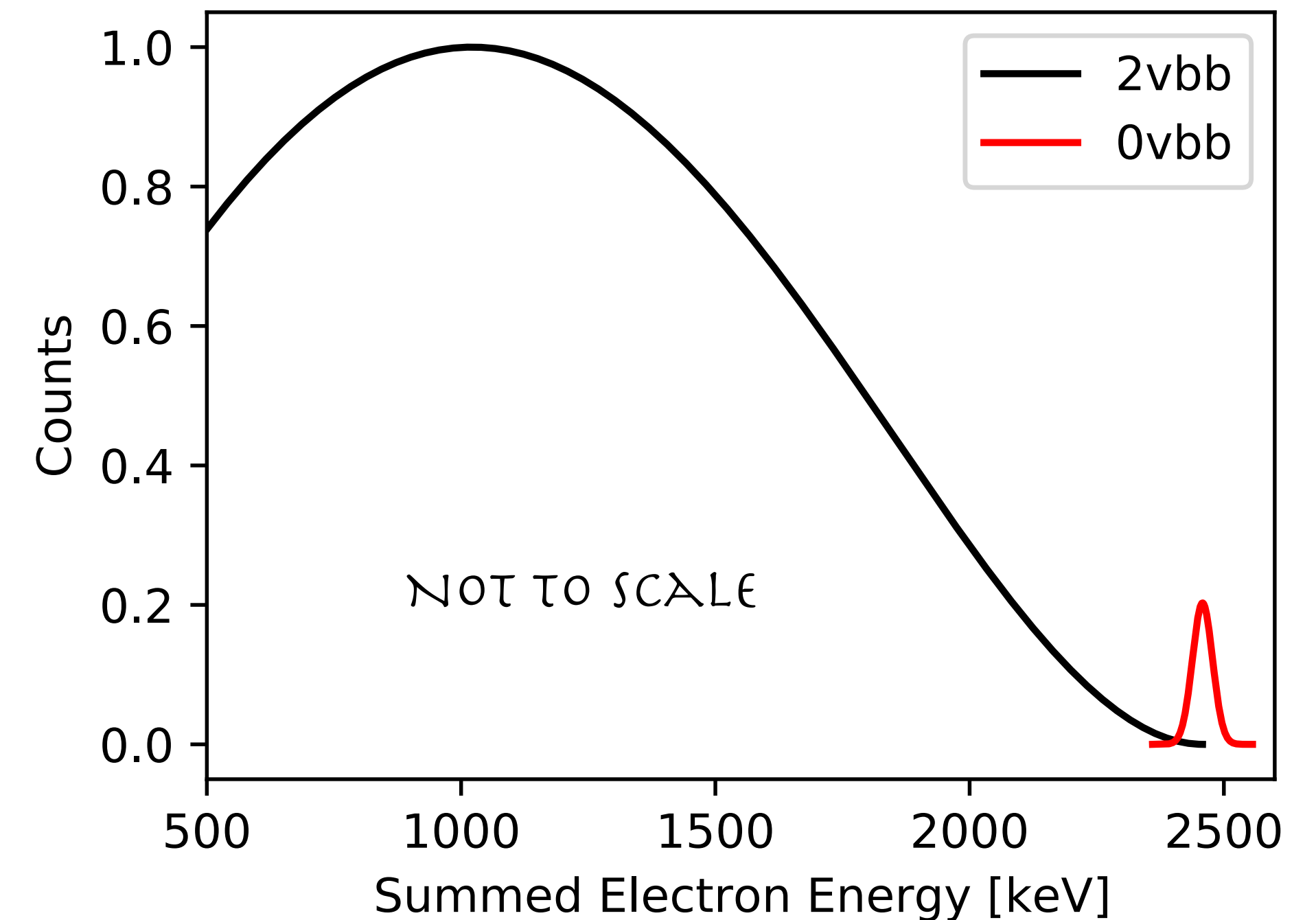
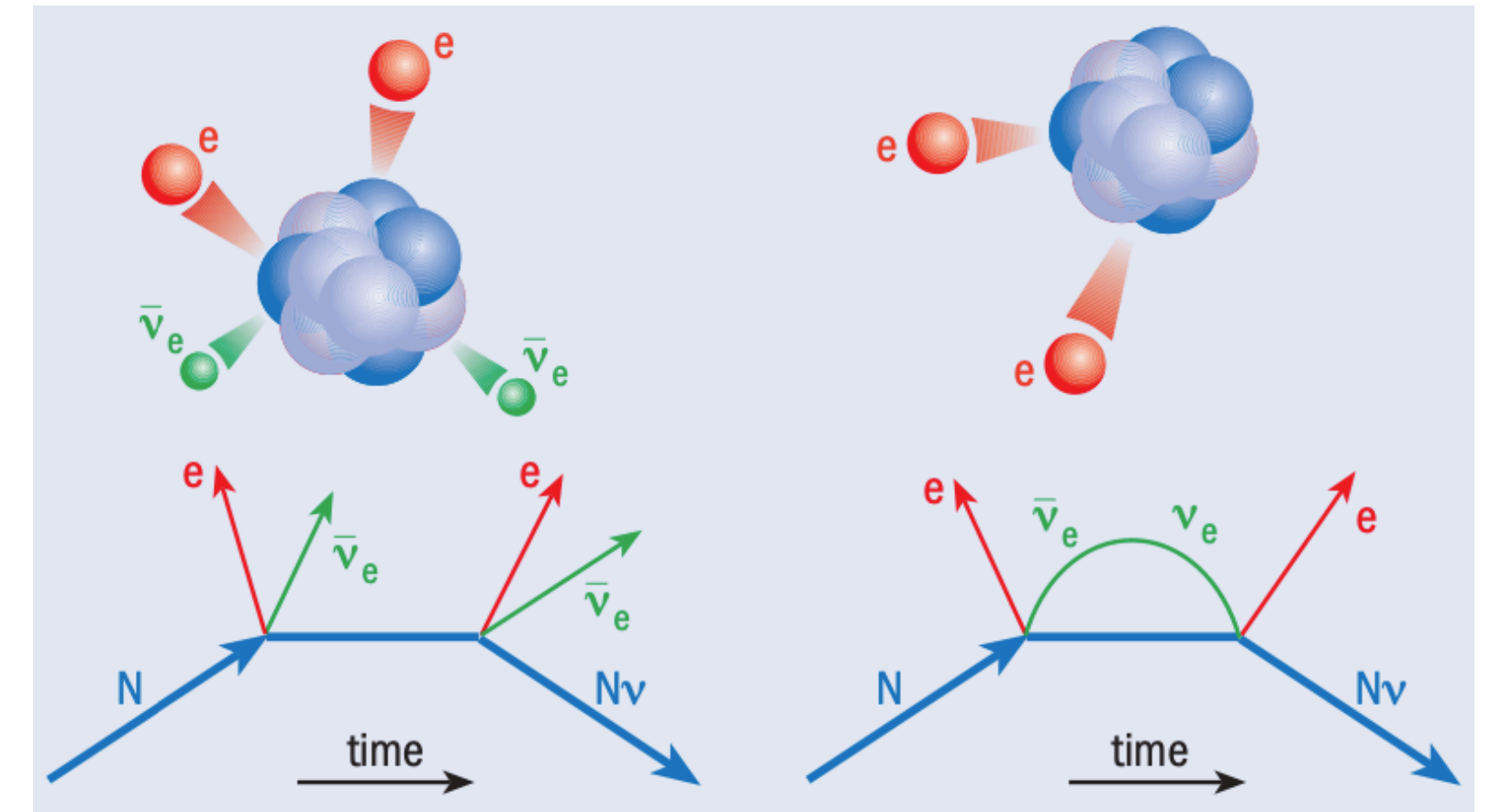
May 12, 2022

CoSSURF, Rapid City



# Motivation for Neutrinoless Double Beta Decay

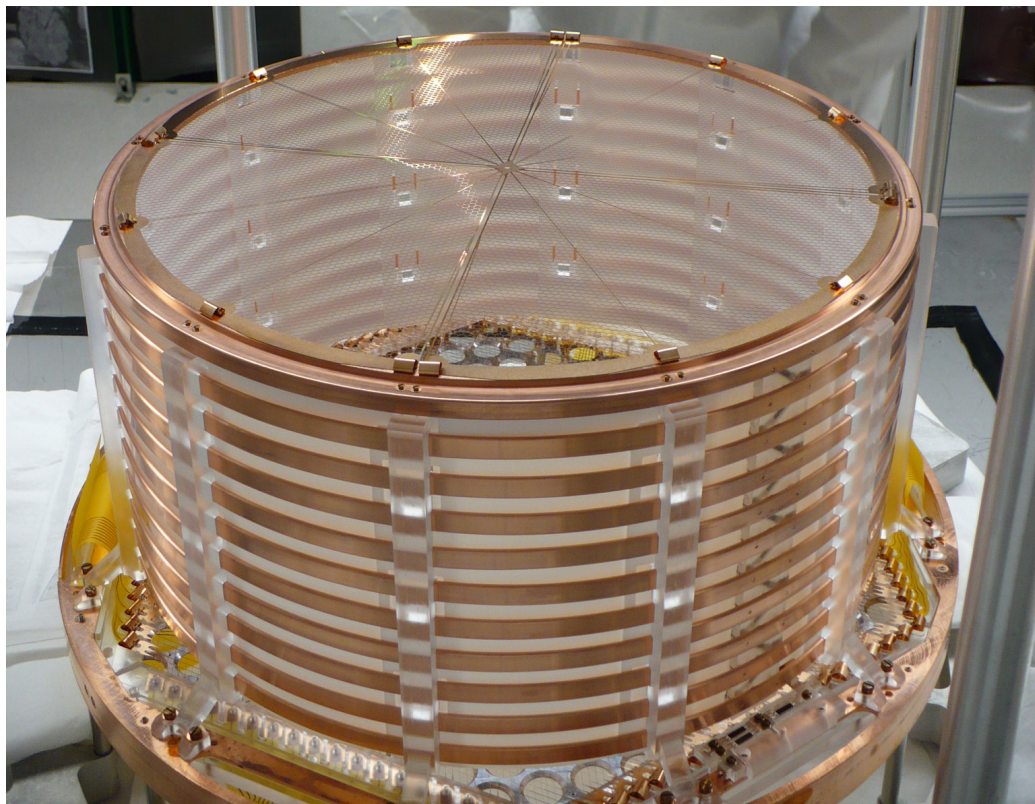
- Finding  $0\nu\beta\beta$  always implies new physics
- Lepton number violation
- Neutrinos are Majorana fermions ( $\nu \equiv \bar{\nu}$ )
- Origin of neutrino masses
- Insight into absolute neutrino mass scale
- Possibly linked to matter and anti-matter asymmetry
- Experimental signature is a peak at the Q-value (2458 keV for  $^{136}\text{Xe}$ )





# Liquid Xenon Detectors for $0\nu\beta\beta$

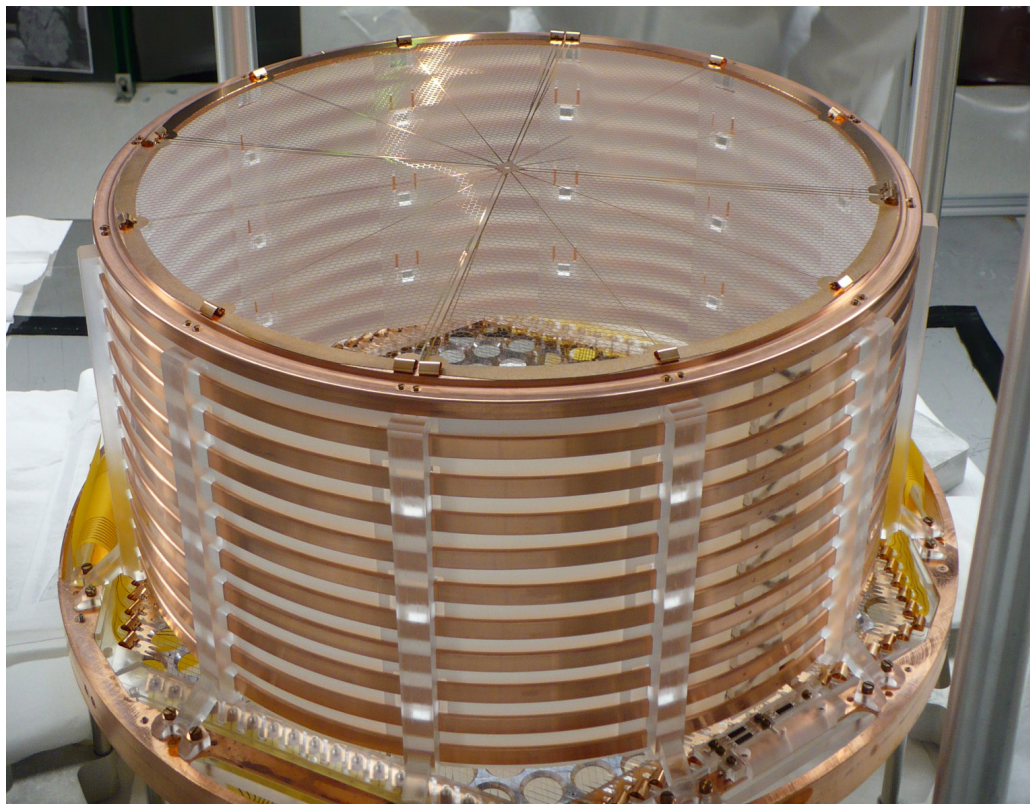
## EXO-200



- First 100 kg-class  $0\nu\beta\beta$  search
- Discovered  $2\nu\beta\beta$  in  $^{136}\text{Xe}$  in 2011
- Limit:  $T_{1/2}^{0\nu} > 3.5 \times 10^{25}$  yr
- Sensitivity:  $T_{1/2}^{0\nu} > 5.0 \times 10^{25}$  yr
- Pioneered ultra low-background LXe TPC technology

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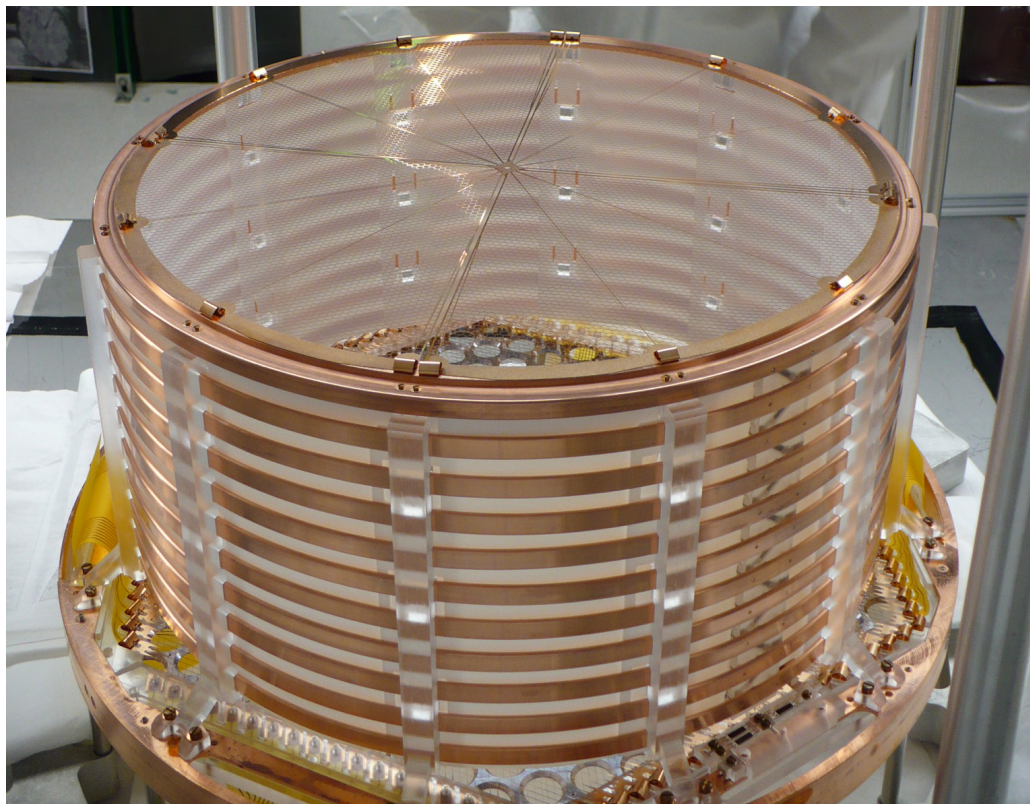
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Search for  $0\nu\beta\beta$  with the Complete EXO-200 Dataset

<https://indico.sanfordlab.org/event/28/contributions/313/>

# Liquid Xenon Detectors for $0\nu\beta\beta$

## EXO-200

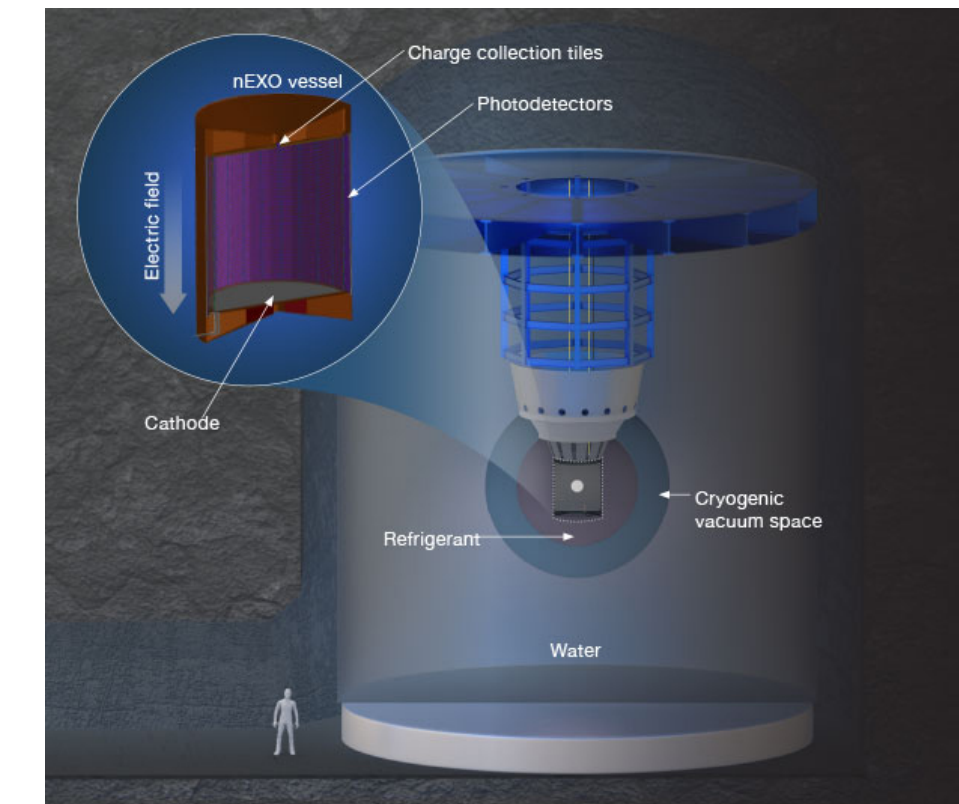


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Phys. Rev. Lett. 123, 161802 (2019)

>2 ORDERS OF MAGNITUDE  
IMPROVEMENT  
IN HALF-LIFE SENSITIVITY

## nEXO

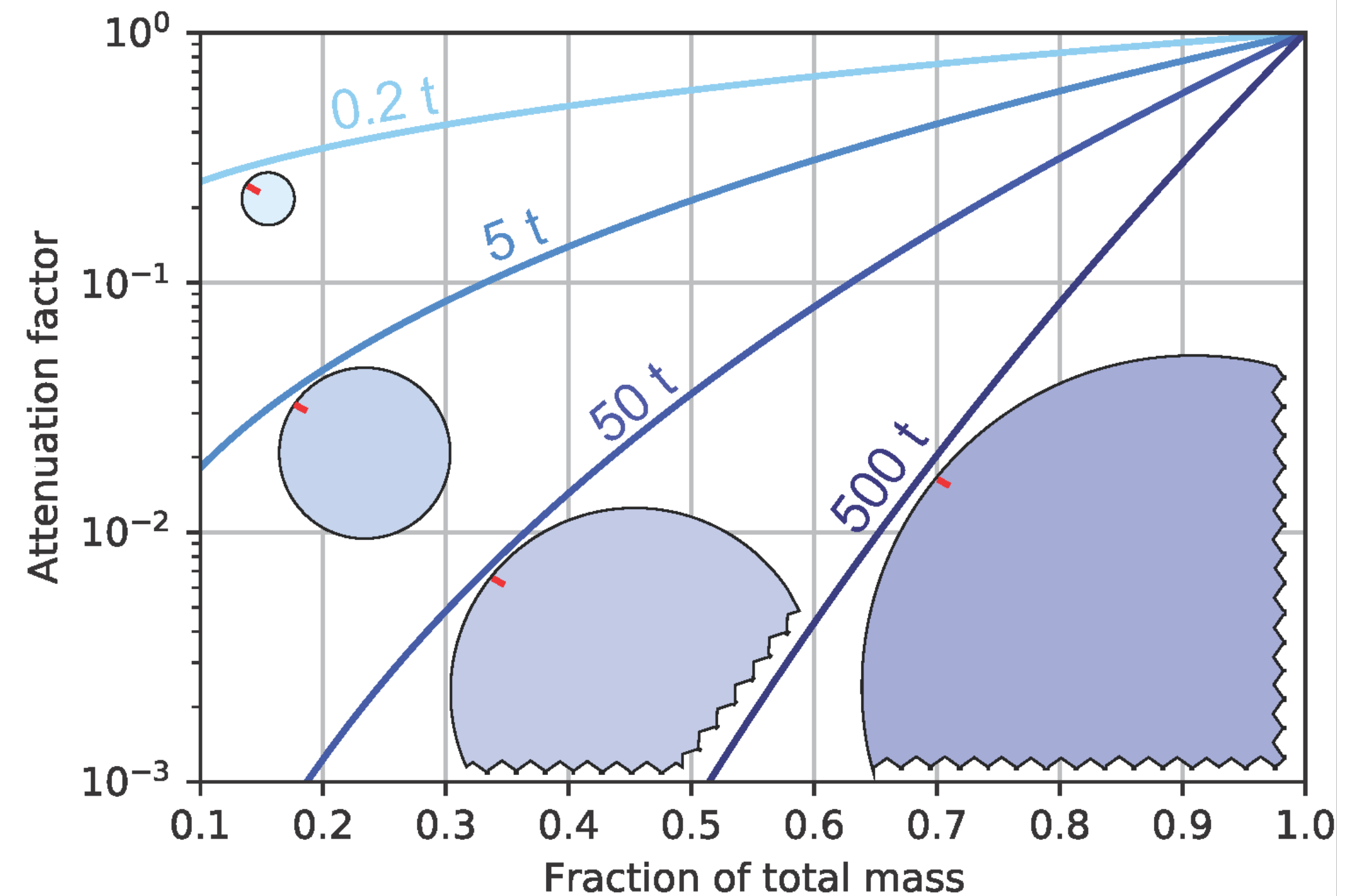


- 5 tonnes of liquid xenon
- Better self-shielding and external shielding
- Improved charge (tiles) and light (SiPM) readout
- Projected Sensitivity:  
 $T_{1/2}^{0\nu} > 1.35 \times 10^{28}$  yr

J. Phys. G: Nucl. Part. Phys. 49, 015104 (2022)

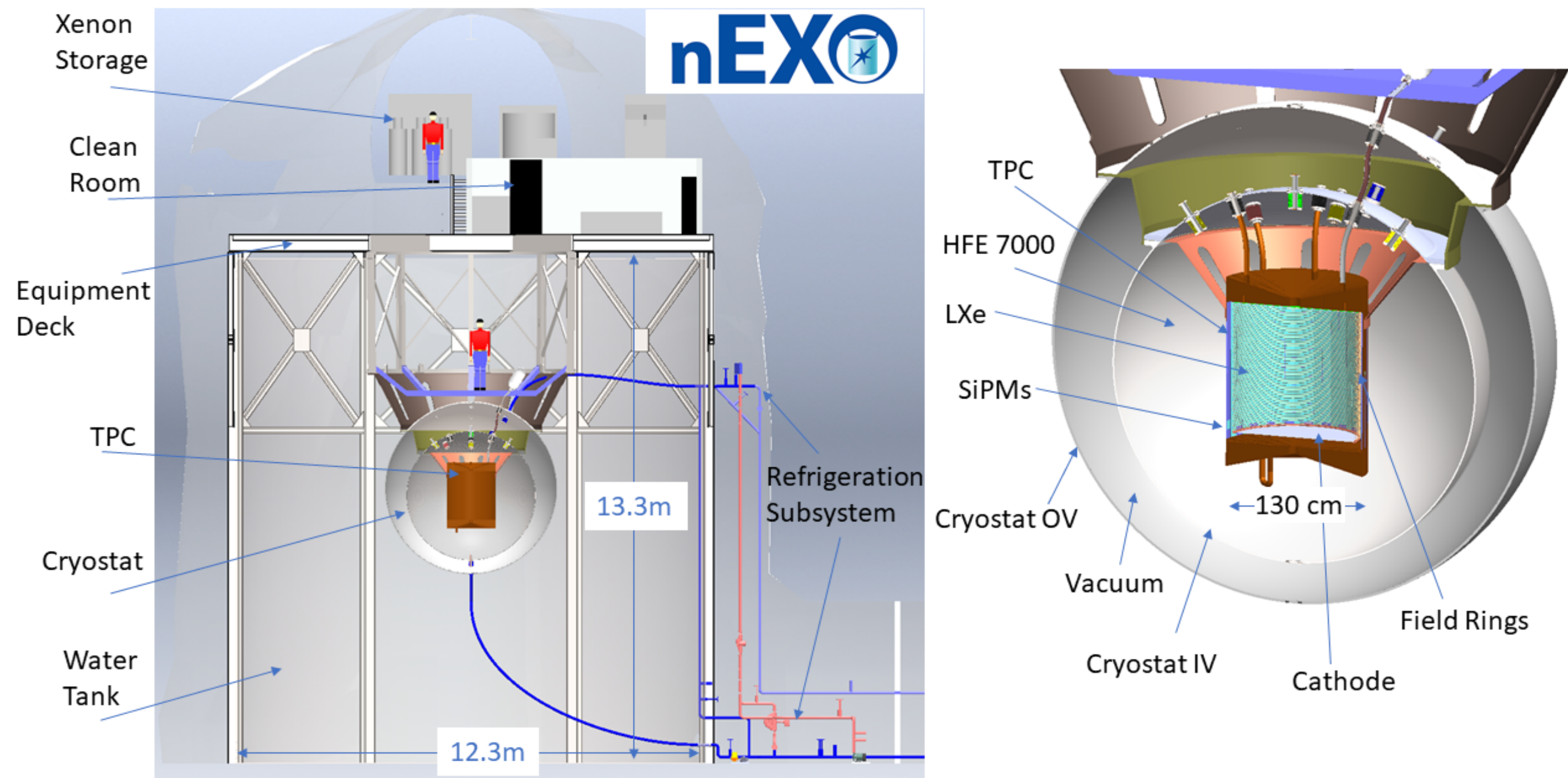
# Why a Liquid Xenon Detector?

- Advantages of the Liquid Xenon Technology for  $0\nu\beta\beta$ 
  - Scalability
  - Low intrinsic background
  - Good energy resolution
- Topological discrimination of backgrounds
- Possibility of a control experiment



# Overview of the nEXO Detector

- Single Phase Time Projection Chamber (TPC)
- **5000 kg** of liquid xenon
- **90 %** enriched in  $^{136}\text{Xe}$
- Single **1.2 m** drift length
- Energy resolution  $\sigma/Q_{\beta\beta} = 0.8 \%$
- Active water Cherenkov muon veto
- Preferred site: SNOLAB



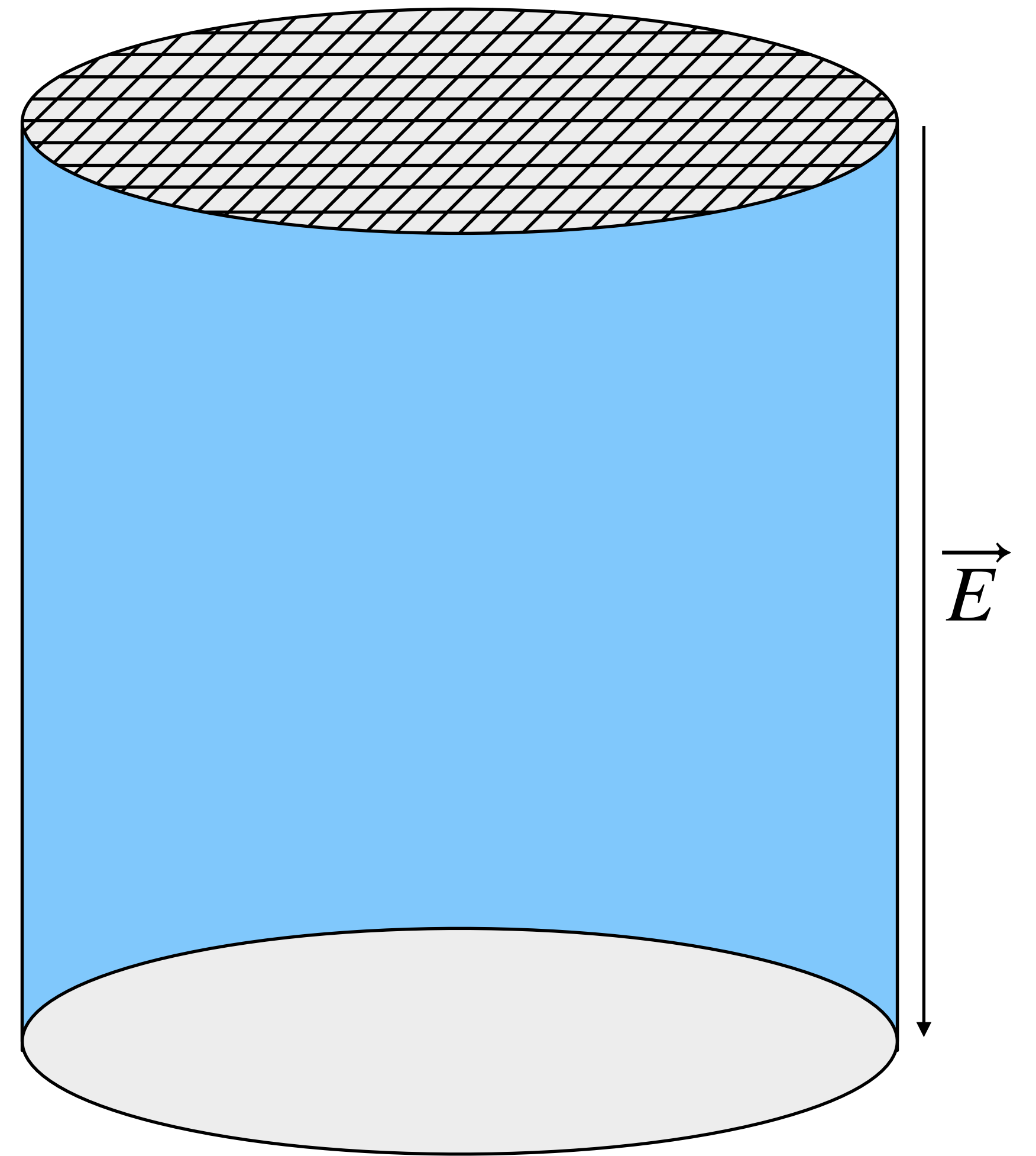
*nEXO pre-conceptual Design Report:*  
**arXiv:1805.11142**

*nEXO Sensitivity and Discovery Potential:*  
**J. Phys. G: Nucl. Part. Phys. 49, 015104 (2022)**



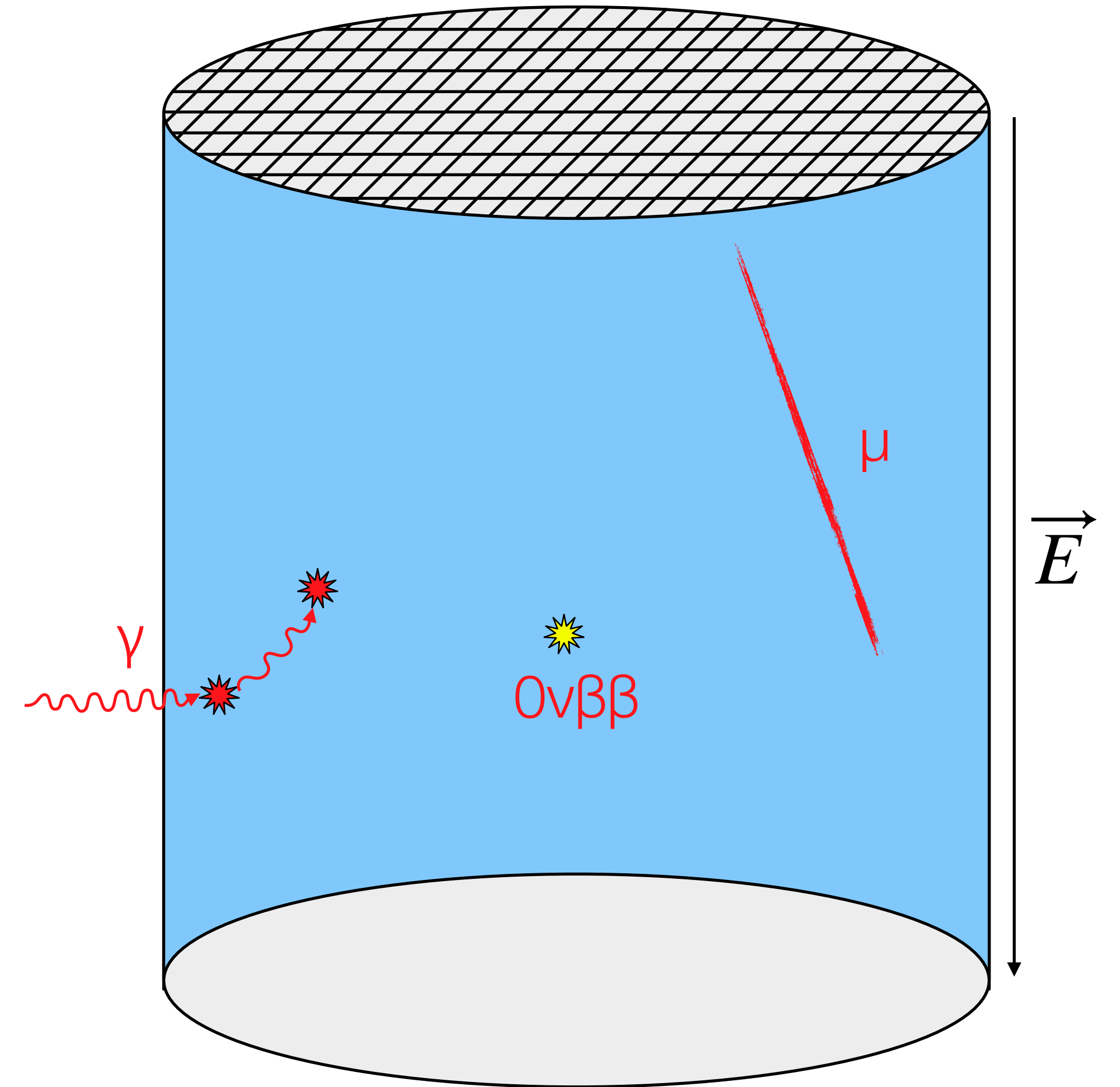
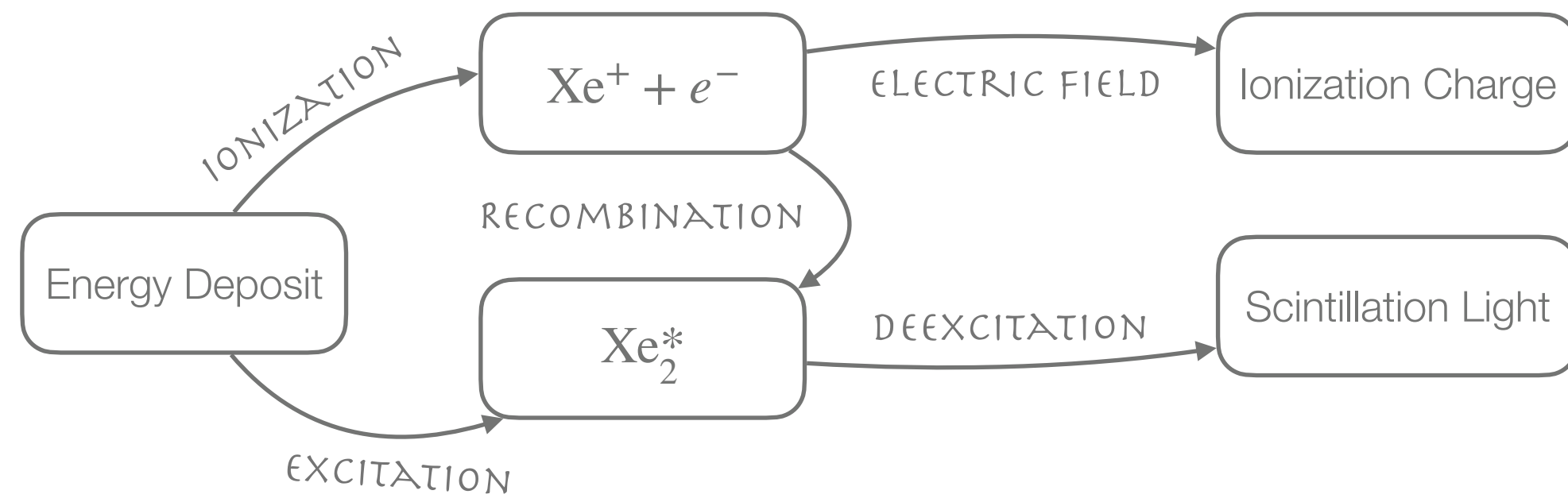
# Signal in a Liquid Xenon TPC

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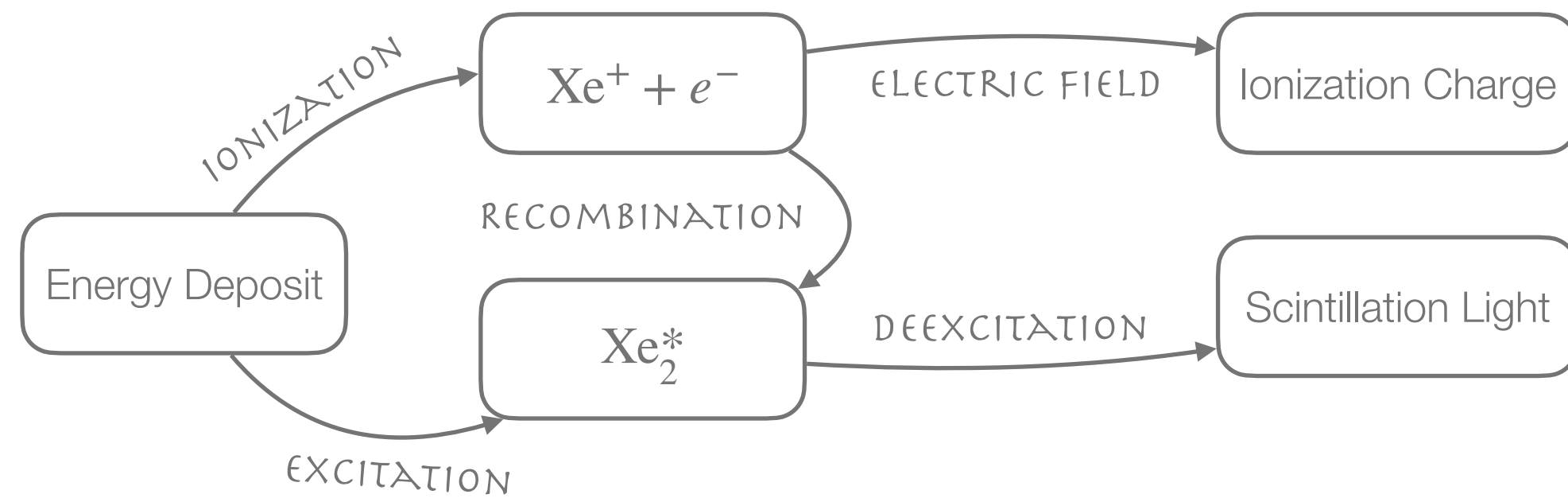
# Signal in a Liquid Xenon TPC

- Ionizing radiation will either ionize or excite Xe atoms

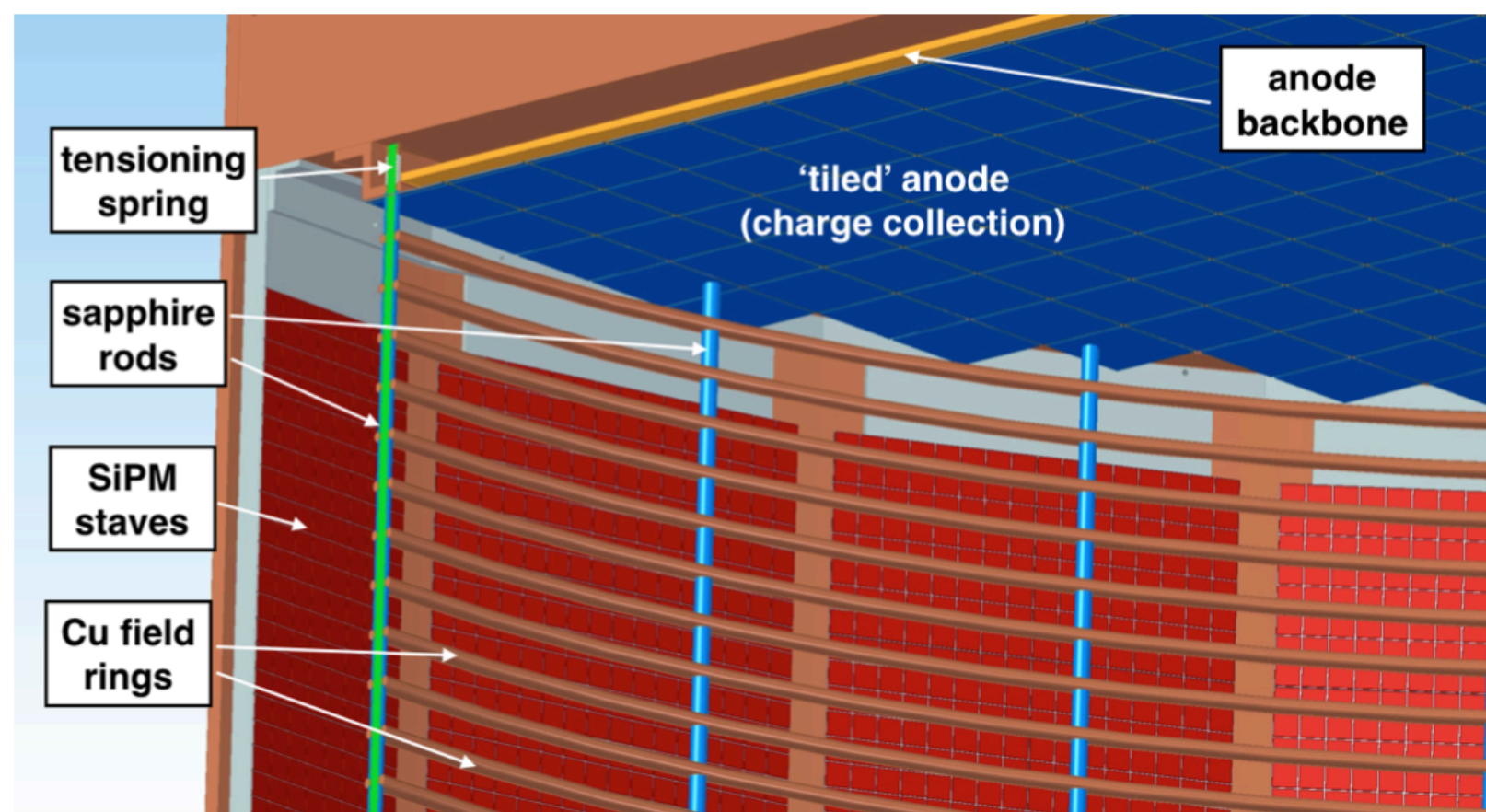
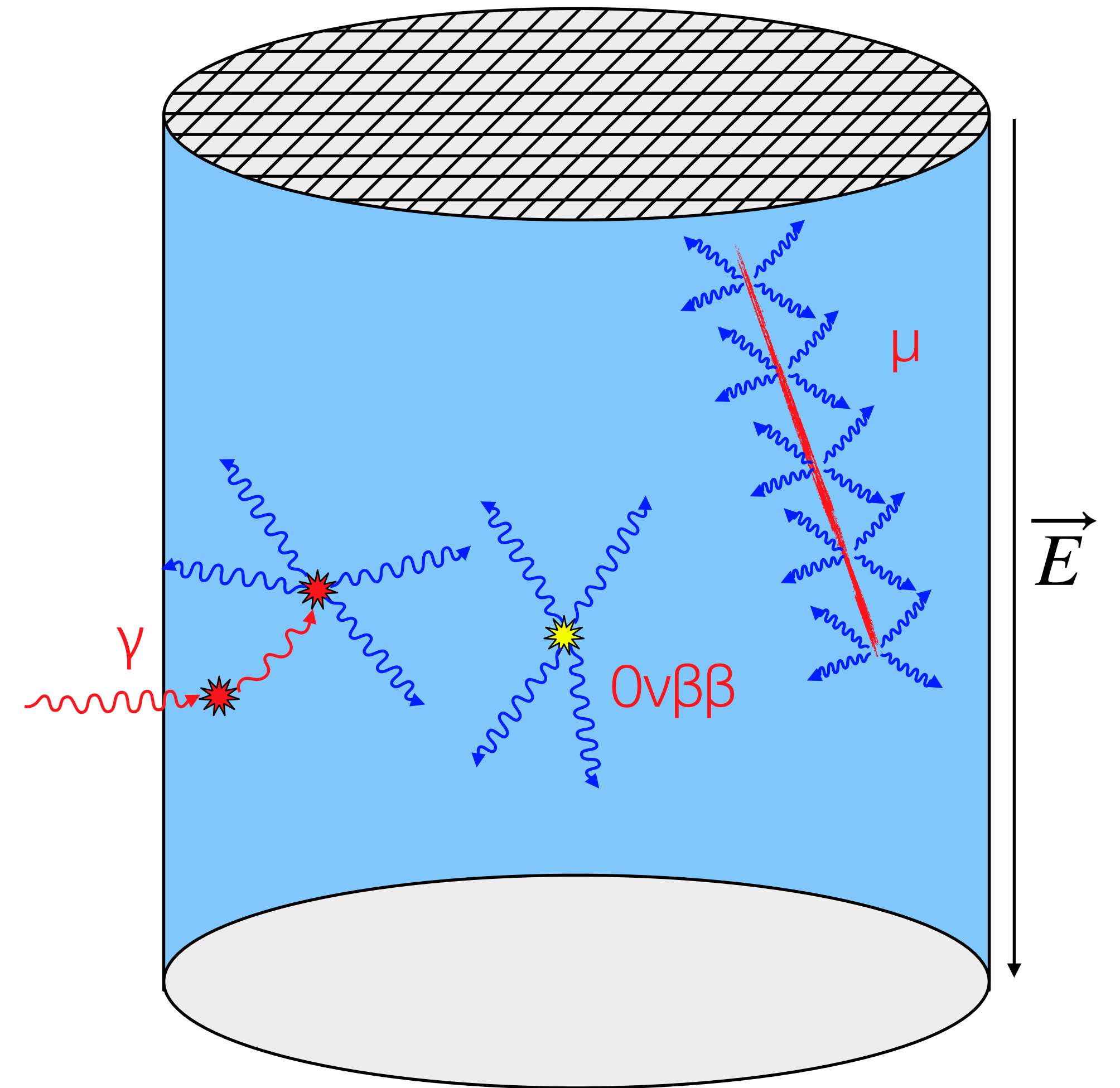


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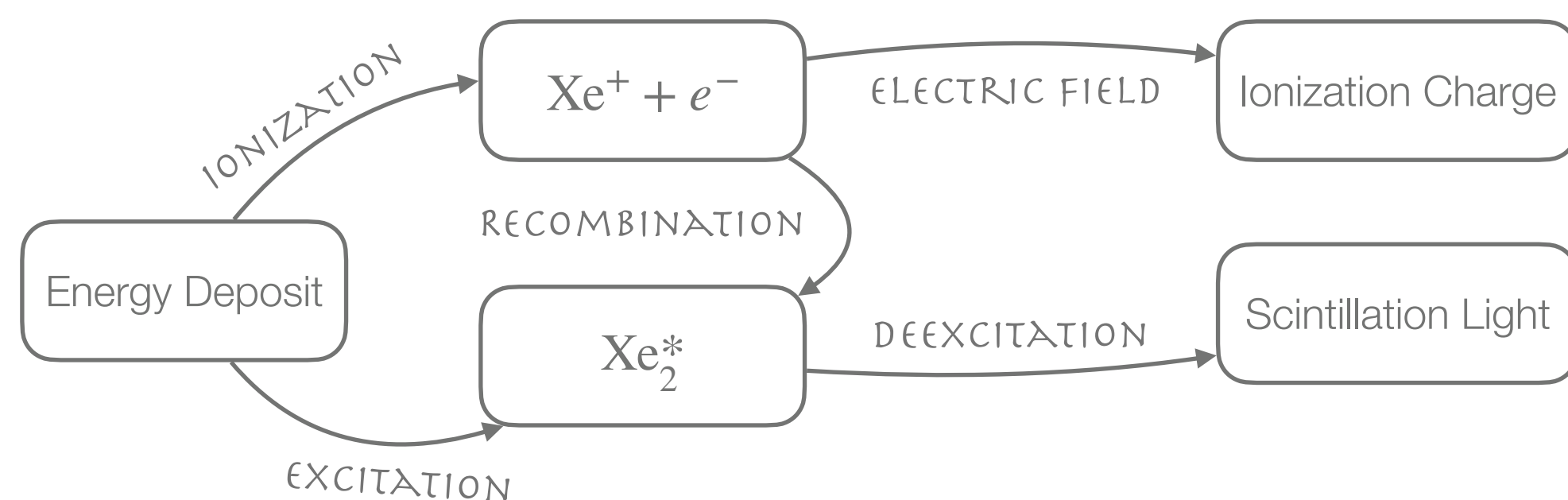


- Photons are immediately detected by the SiPMs around the barrel and provide a time stamp

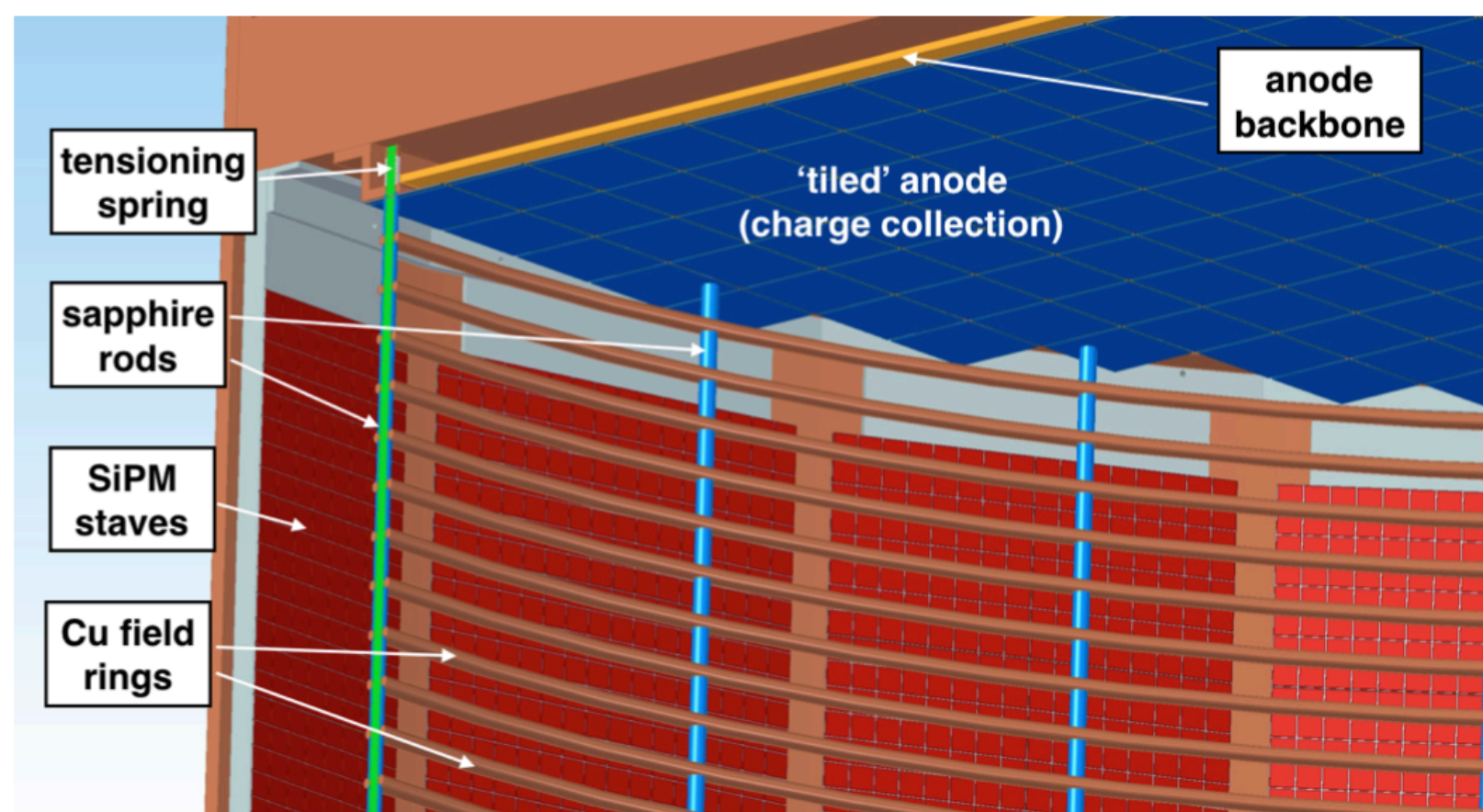
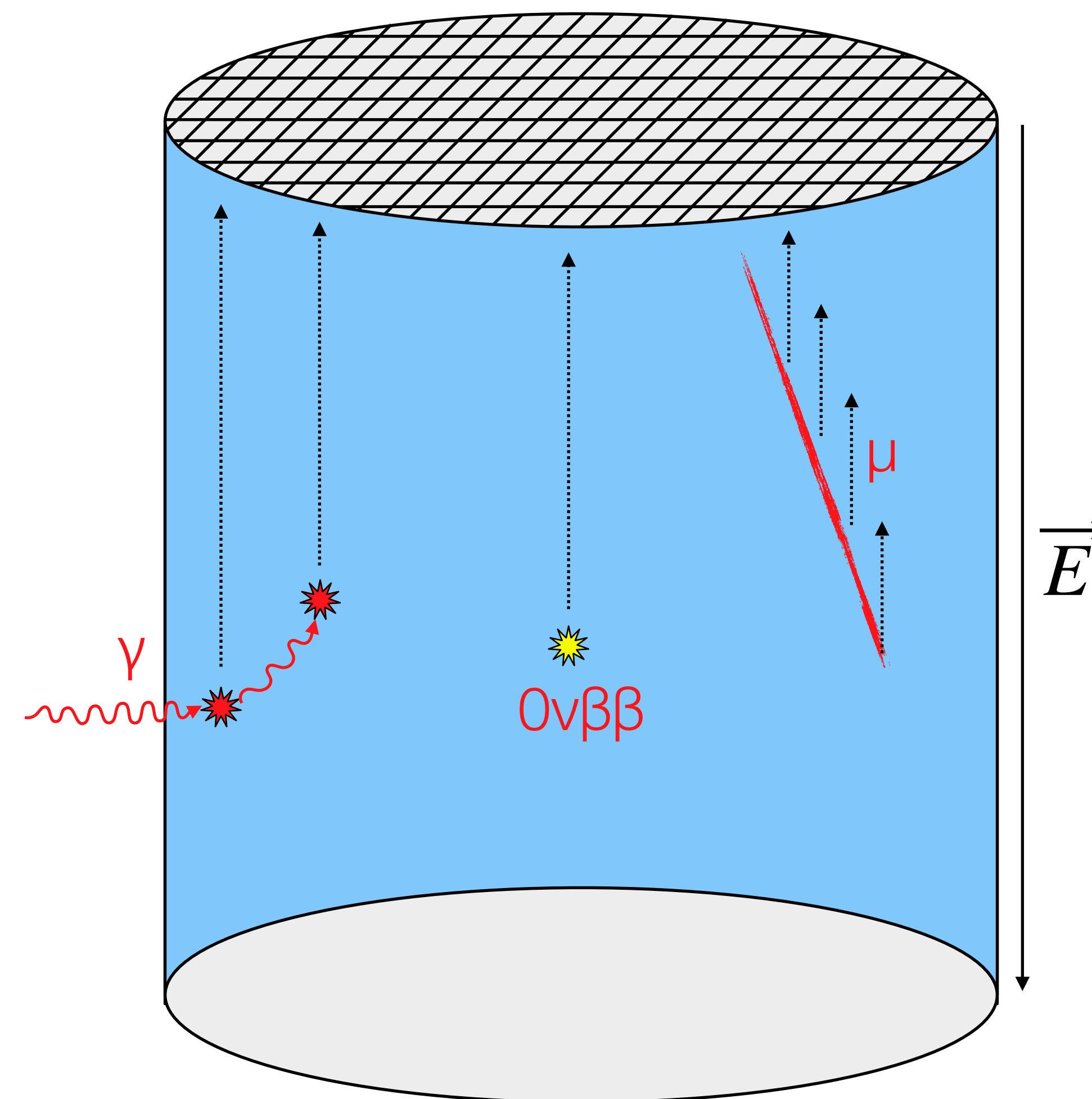


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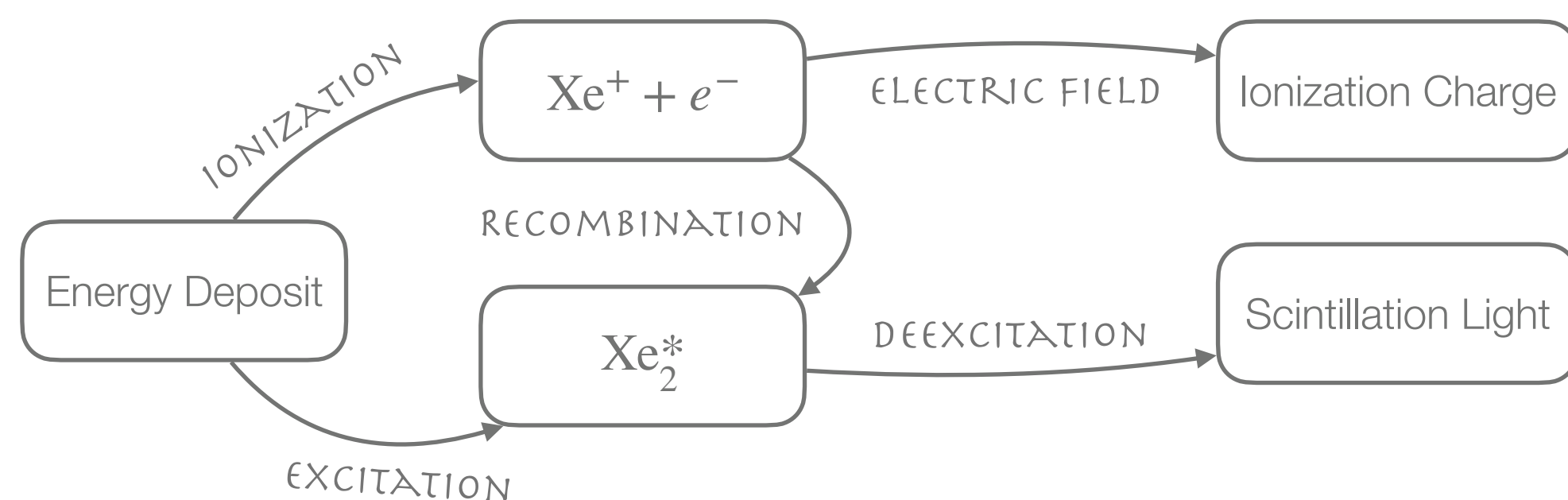


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- Electrons are drifted to charge collection tiles at the top

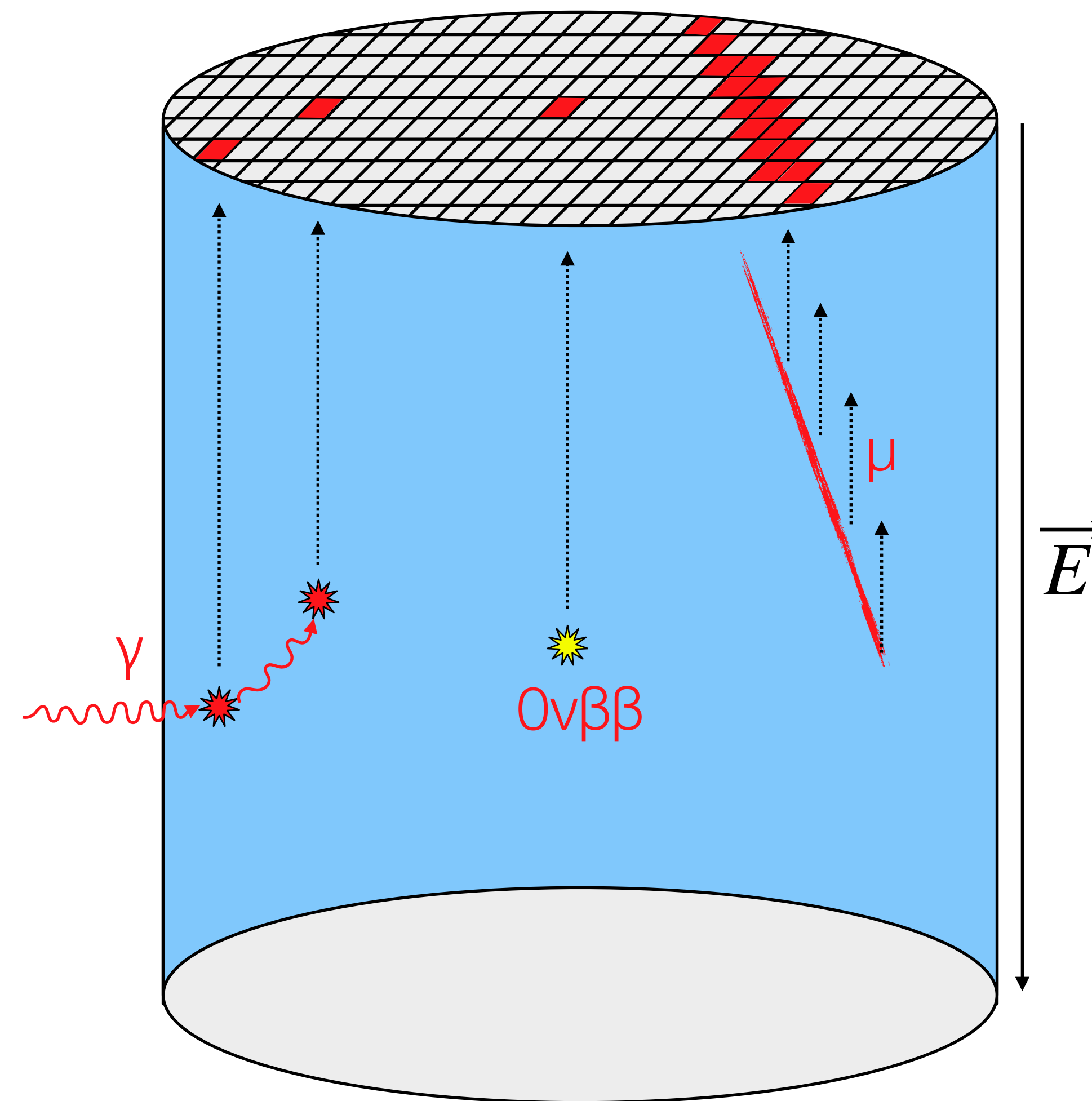


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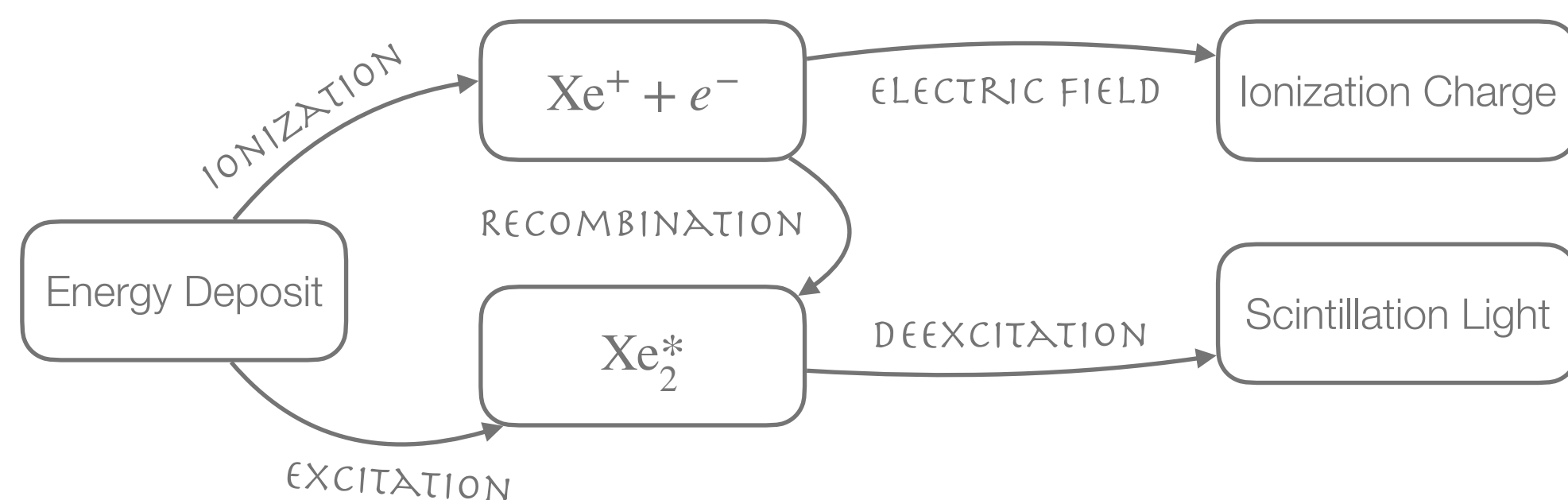


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- Charge collection tiles with 3mm pitch strips detect  $e^-$

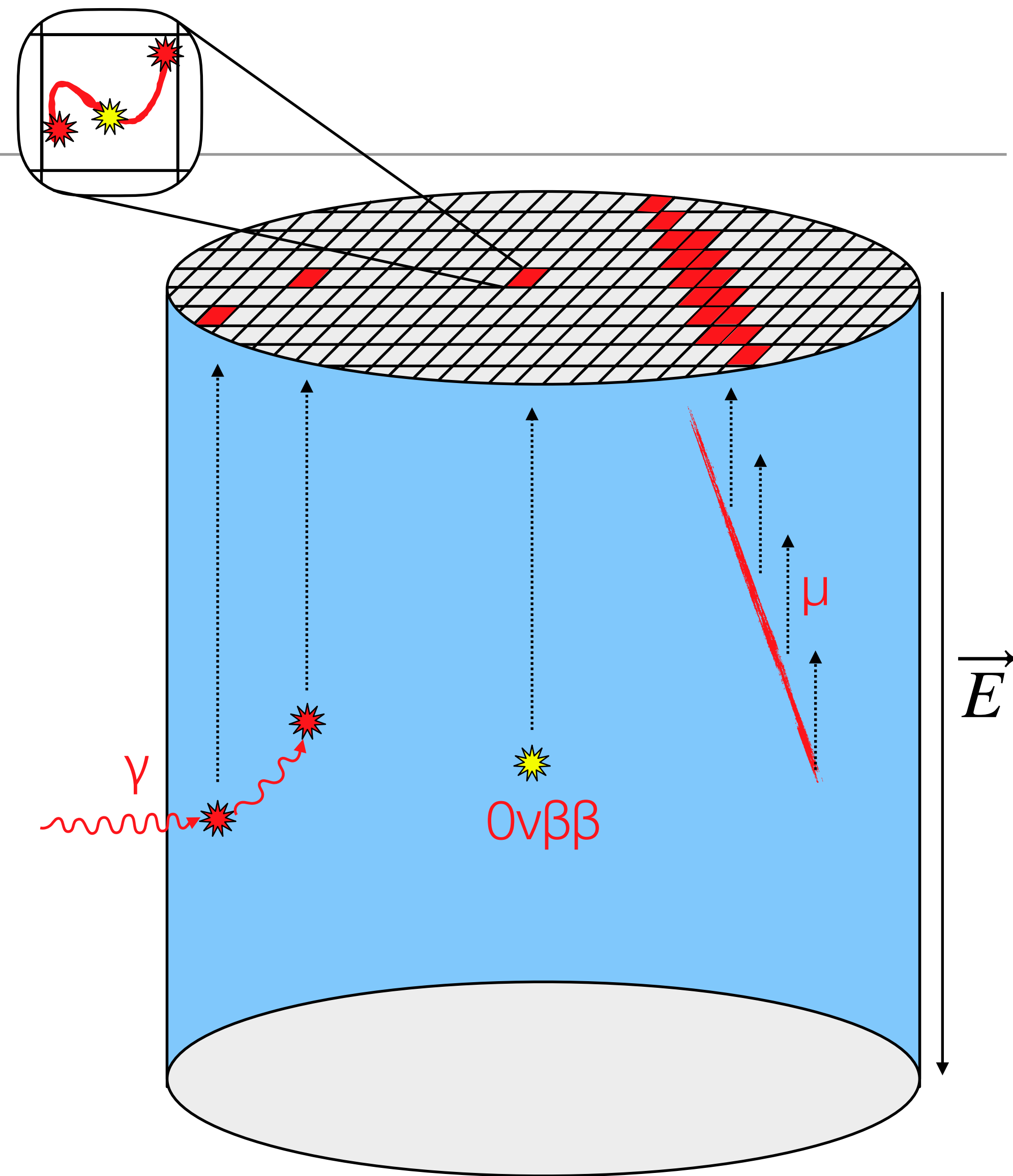


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- Photons are immediately detected by the SiPMs around the barrel and provide a time stamp
- Electrons are drifted to charge collection tiles at the top
- Charge collection tiles with 3mm pitch strips detect  $e^-$
- $0\nu\beta\beta$  charge is mostly spatially contained (unlike  $\gamma$ 's)



# Multidimensional Event Analysis

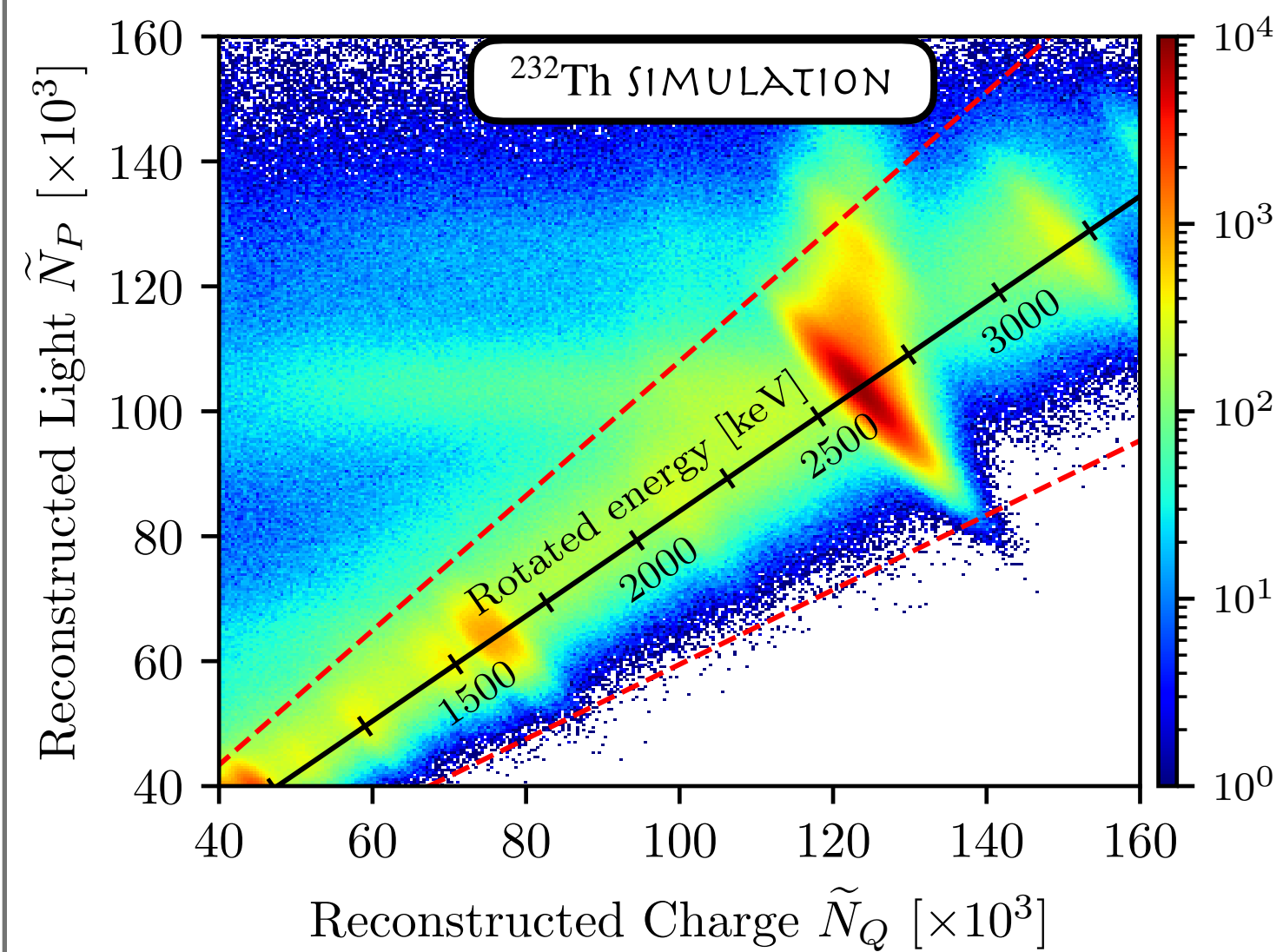
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# Multidimensional Event Analysis

## Energy

Anti-correlation between number of light and charge quanta can be exploited for improved energy resolution

—> Expect  $\sigma/Q_{\beta\beta} = 0.8\%$



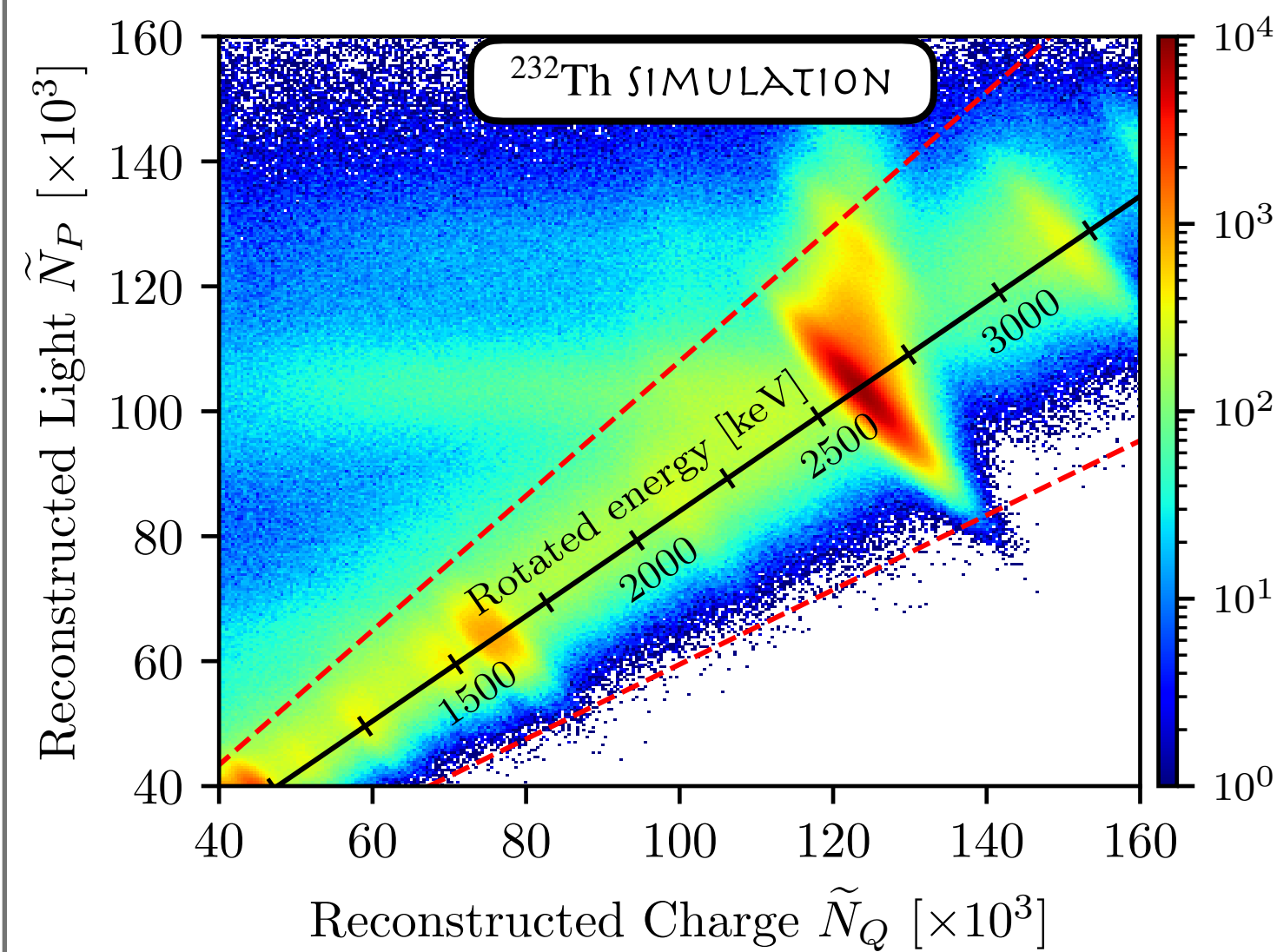


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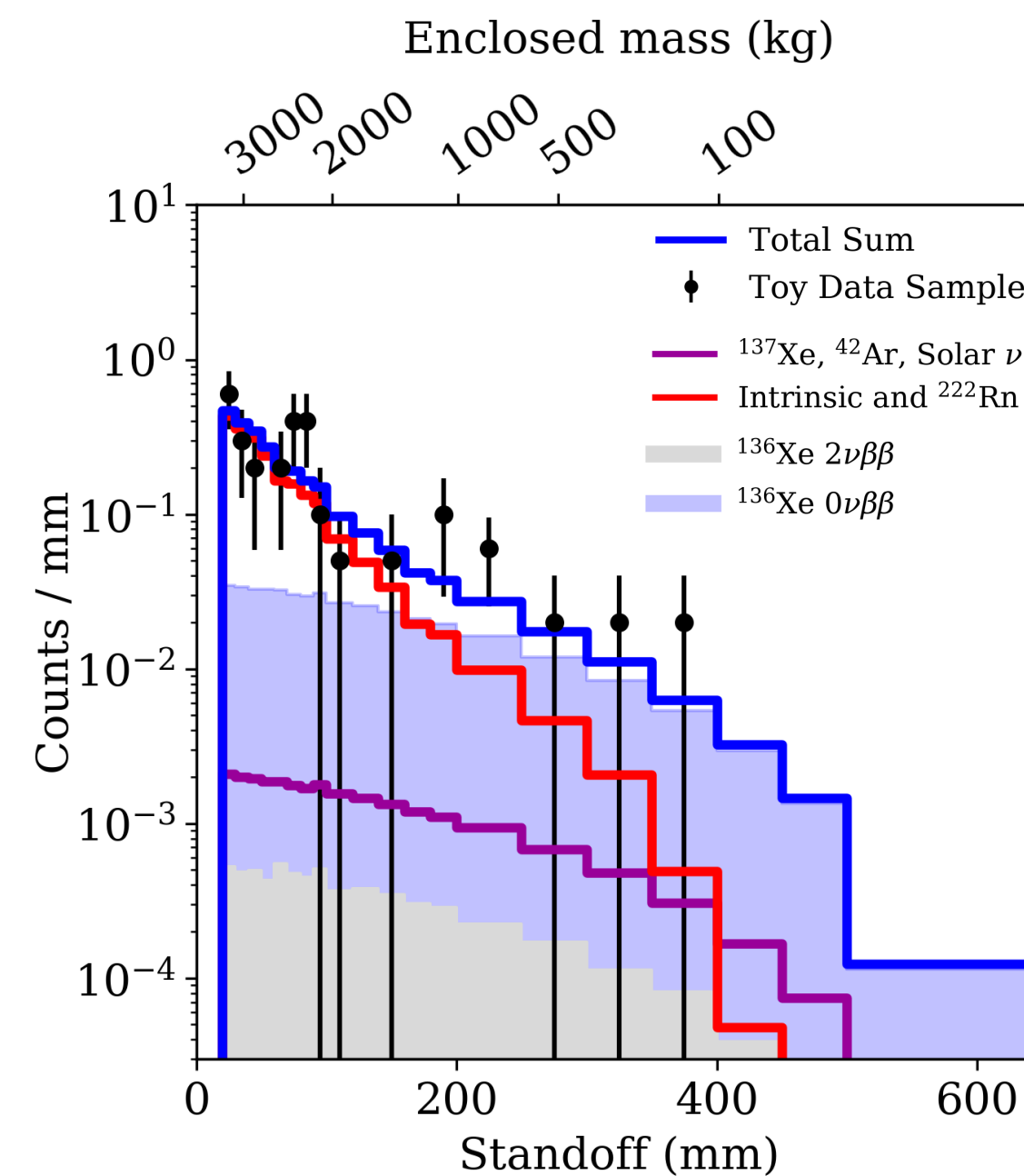
→ Expect  $\sigma/Q_{\beta\beta} = 0.8\%$



## Standoff Distance

Most backgrounds originate from outside the fiducial volume

$0\nu\beta\beta$  events are uniformly distributed

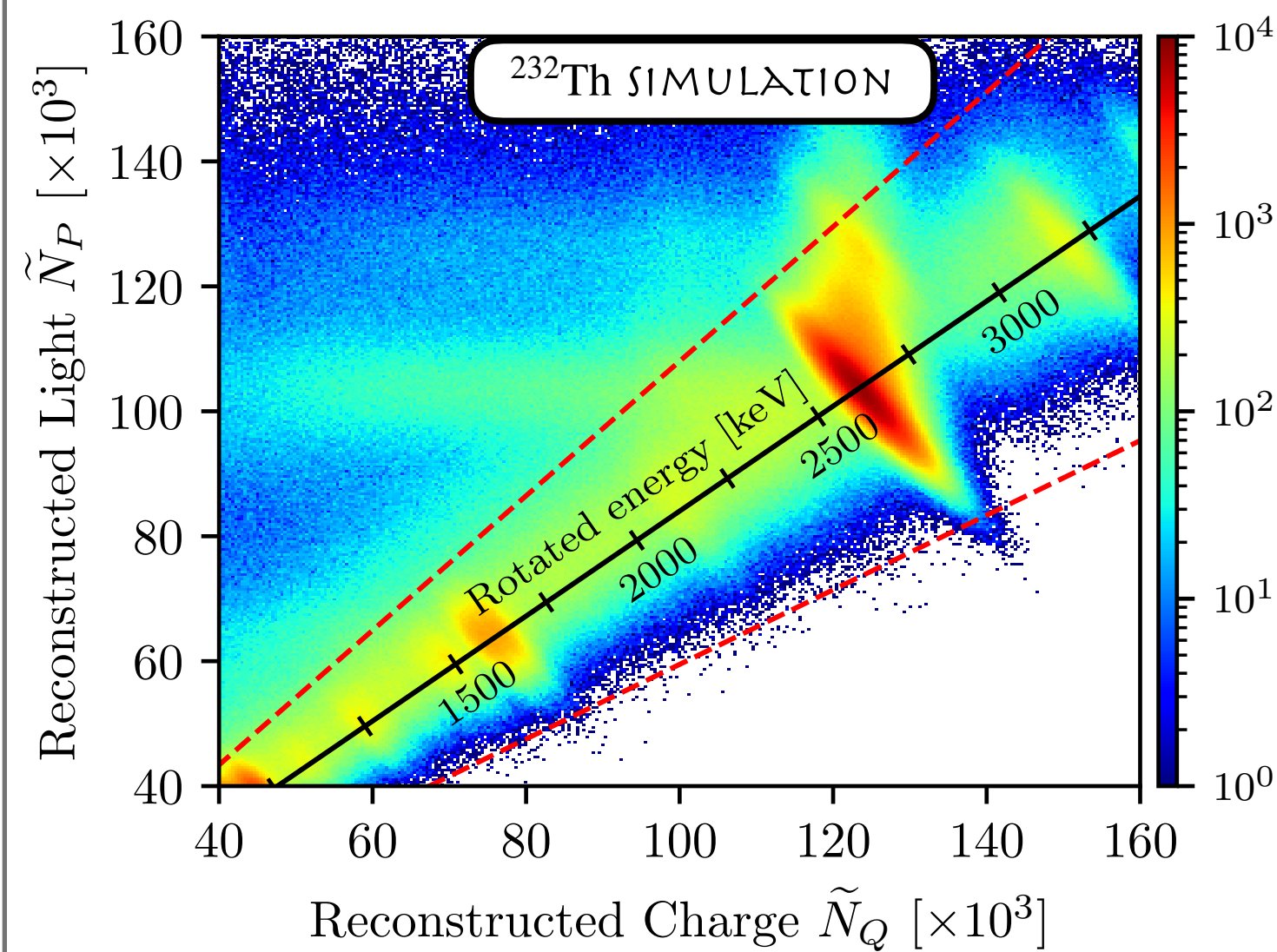


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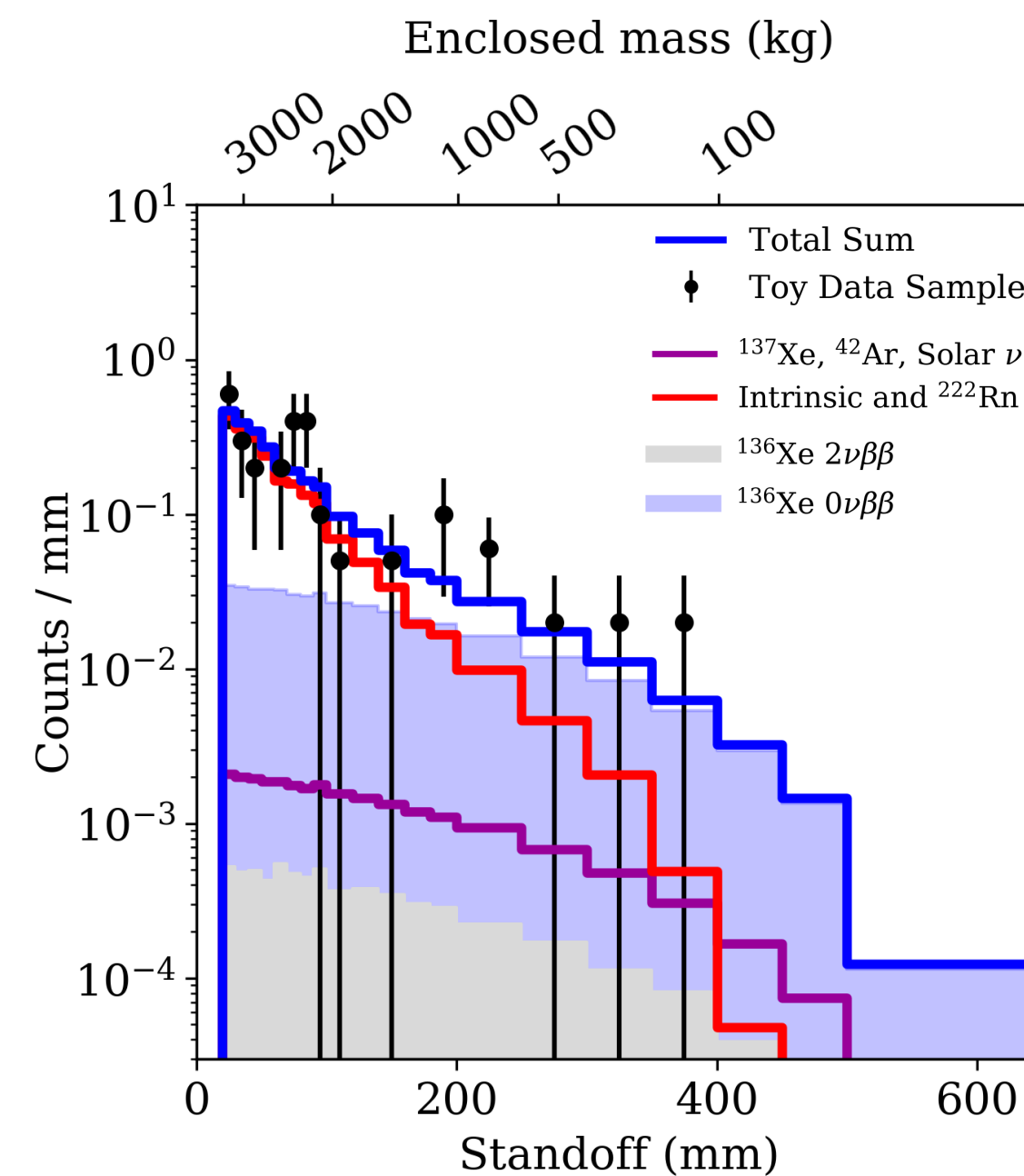
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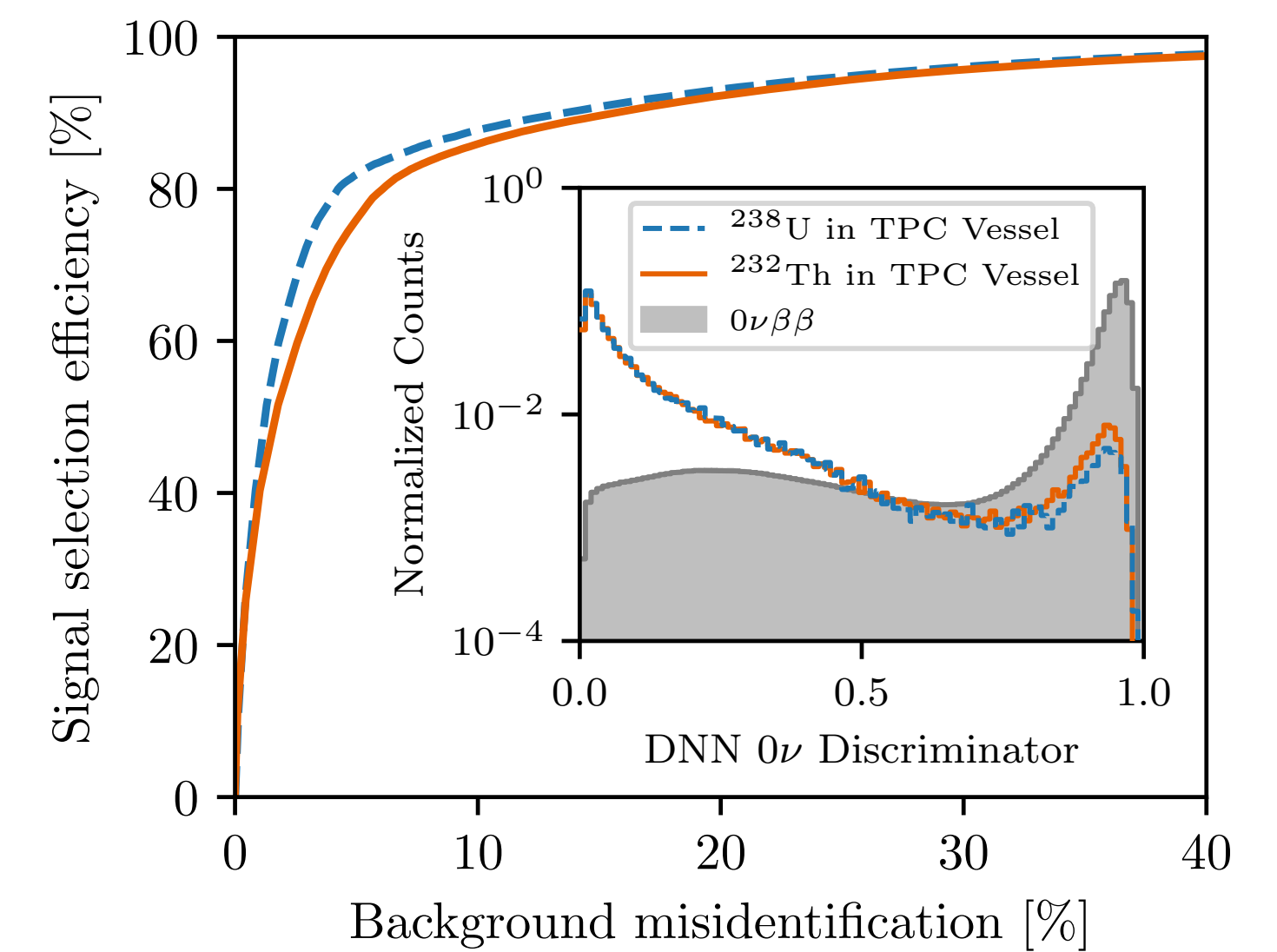
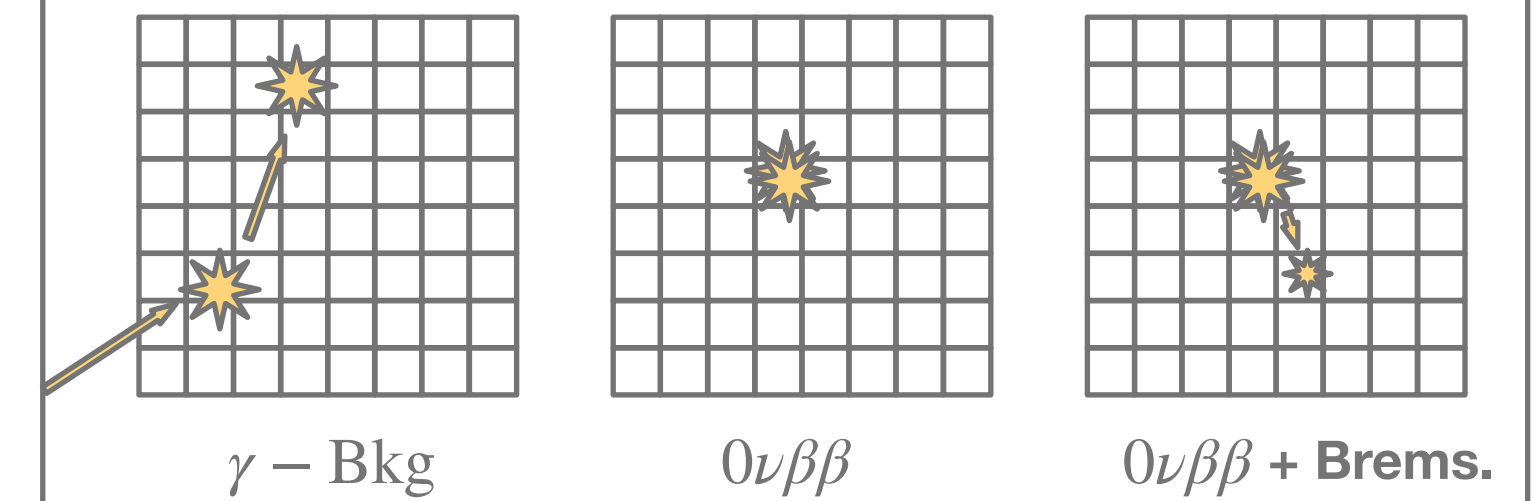
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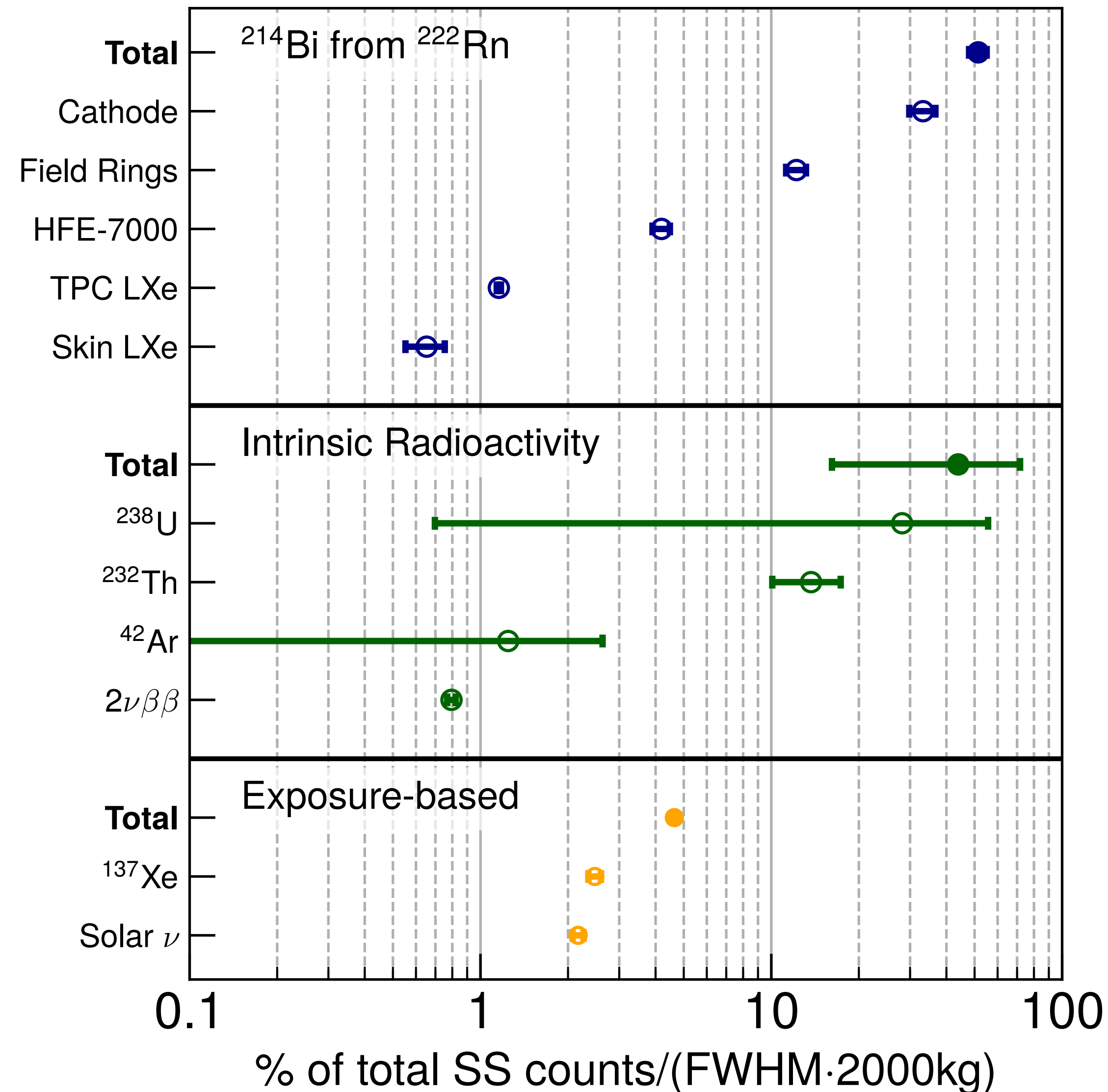


## Topology



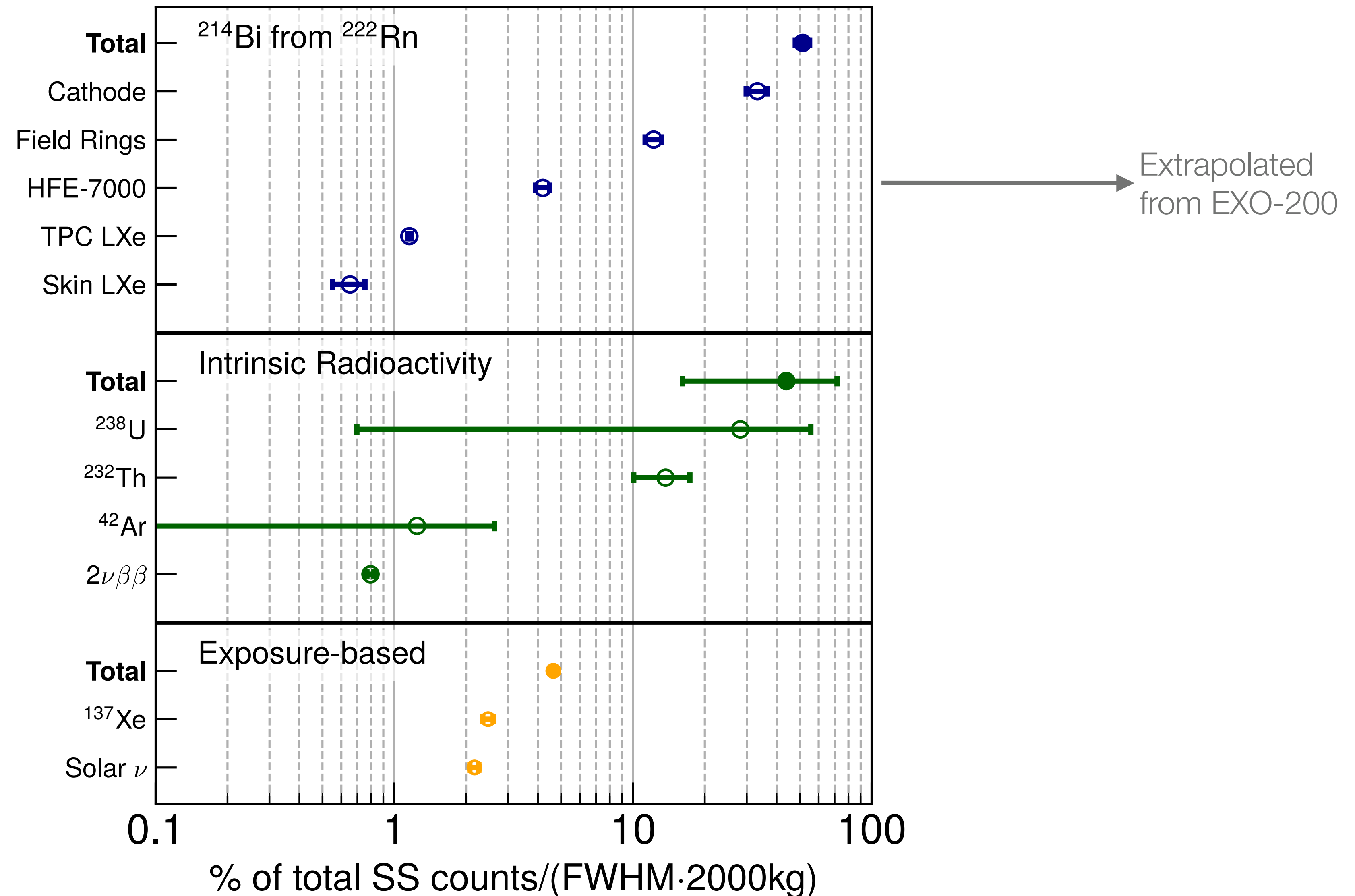
# Data-Drive Background Model

- Bottom-up estimate of background budget through extensive screening of all detector materials
- In-house electro-formed copper for some TPC components to significantly reduce intrinsic radioactivity from  $^{238}\text{U}$  and  $^{232}\text{Th}$
- Ongoing R&D looking into further reduction of  $^{222}\text{Rn}$
- Cosmogenically produced  $^{137}\text{Xe}$  can be vetoed with at least 70% efficiency and negligible lifetime loss



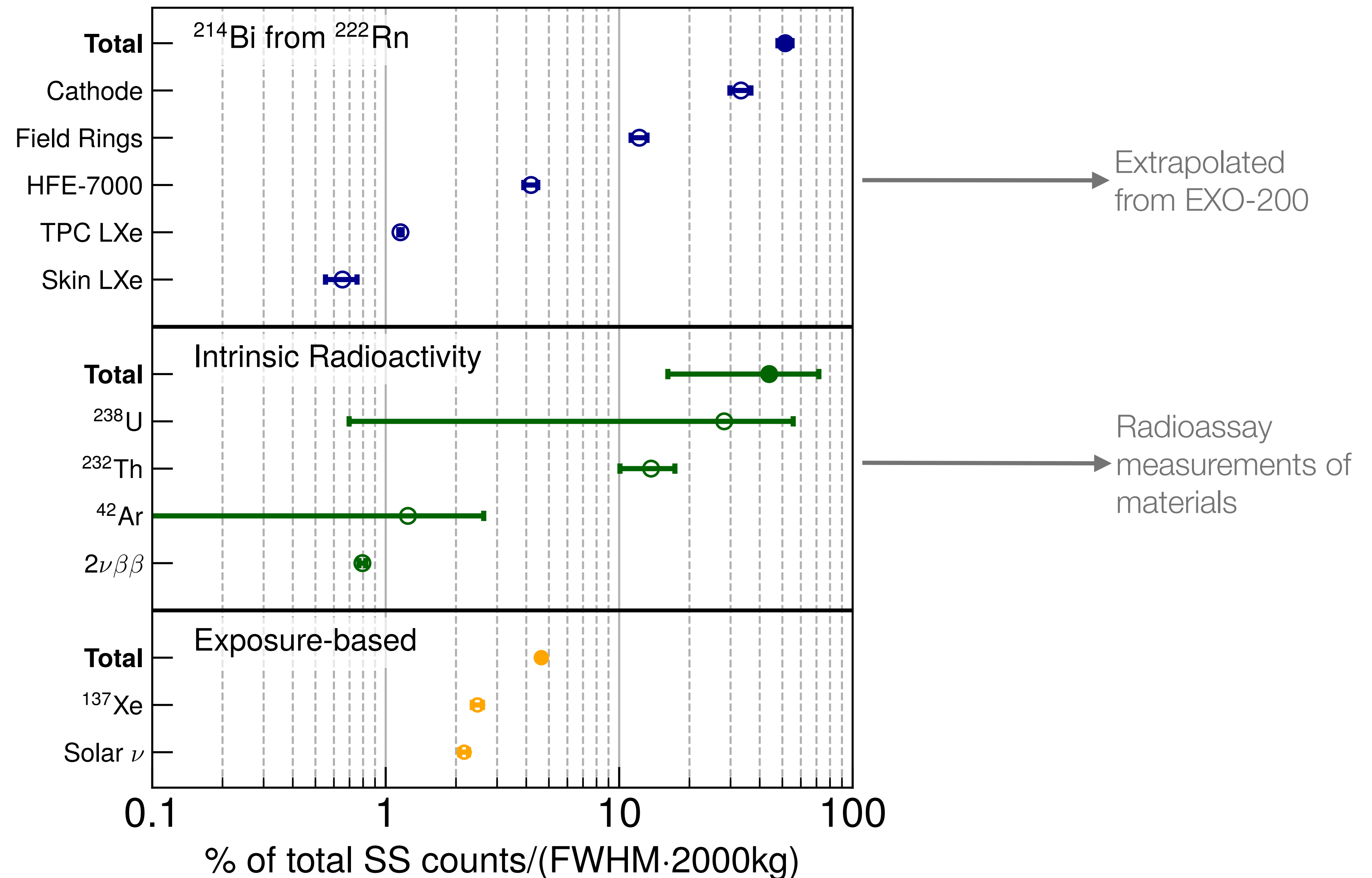
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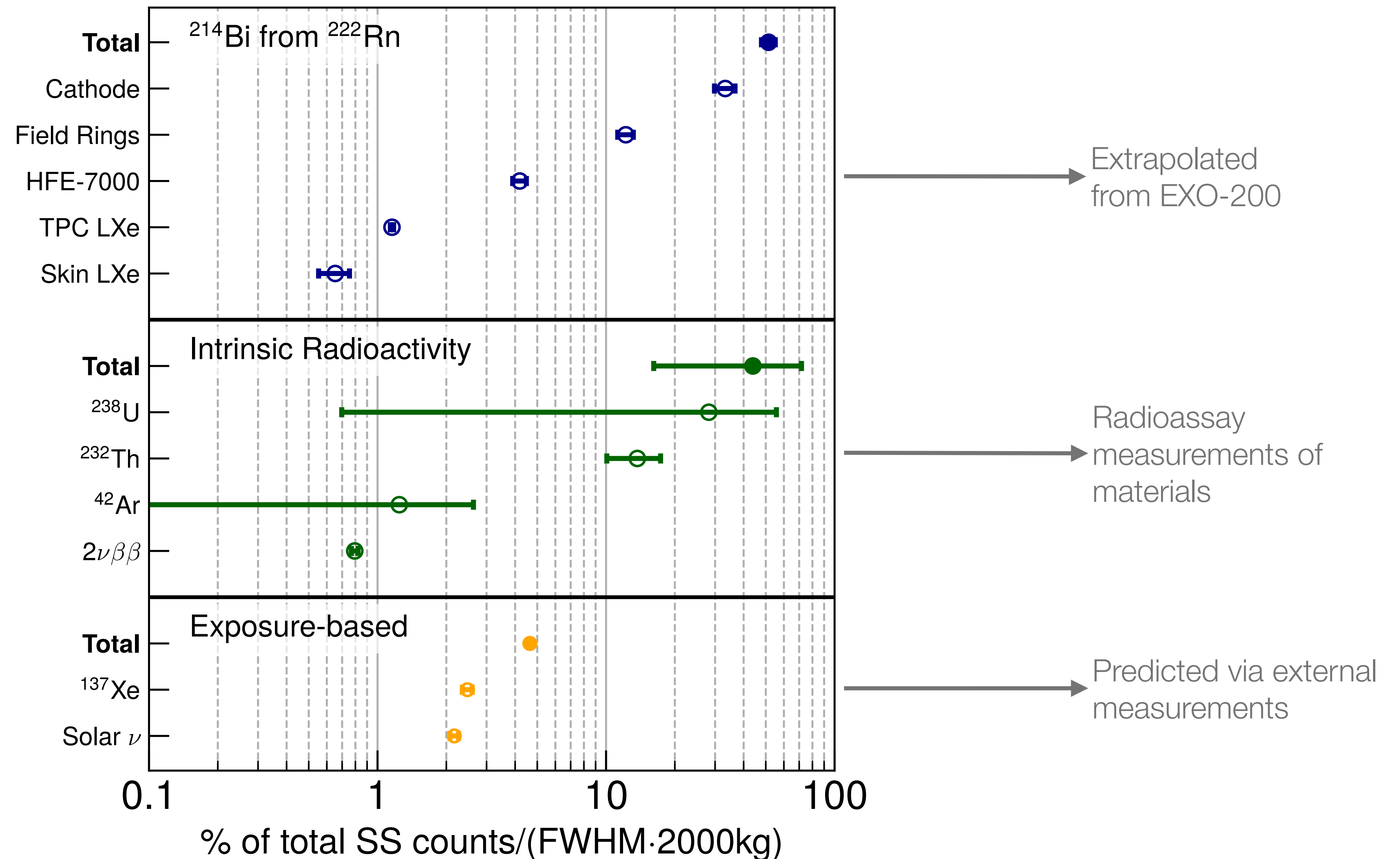
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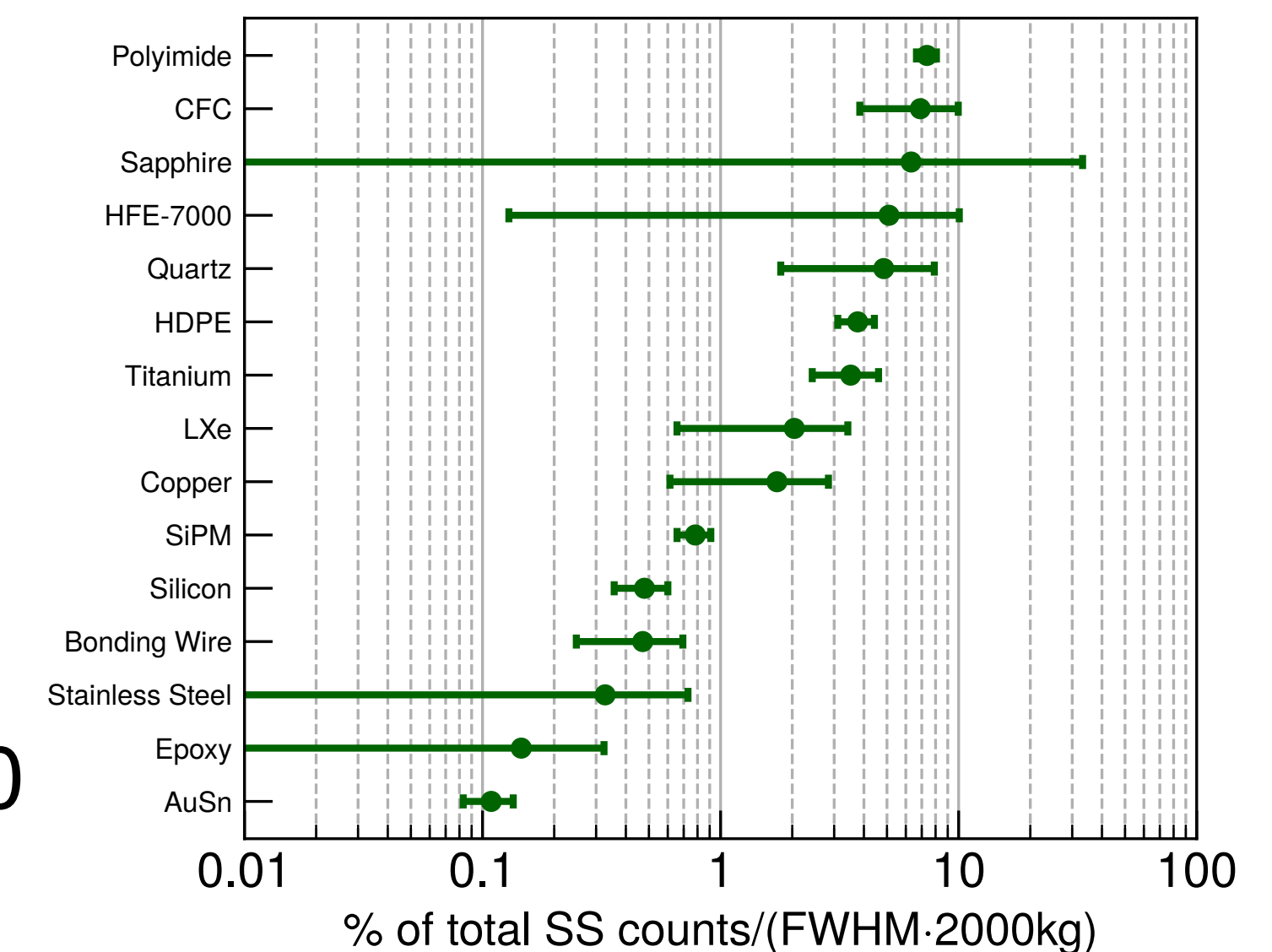
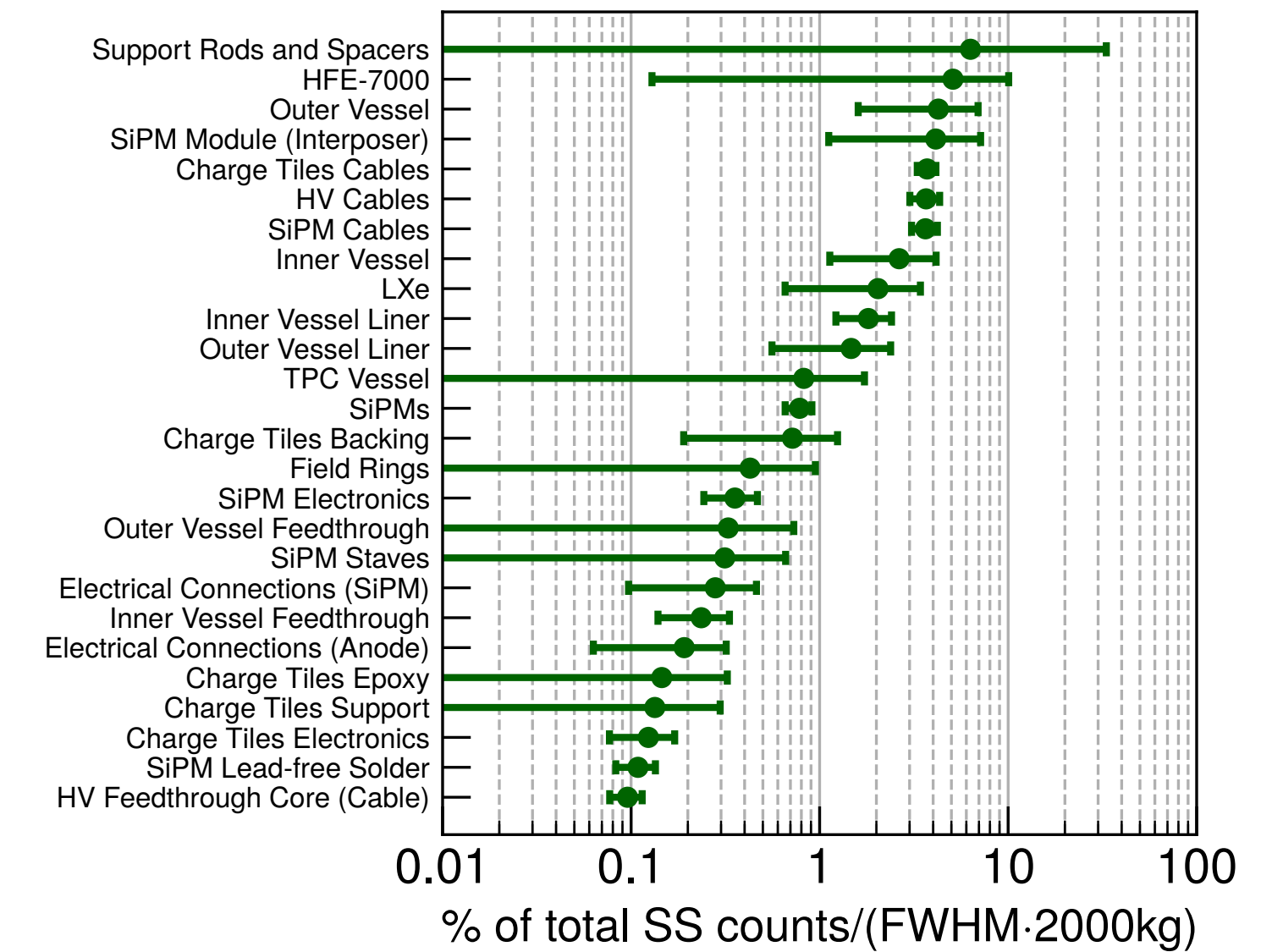
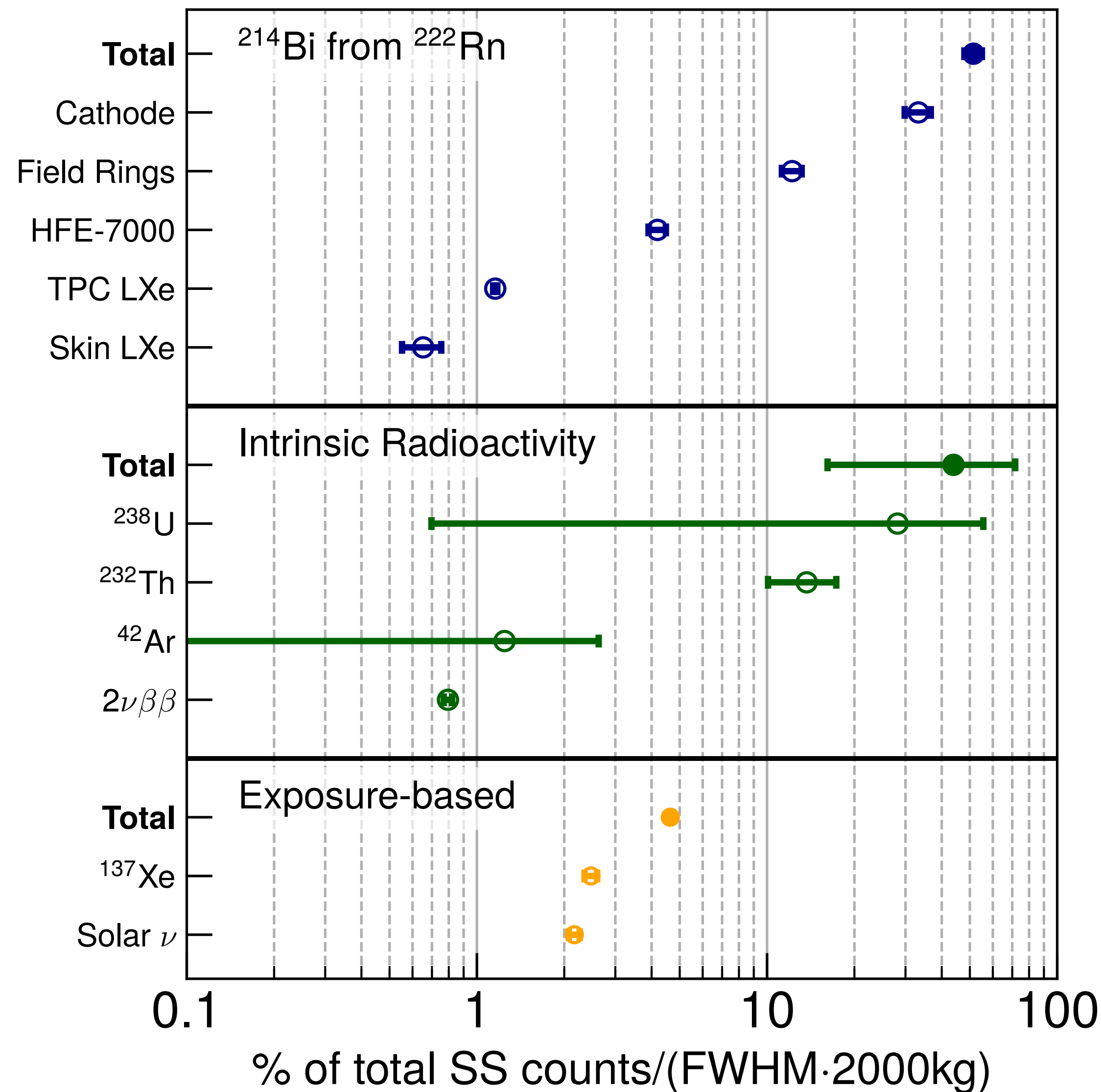
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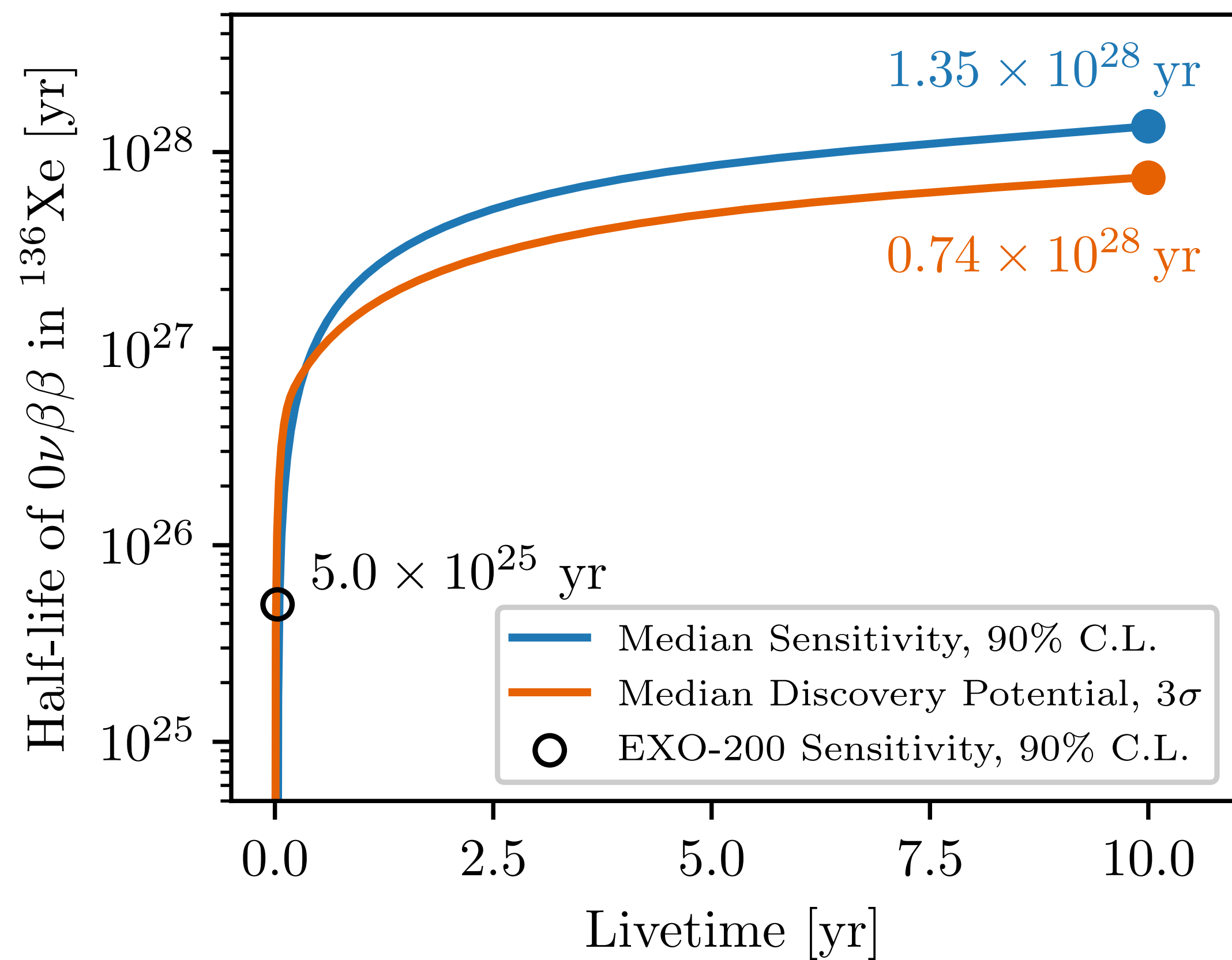


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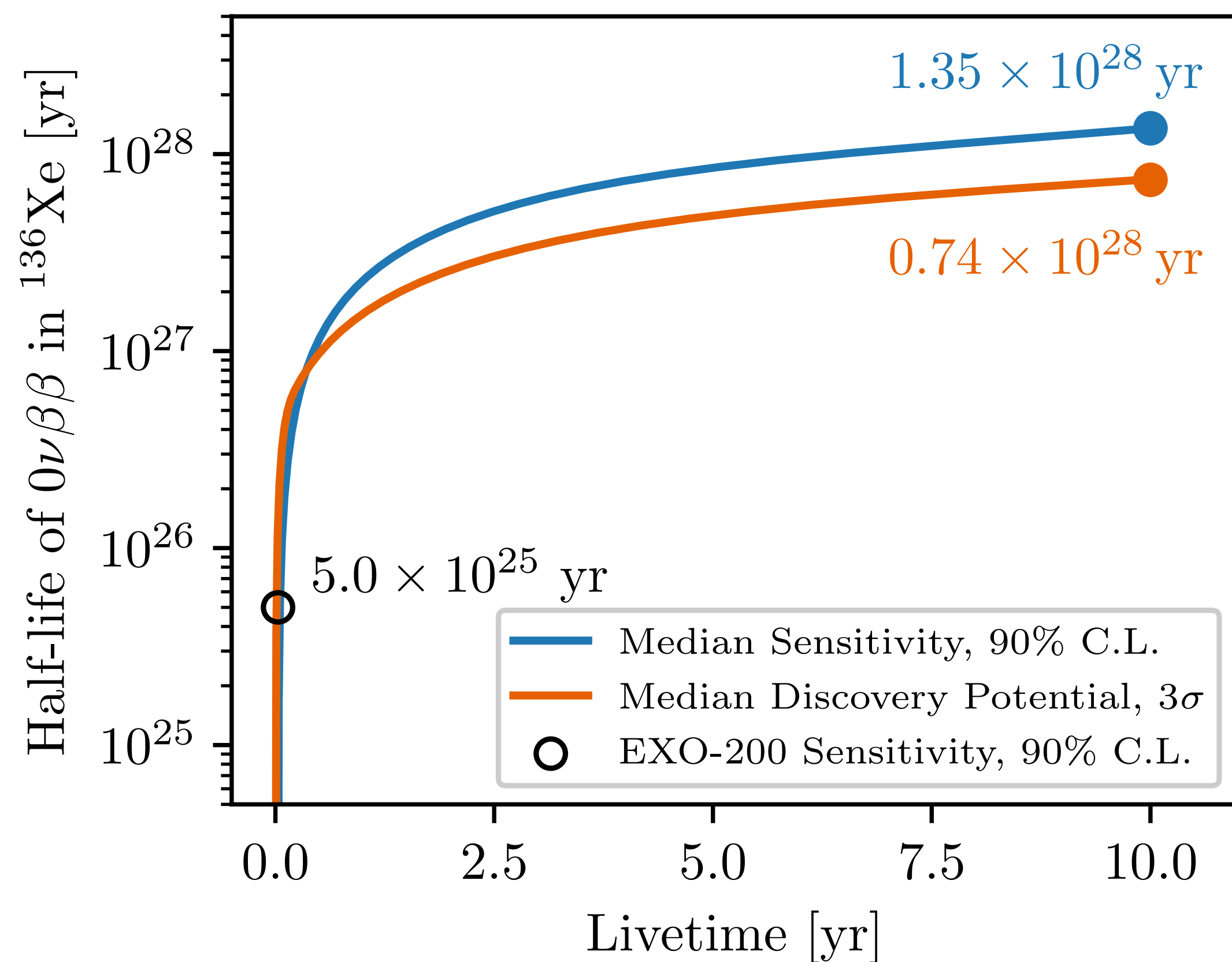
# nEXO's Sensitivity and Discovery Potential





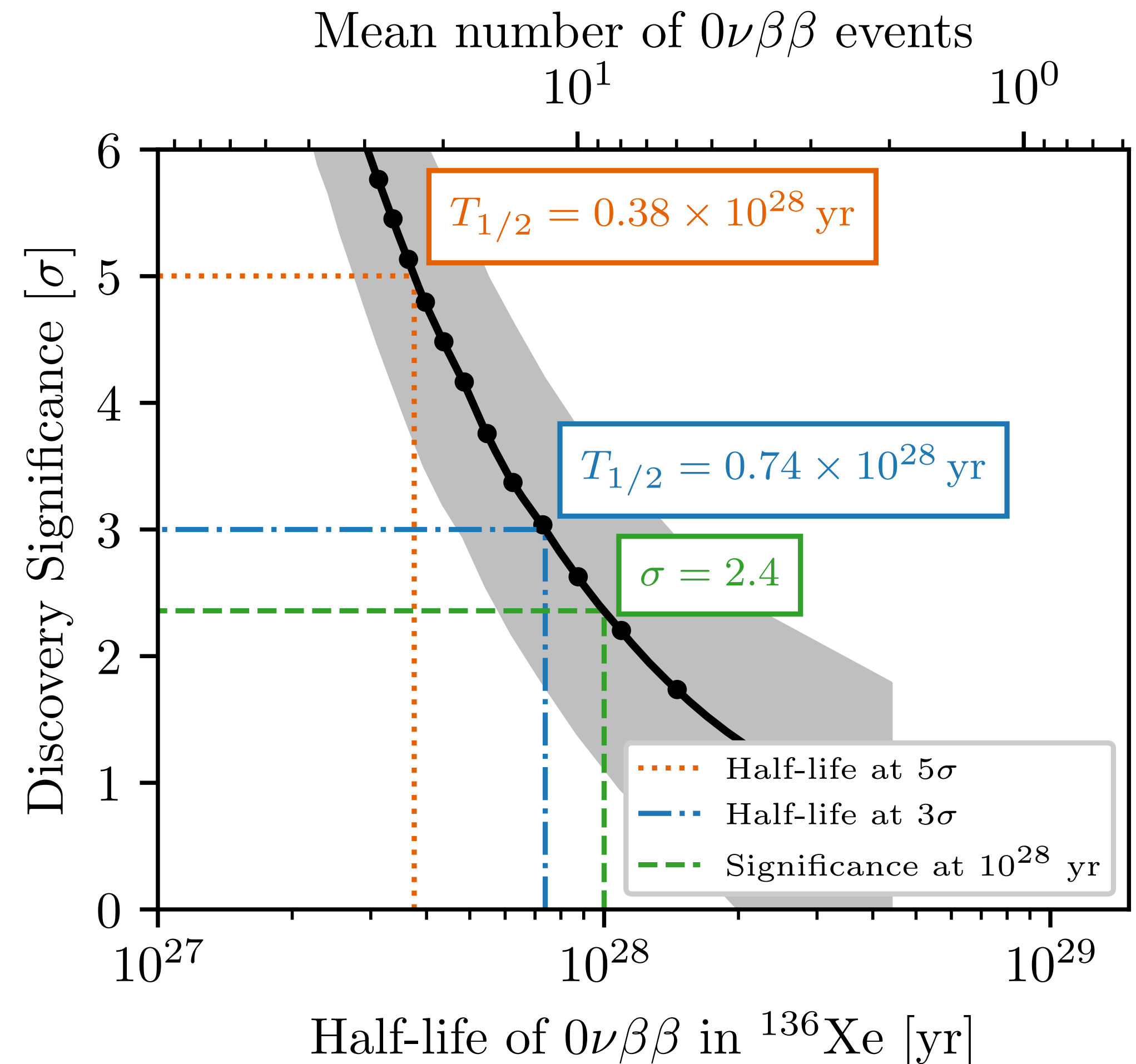
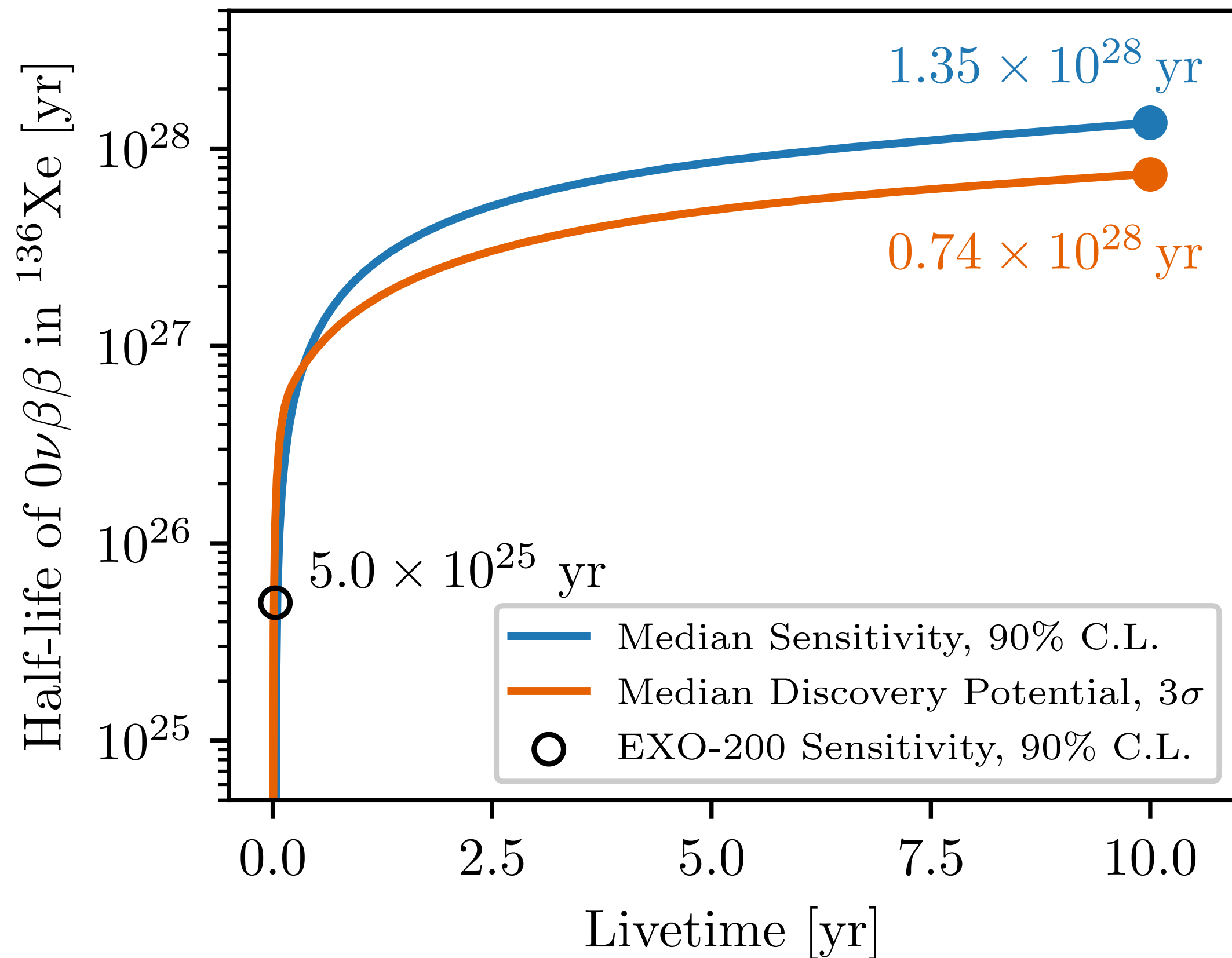
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Increase by **two orders of magnitude** in half-life sensitivity over current generation experiments!



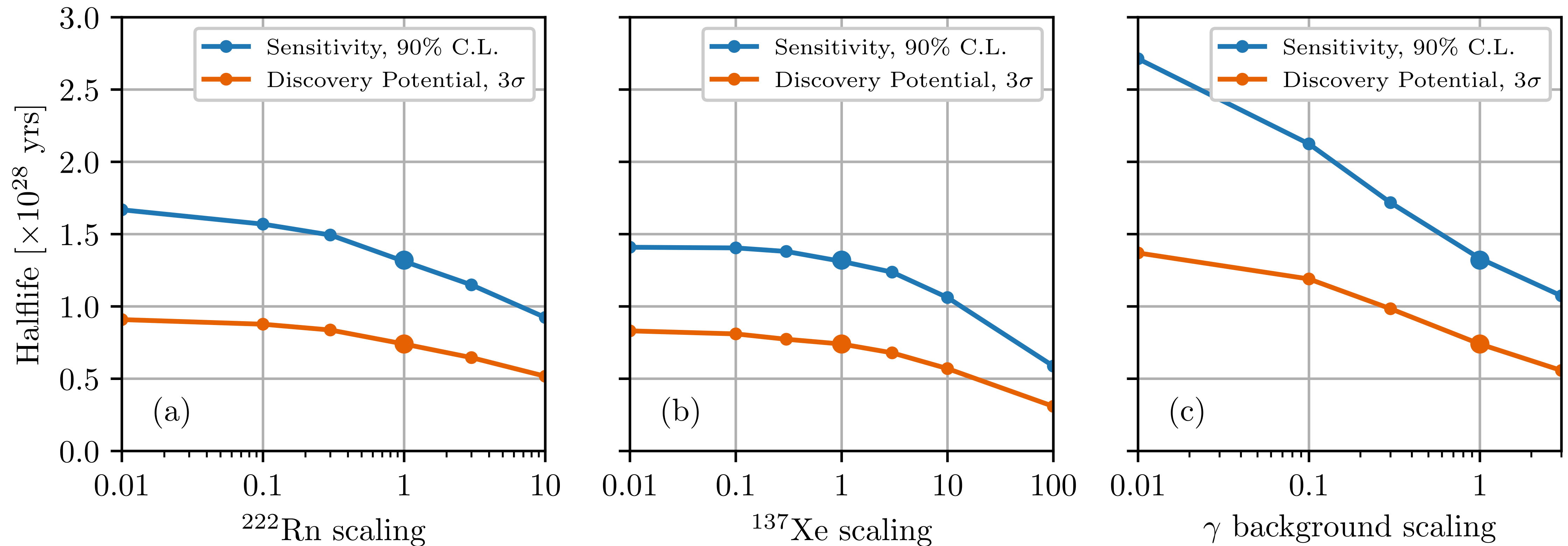
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# Impact of Background Modeling on the Sensitivity

- Credible background estimation, founded on radioassay data and coupled with a detailed MC model
  - Good agreement between prediction and data demonstrated by EXO-200
- nEXO's sensitivity is robust even against misestimates of background components

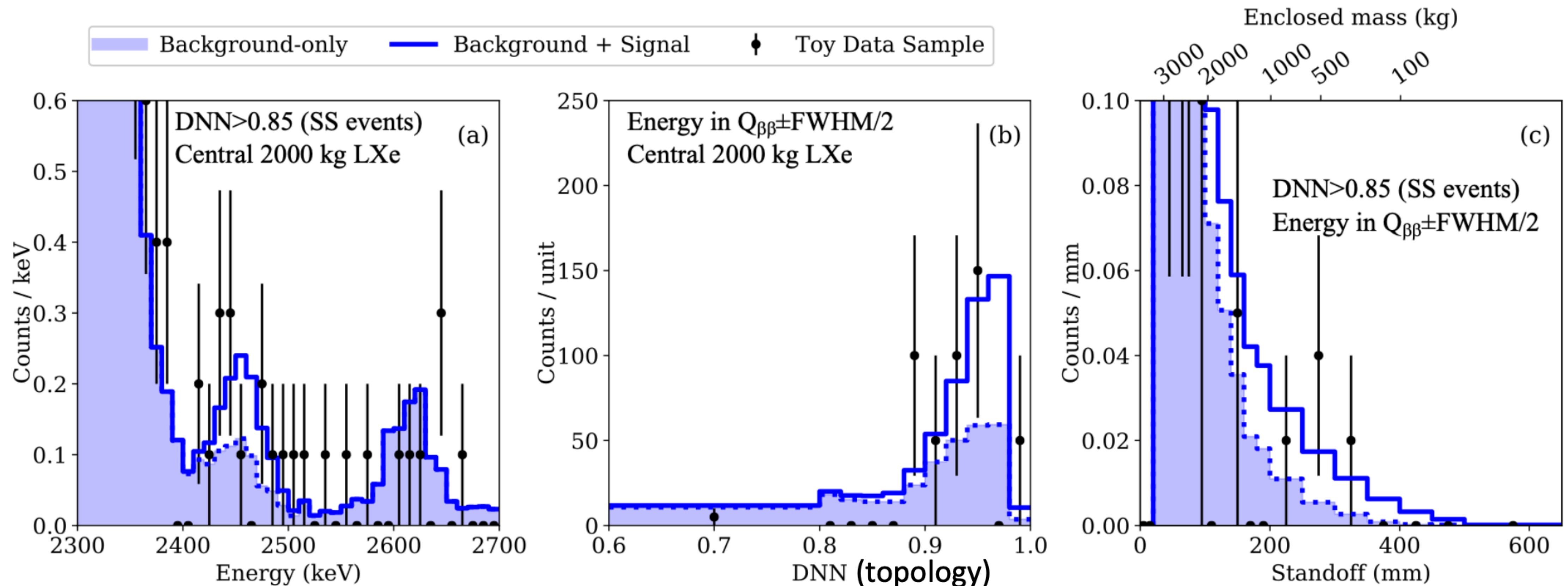


# Potential Discovery of $0\nu\beta\beta$ in nEXO

- Simulated  $0\nu\beta\beta$  signal strength at  $T_{1/2} = 0.74 \times 10^{28}$  yr
- A positive signal in nEXO unlikely to be mistaken for an unknown background!
- Possible to run control experiment with depleted xenon to confirm discovery

Region of Interest:

- $E \in Q_{\beta\beta} \pm \text{FWHM}/2$
- $\text{DNN} > 0.85$  (signal-like events)
- Innermost 2 tonnes of LXe (Standoff > 20 mm)



# Physics Reach of nEXO

$$\frac{1}{T_{1/2}^{0\nu}} = \frac{\langle m_{\beta\beta} \rangle^2}{m_e^2} G^{0\nu} \left| \mathcal{M}^{0\nu} \right|^2$$

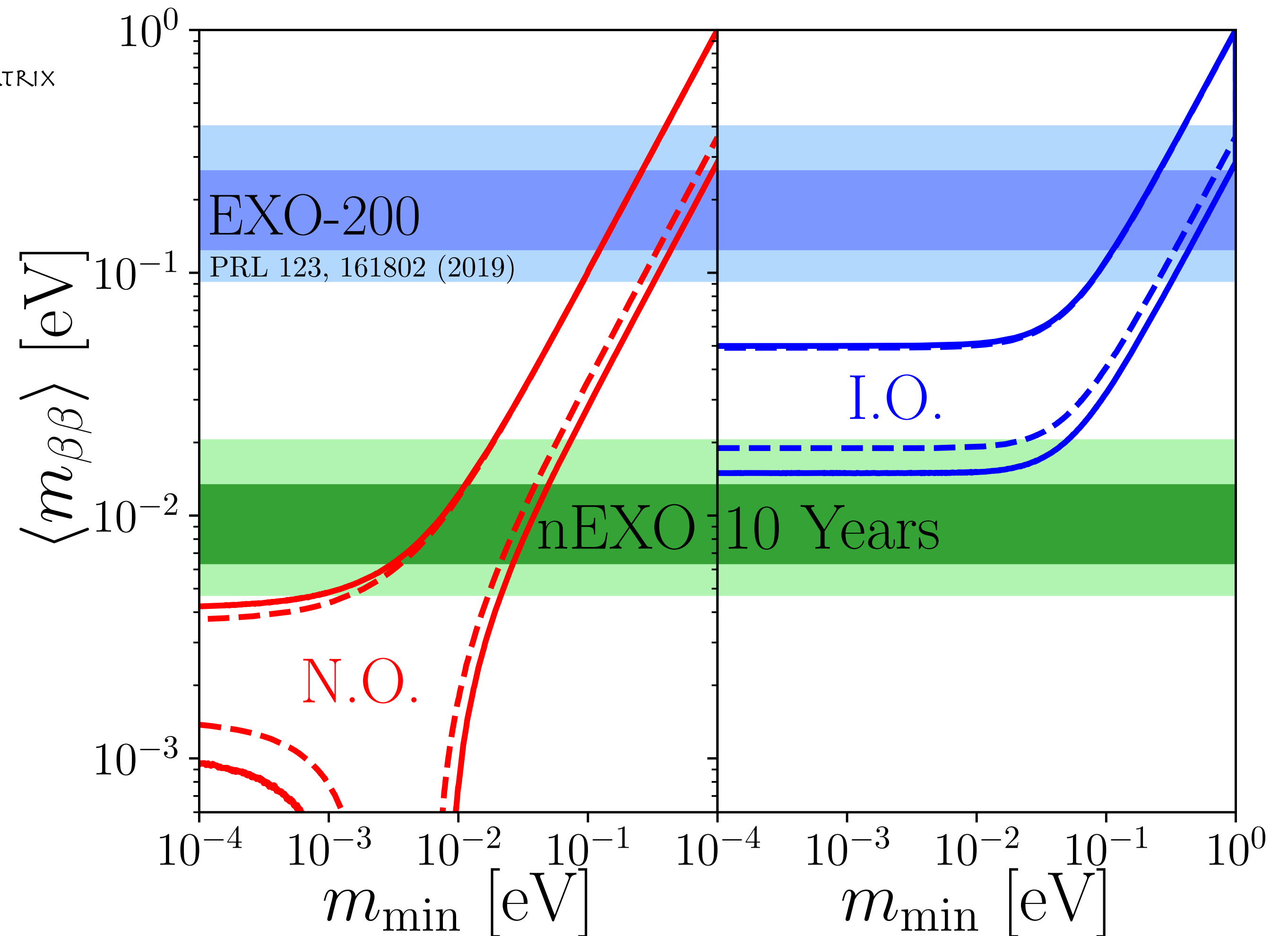
EFFECTIVE MAJORANA NEUTRINO MASS:  $\langle m_{\beta\beta} \rangle^2$   
 PHASE SPACE FACTOR:  $G^{0\nu}$   
 NUCLEAR MATRIX ELEMENT:  $\left| \mathcal{M}^{0\nu} \right|^2$

- Experimental limit on  $T_{1/2}^{0\nu}$  translated into limit on  $m_{\beta\beta}$

- $T_{1/2}^{0\nu} > 1.35 \times 10^{28} \text{ yr}$

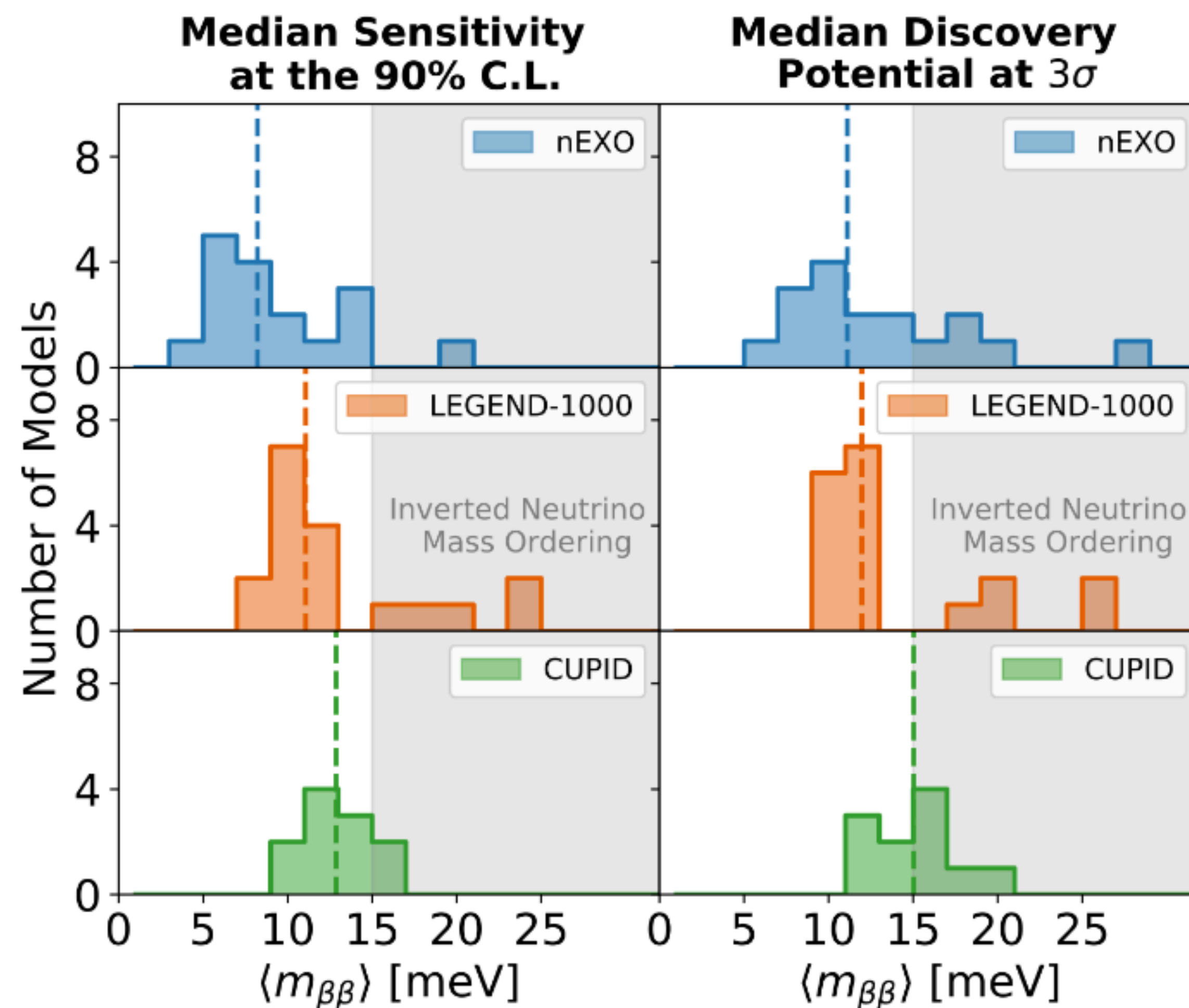
- $\langle m_{\beta\beta} \rangle < (4.7 - 20.4) \text{ meV}$

- nEXO will completely cover the Inverted Neutrino Mass Ordering



# Tonne-Scale $0\nu\beta\beta$

- Agnostic selection of NMEs over last 20 years
- Great promise for next-generation experiments to make a discovery
- Multi-isotope confirmation of  $0\nu\beta\beta$  needed to understand physics mediating this decay
- Additional mechanism beyond light Majorana neutrino exchange

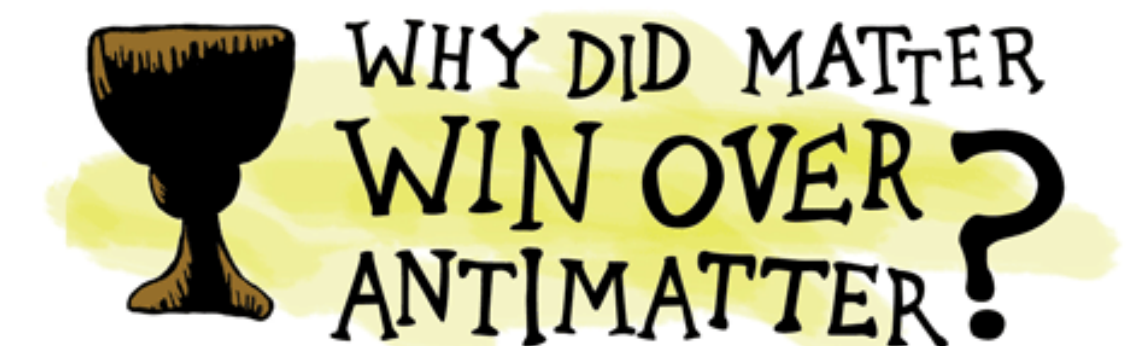
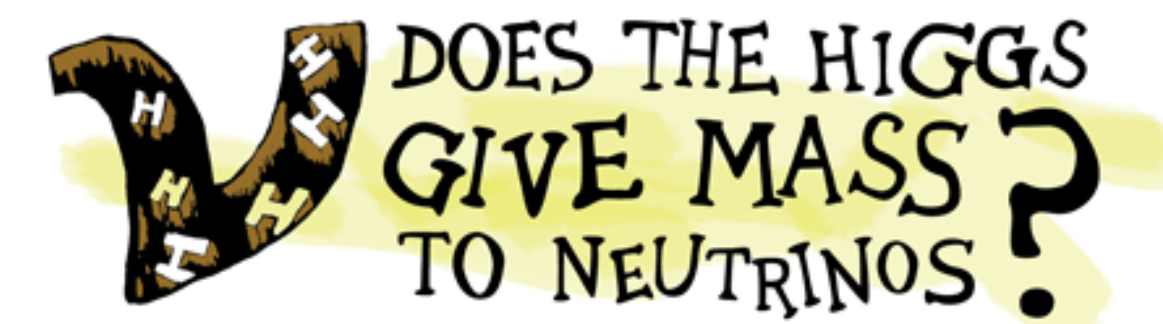


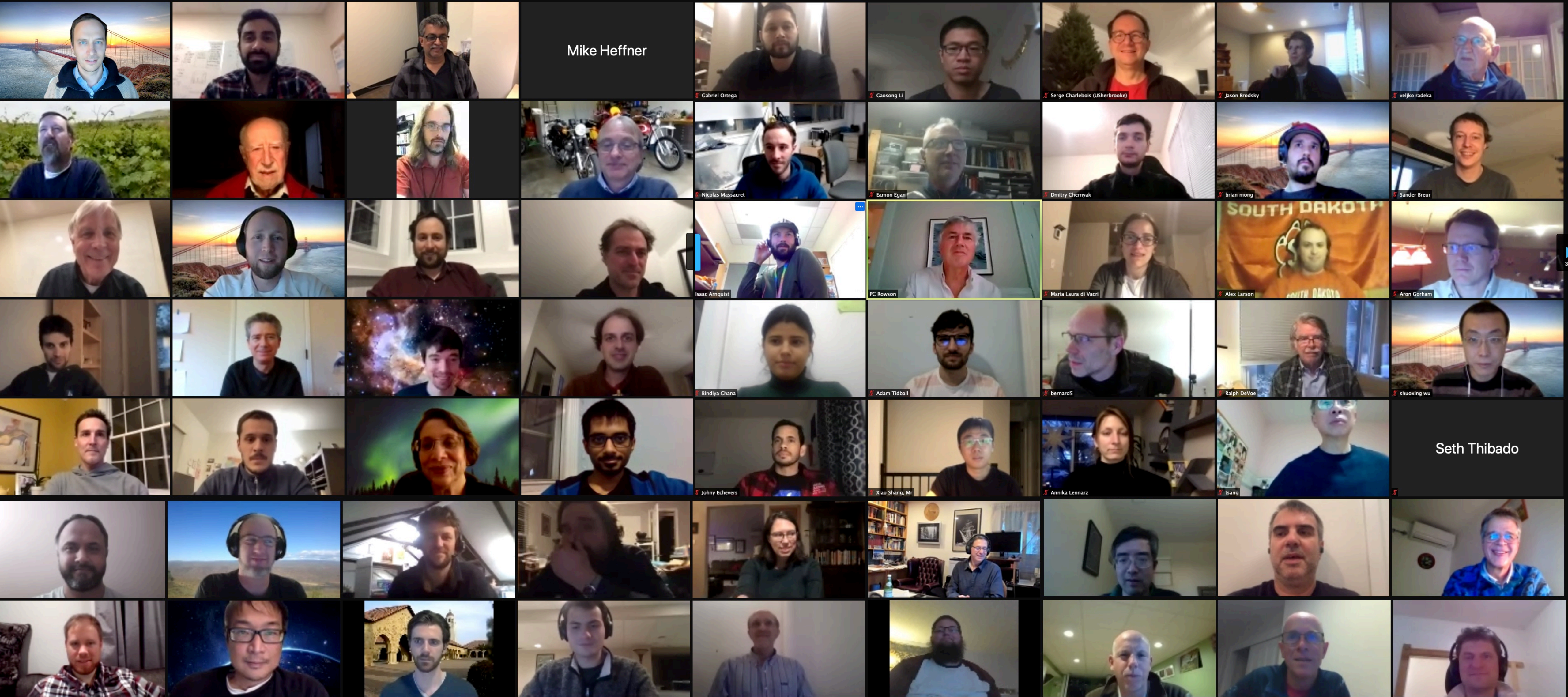
NMEs values don't follow a statistical distribution. Only a single value is true. However, calculations are difficult and have large uncertainties

# Summary

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- Searches for  $0\nu\beta\beta$  are a powerful tool to probe physics beyond the Standard Model
- nEXO utilizes a tonne-scale LXe TPC to search for  $0\nu\beta\beta$  in  $^{136}\text{Xe}$ 
  - Low intrinsic and well-understood background
  - Good energy resolution
  - Powerful background discrimination using multi-variate analysis
  - Capability for running a control experiment with natXe
- This is a very exciting time for  $0\nu\beta\beta$  and a discovery might be just around the corner!!





Mike Heffner

Seth Thibado



Collaboration Meeting Dec 2020



BACKUP SLIDES

# Improved Realism of Detector Simulation and Reconstruction

## Light Simulation

- Data-driven optical properties for SiPMs
- Conservative values for reflectivities of passive TPC components
- Detailed understanding of photon propagation

## Charge Simulation

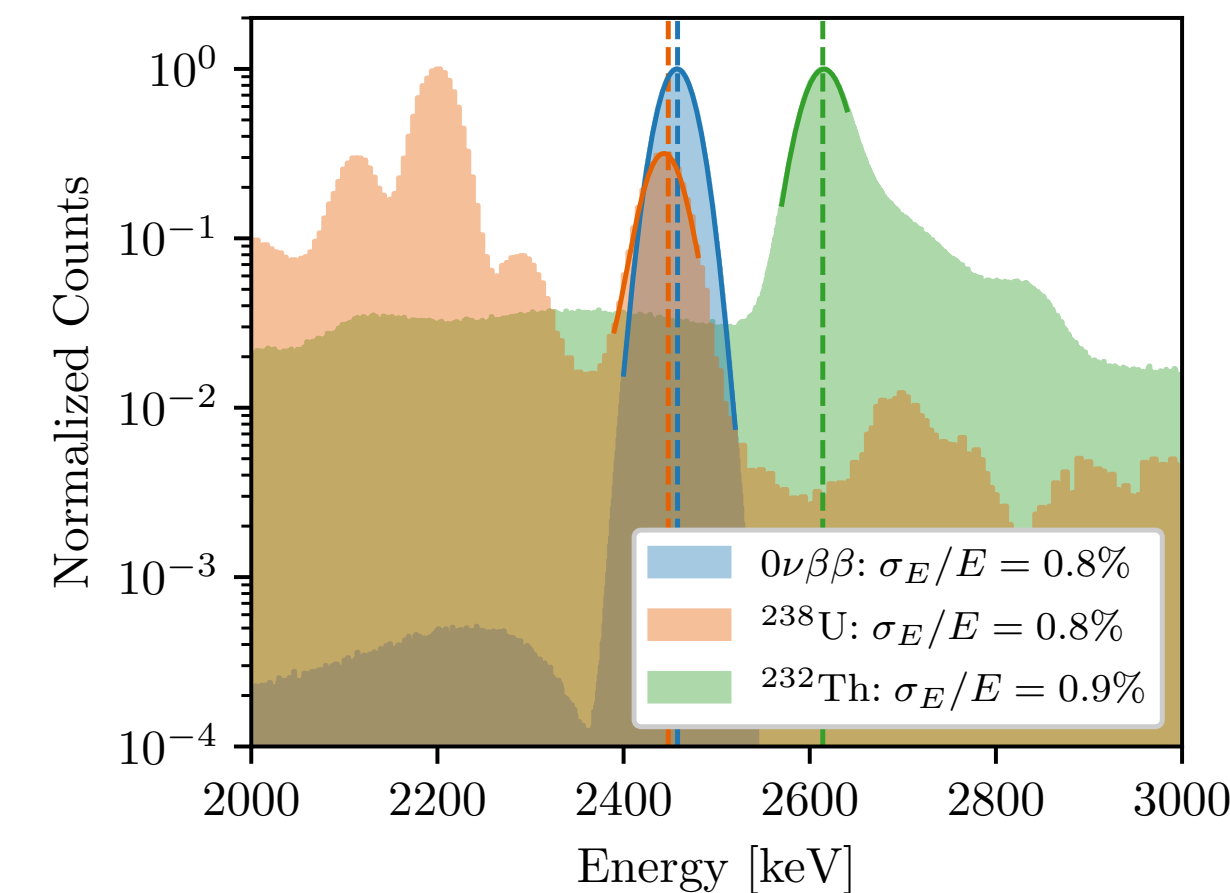
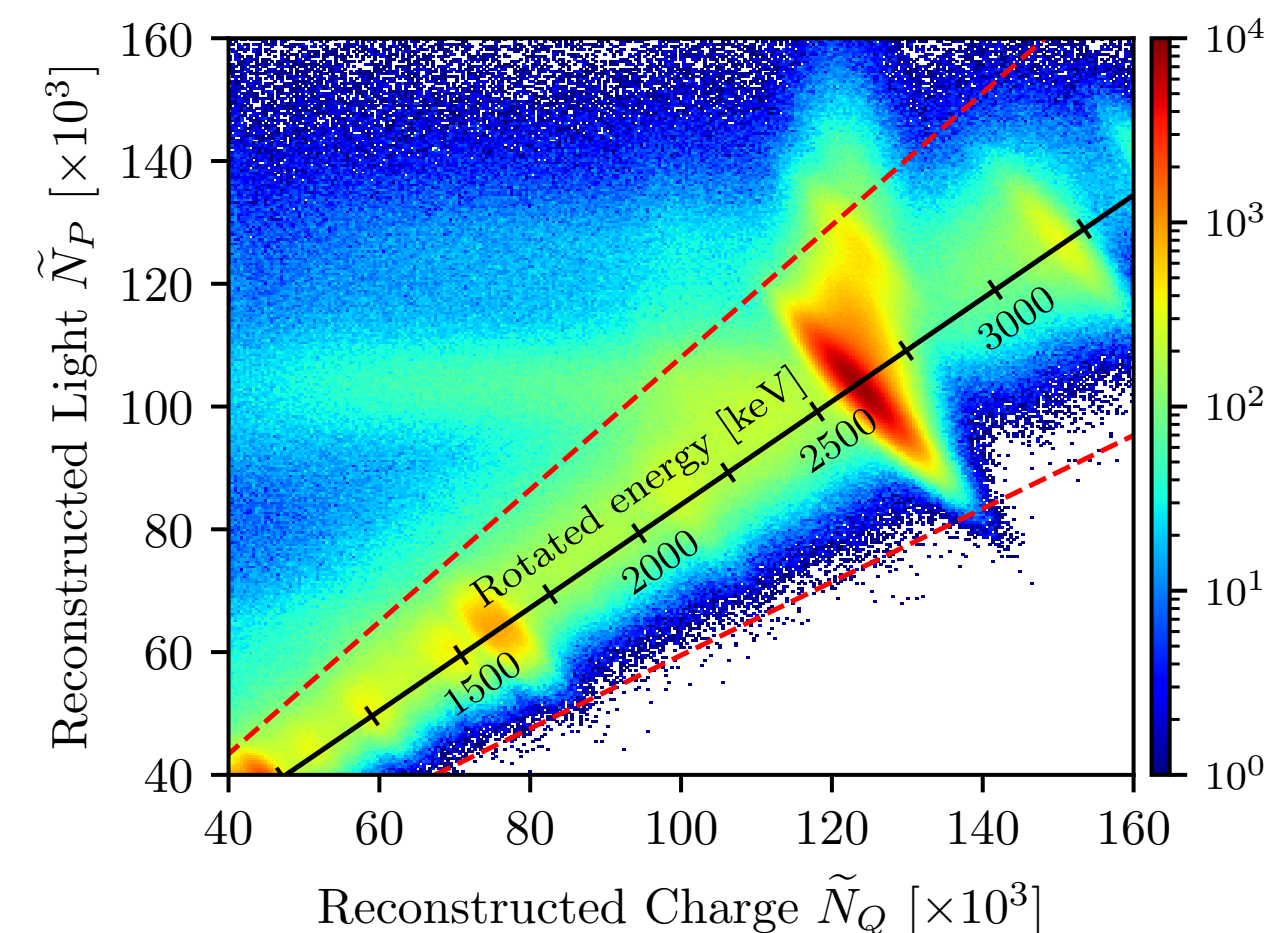
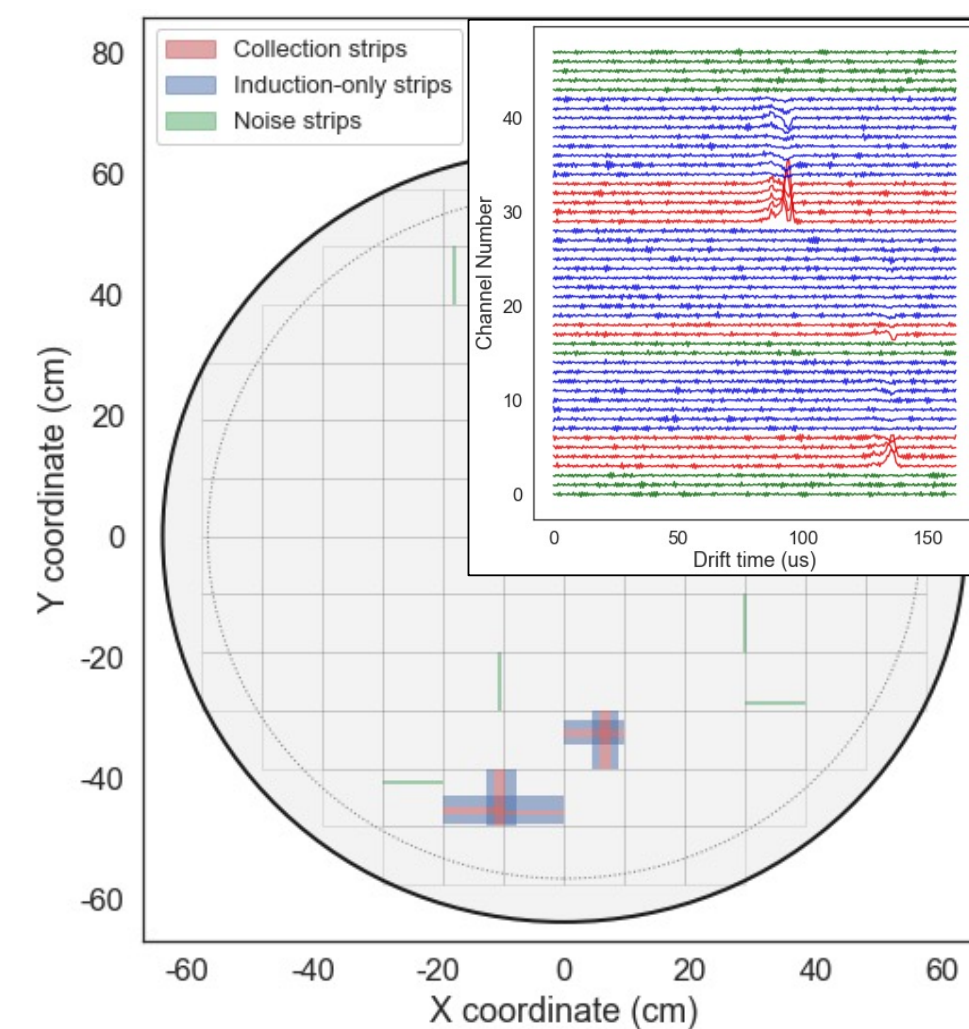
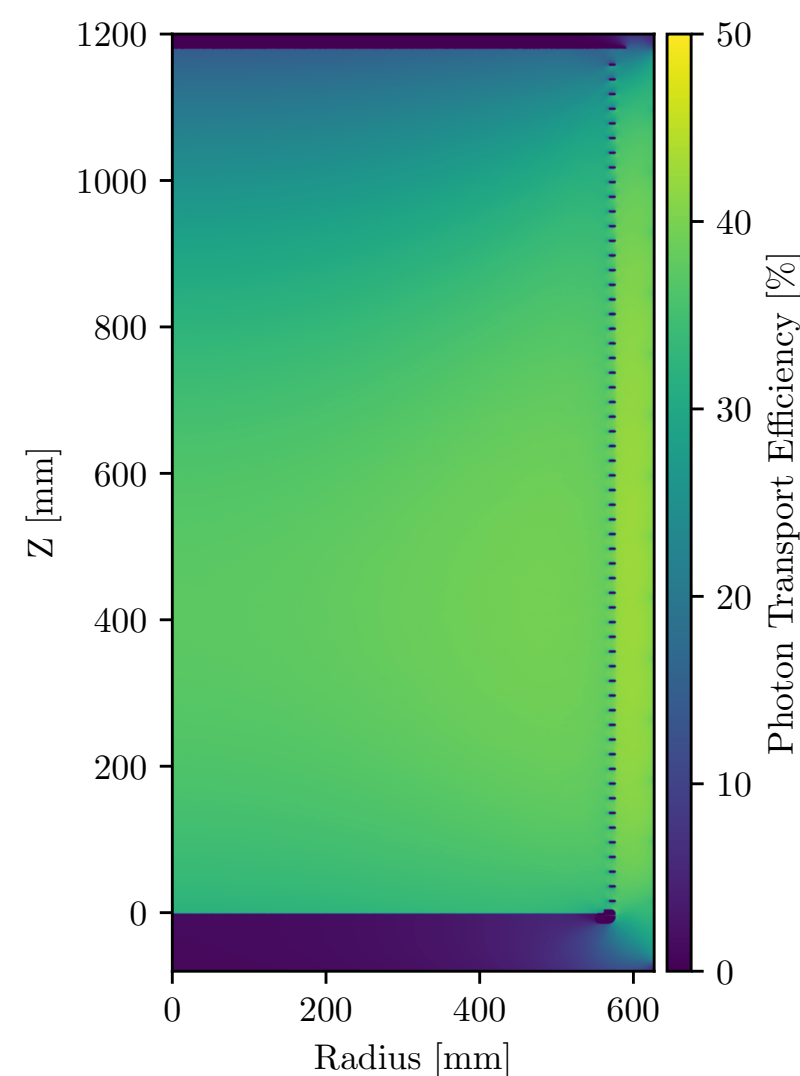
- Complete charge propagation through TPC
- Inclusion of diffusion & electron lifetime effects
- Realistic ASIC electronics noise added to waveforms

## Rotated Energy

- Anti-correlation between light and charge exploited by combining both for energy estimate
- LXe “skin” effects due to open-field cage design is well understood ([arXiv:2009.10231v2](https://arxiv.org/abs/2009.10231v2))

## Energy Resolution

- Not dominated by either light or charge channel
- Estimated to be  $\sigma_Q/Q = 0.8\%$
- In good agreement with semi-empirical resolution model (validated by EXO-200)



# Light and Charge Detection

ENERGY RESOLUTION IN NEXO

Fano noise

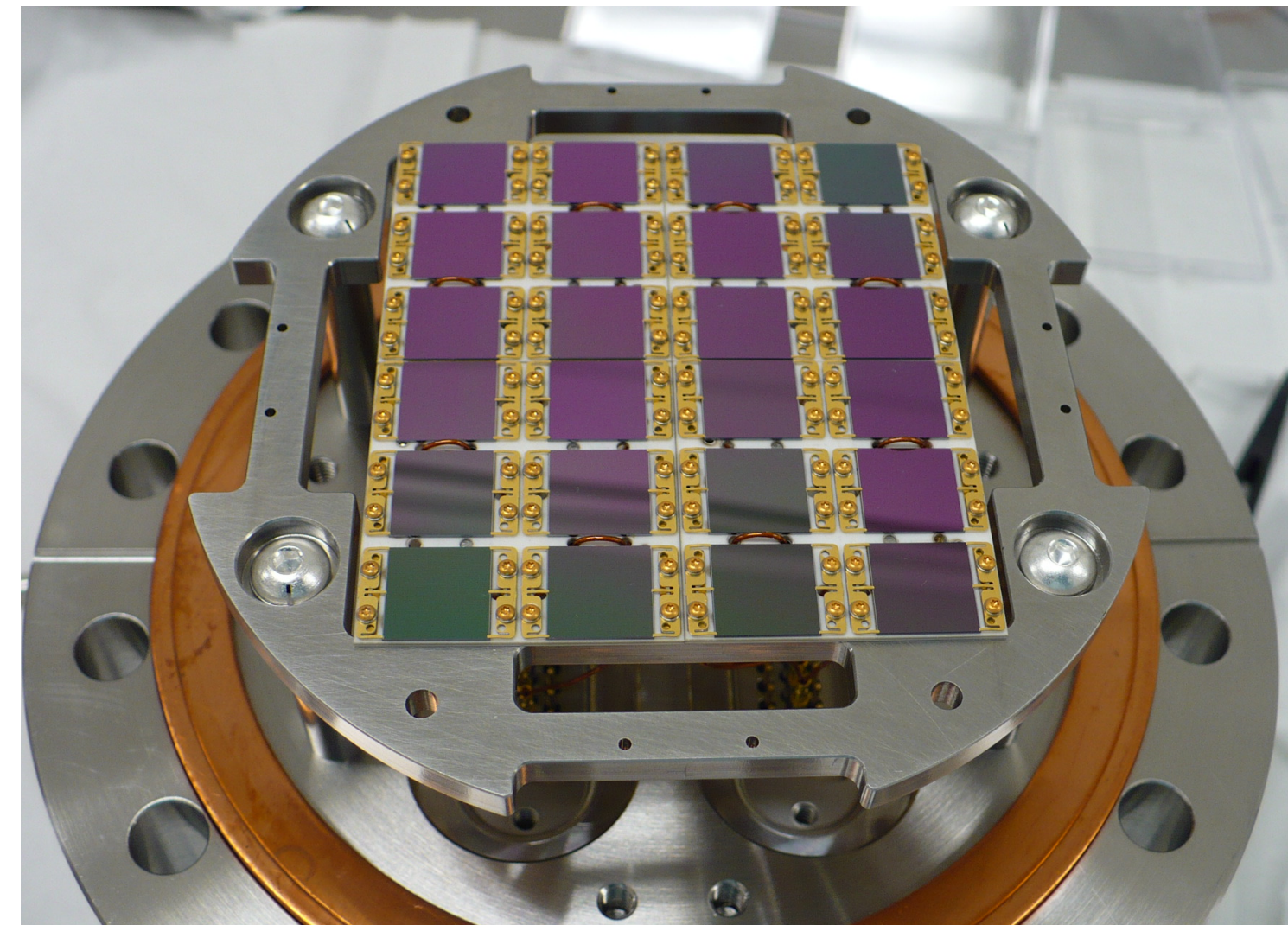
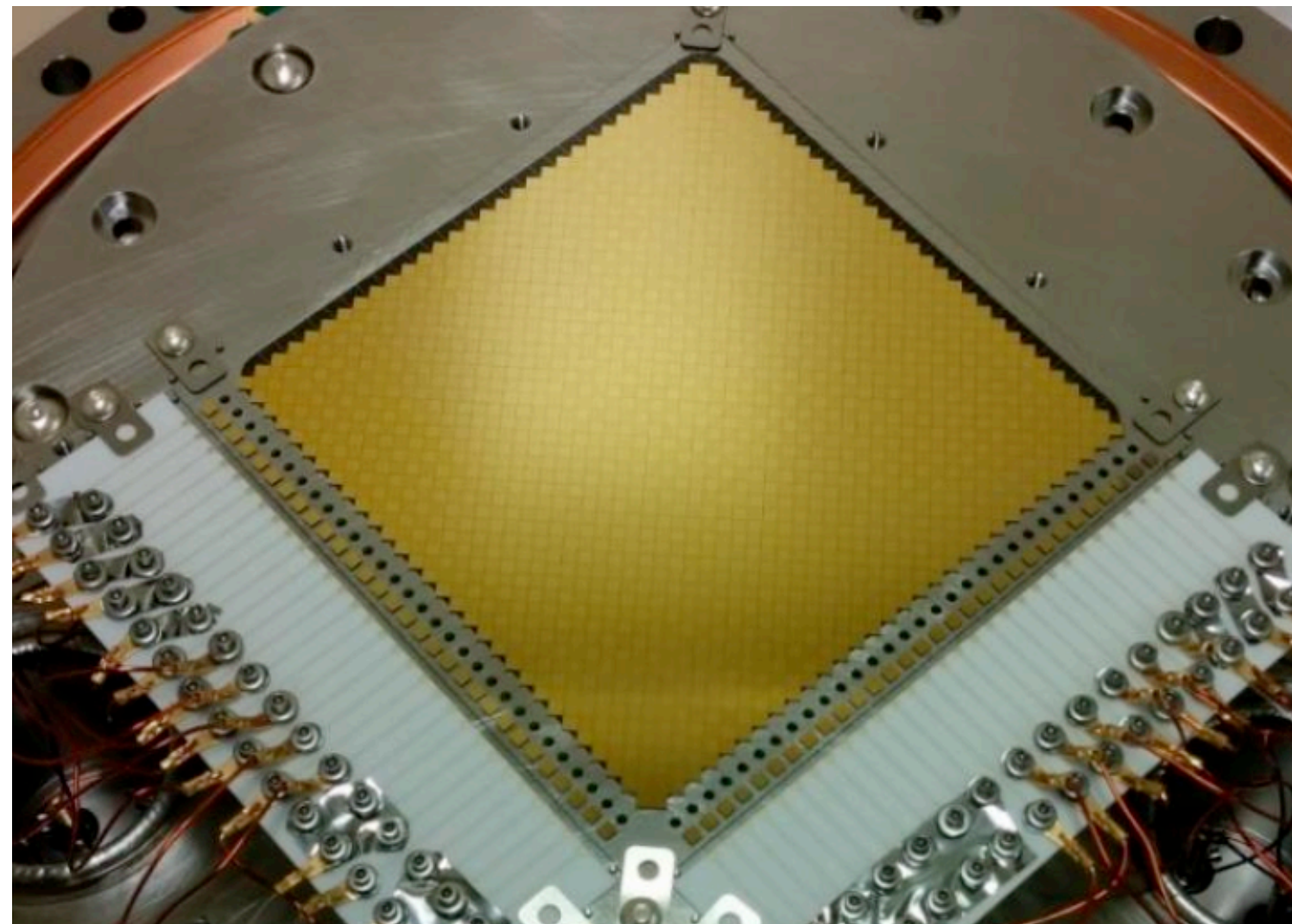
Charge Channel

- Charge collection
- e- lifetime systematic
- charge electronics

Light Channel

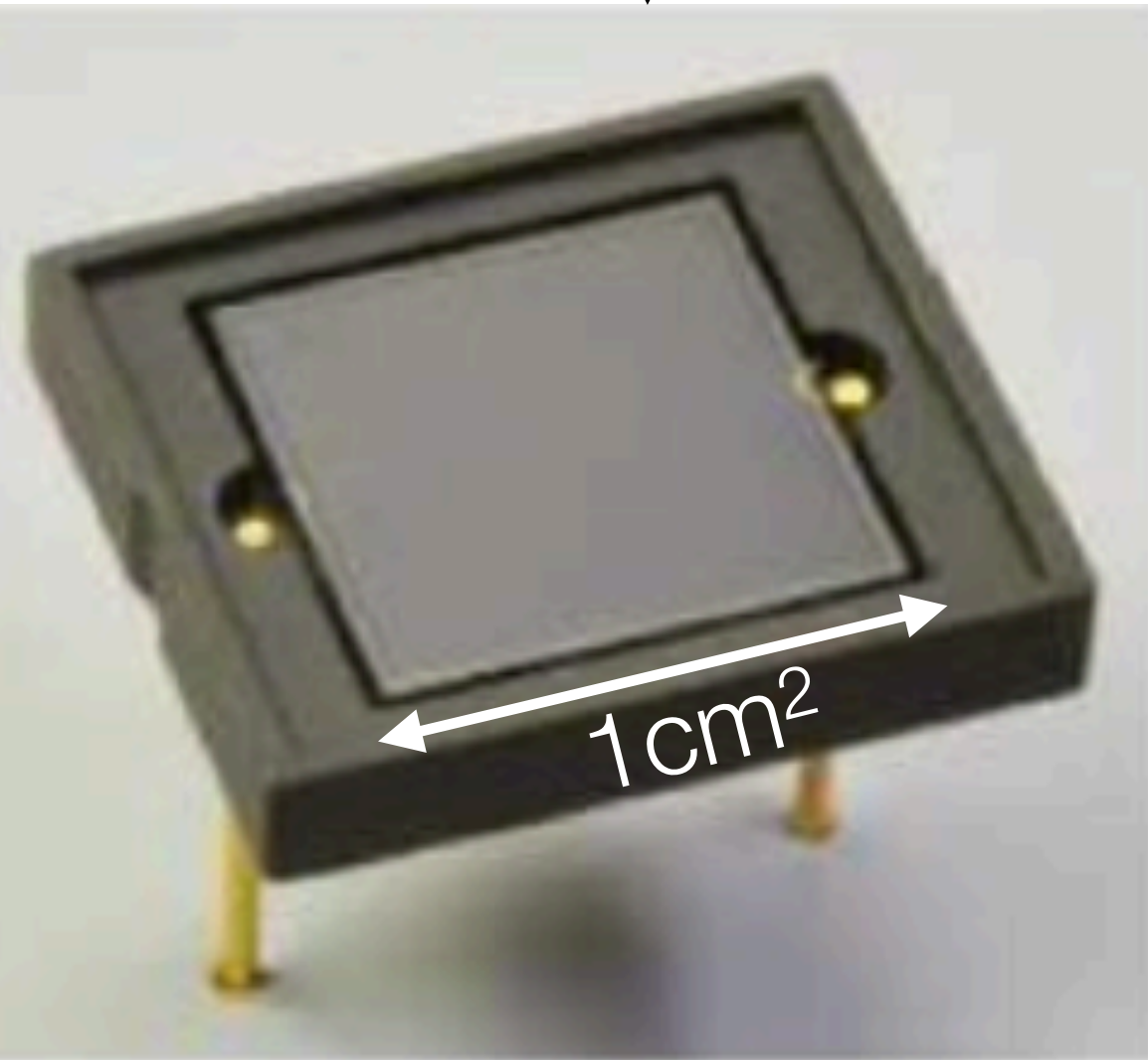
- Light collection
- Correlated avalanches
- Dark current and gain
- Lightmap Error

$$\frac{\sigma_E}{E} = \frac{\sqrt{F \cdot (Q + S) + \frac{Q(1 - e^{-t/\tau})}{e^{-t/\tau}} + \frac{Q^2}{(e^{-t/\tau})^2} \left(\frac{t}{\tau}\right)^2 \sigma_\tau^2 + \frac{\sigma_{Q,\text{noise}}^2}{(e^{-t/\tau})^2}}}{Q + S} + \frac{S}{\epsilon} \left[ (1 - \epsilon) + \frac{\Lambda}{(1 + \Lambda)^2} \right] + \frac{DC + \eta_{\text{noise}}}{\epsilon^2(1 + \Lambda)^2} + S^2 \sigma_{lm}^2$$

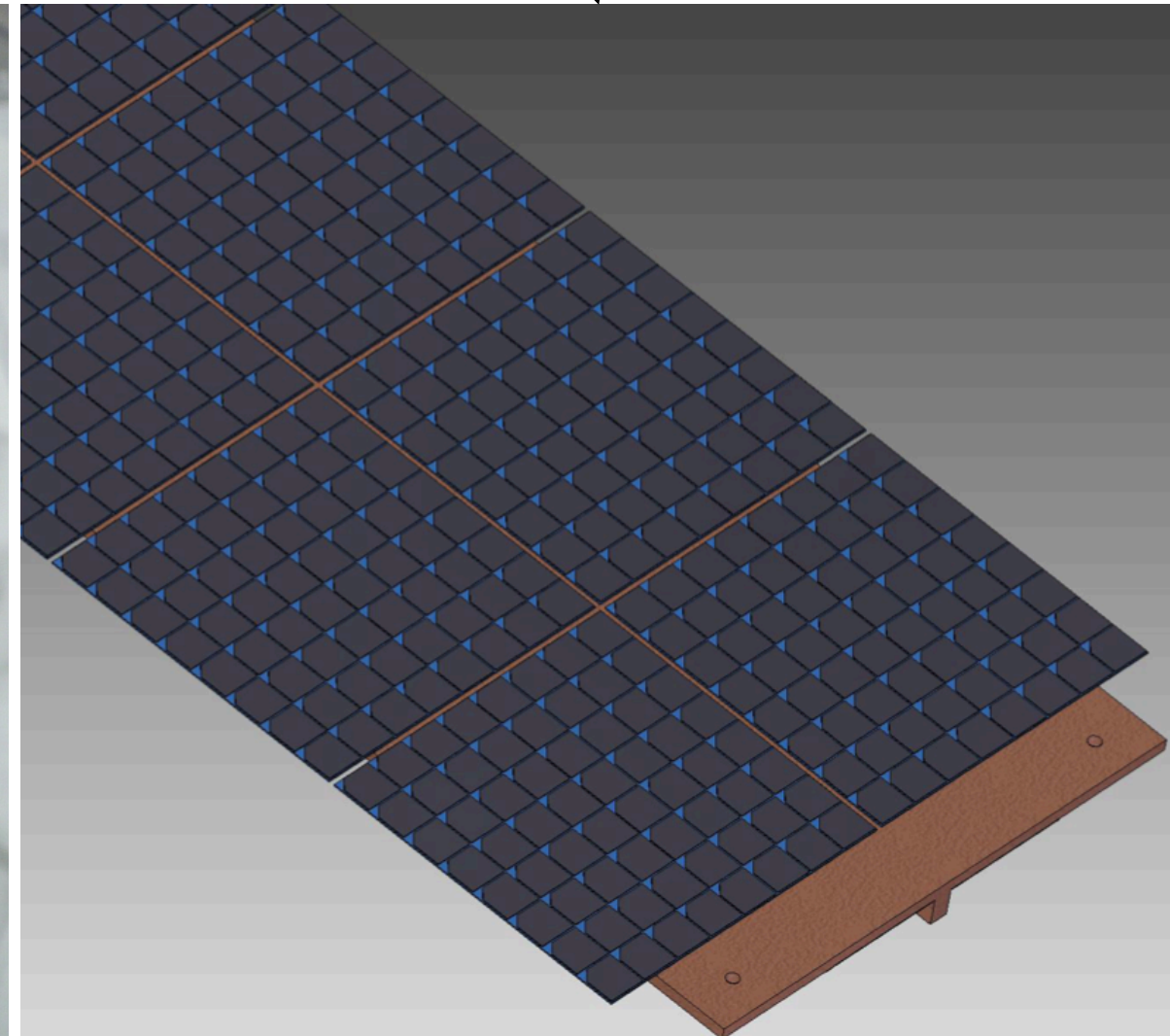


# Light Detection in nEXO

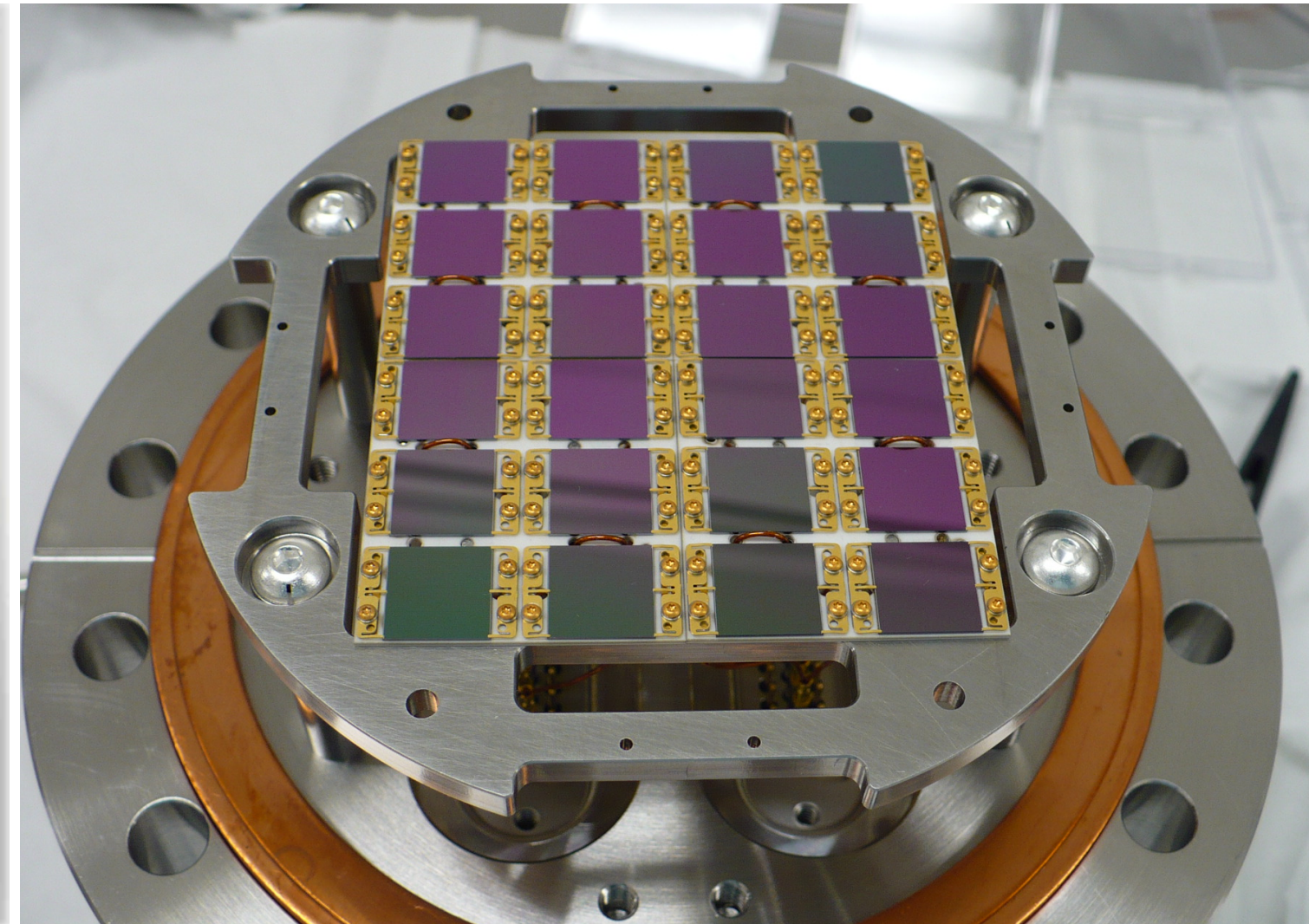
SINGLE SIPM



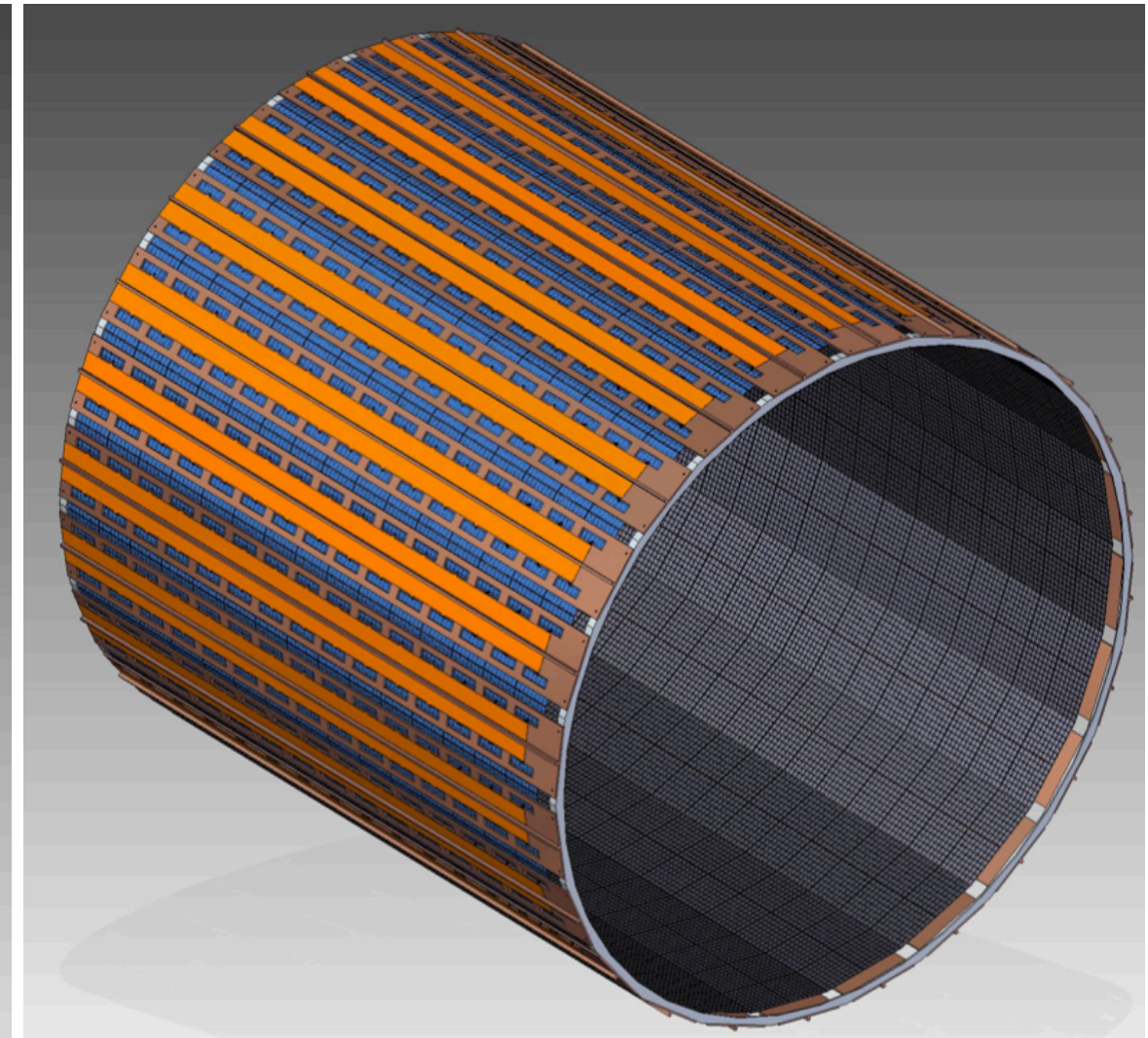
SIPM STAVE



SIPM ARRAY



FULL LIGHT DETECTION SYSTEM



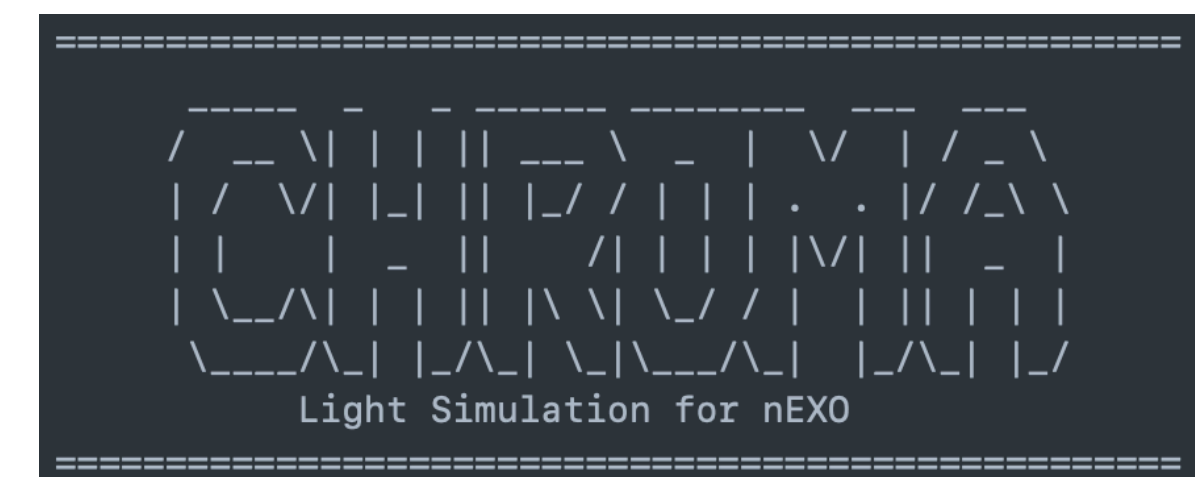
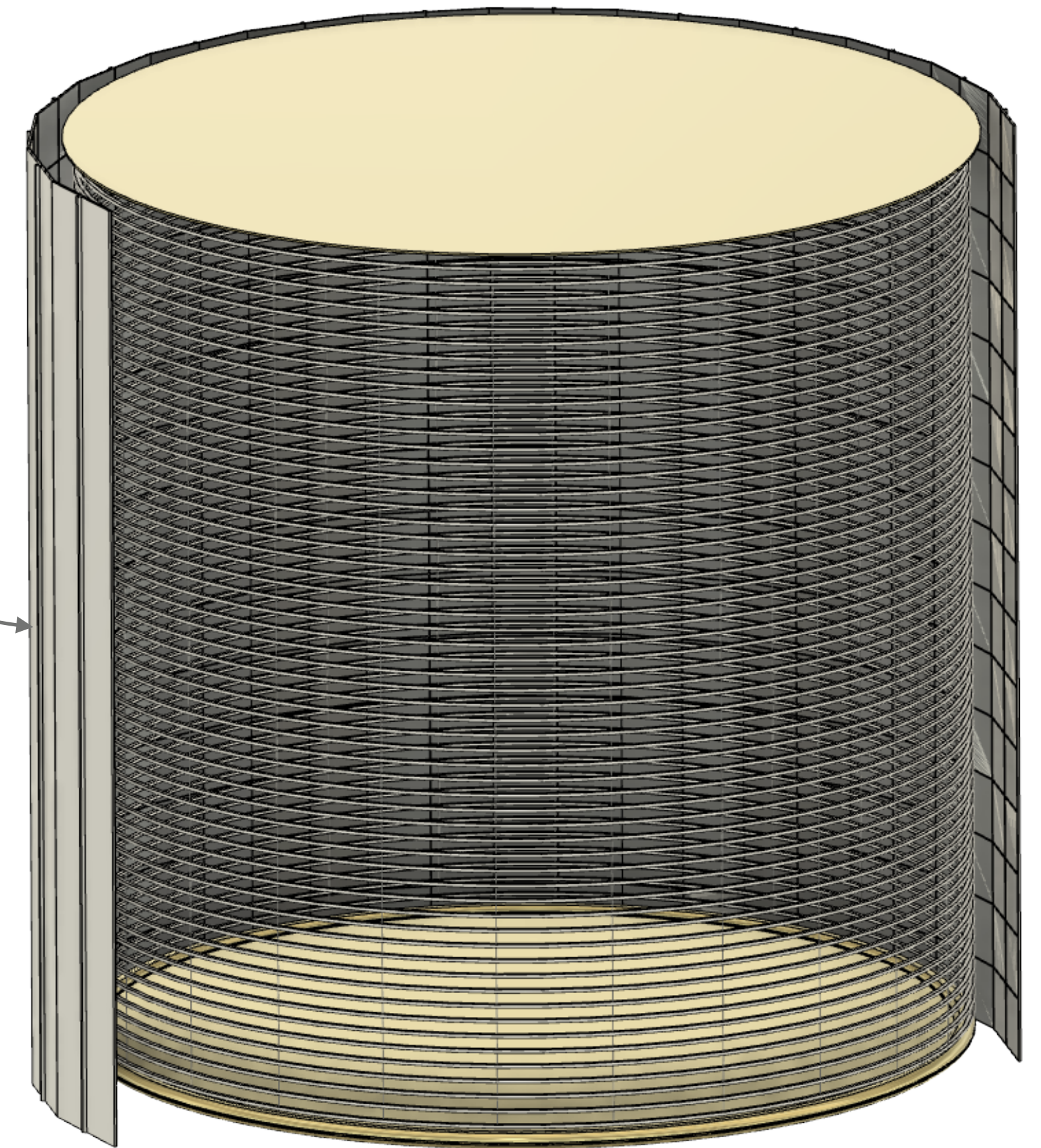
# Light Transport Simulations

- Crucial to optimize Photon Transport Efficiency for increased light collection

$$\epsilon = \text{PDE} \cdot \text{PTE}$$

- Highly dependent on reflectivity of TPC components
- Developed new light simulation of nEXO with GPU-based Chroma software  
<https://github.com/nEXO-collaboration/chroma>
- > 300x faster
- More detailed geometry

CAD MODEL  
OF NEXO



# Lightmap Simulation

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nEXO Sensitivity and Discovery Potential: [arXiv:2106.16243](https://arxiv.org/abs/2106.16243)

# Lightmap Simulation

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nEXO Sensitivity and Discovery Potential: [arXiv:2106.16243](https://arxiv.org/abs/2106.16243)

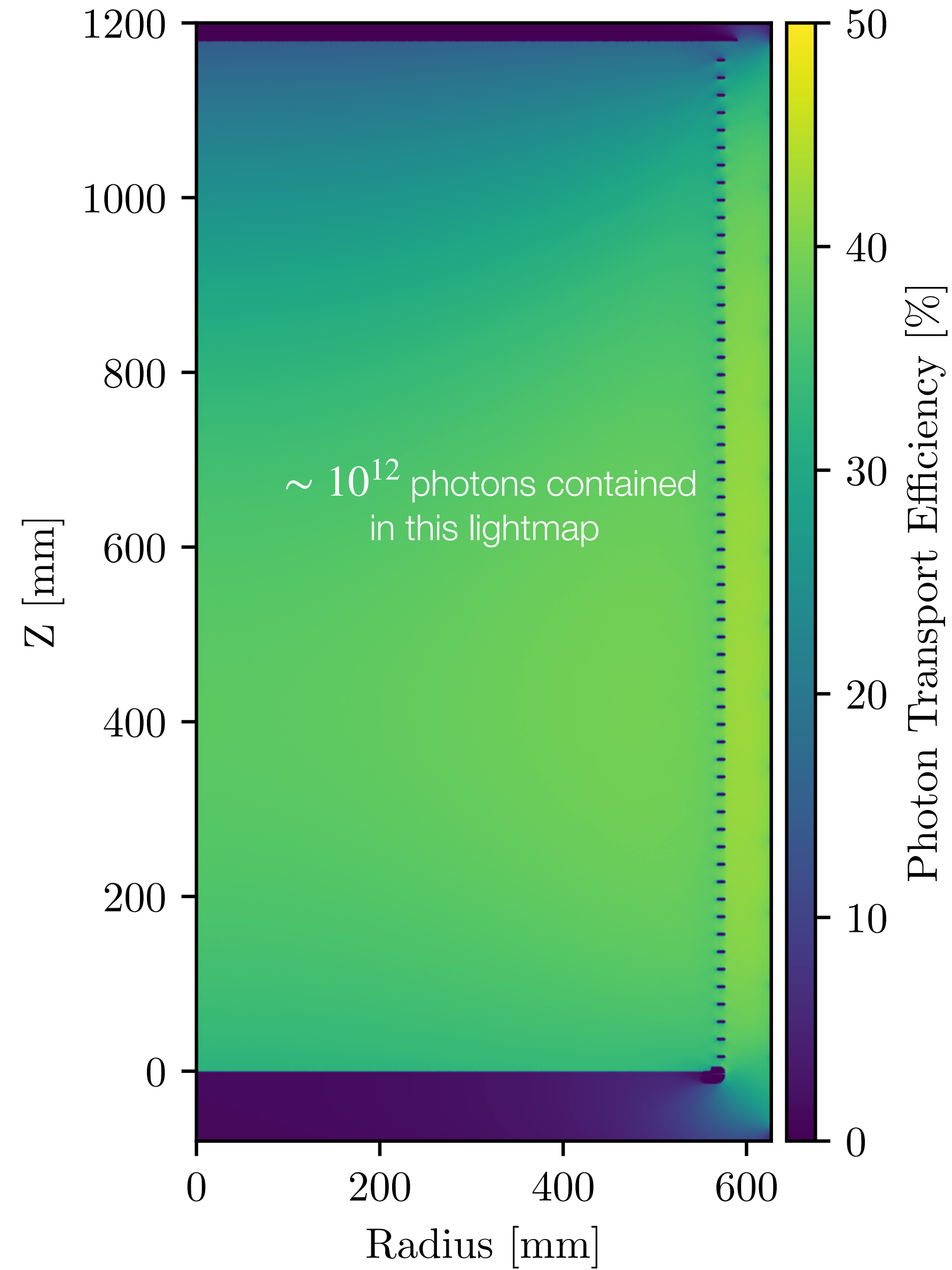
- Varied optical parameters and evaluated systematic error

# Lightmap Simulation

nEXO Sensitivity and Discovery Potential: [arXiv:2106.16243](https://arxiv.org/abs/2106.16243)

- Varied optical parameters and evaluated systematic error
- Estimated PTE combined with measured PDE results in

$$\epsilon = \text{PDE} \cdot \text{PTE} \approx 6.7 \%$$



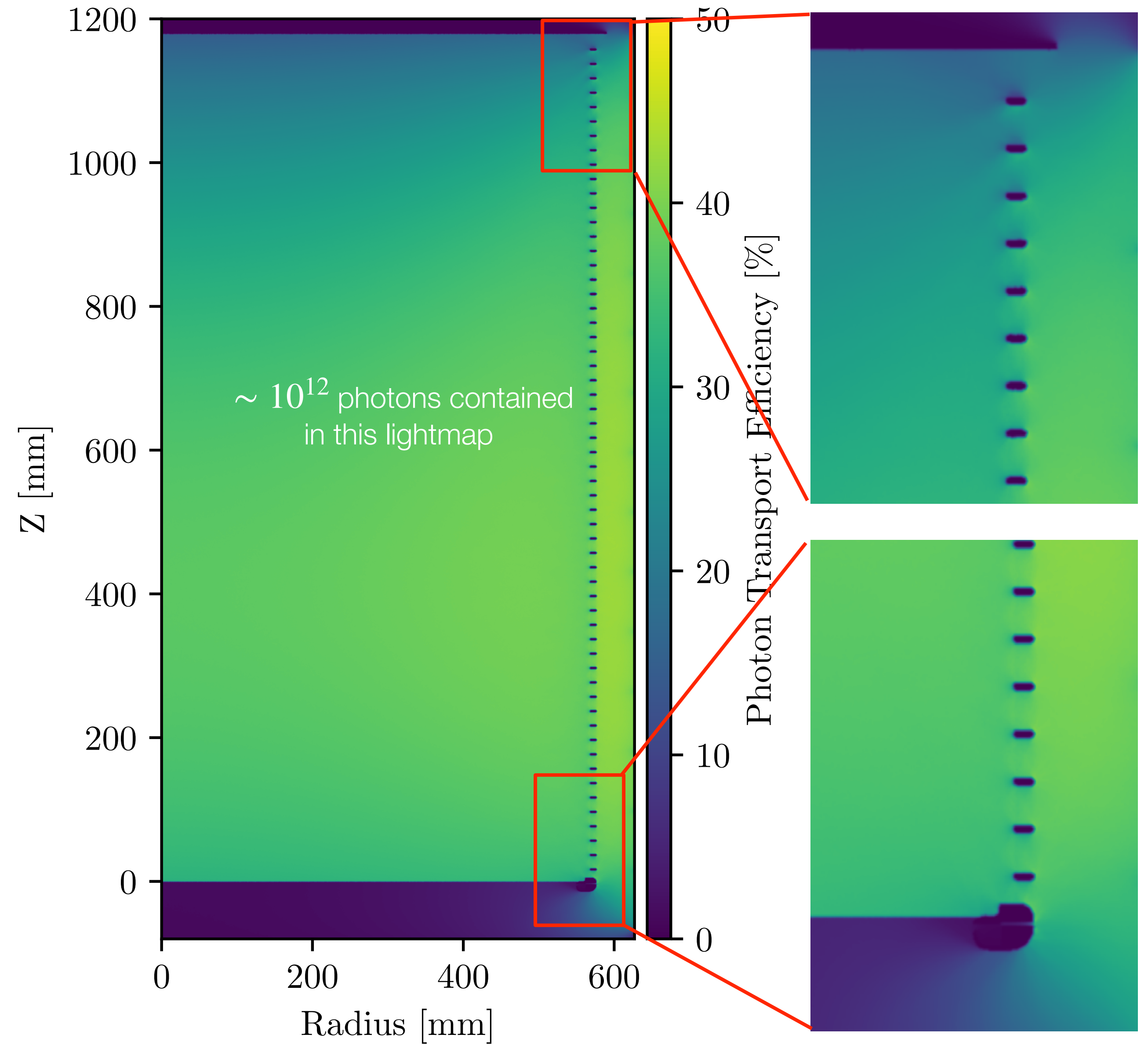


# Lightmap Simulation

nEXO Sensitivity and Discovery Potential: [arXiv:2106.16243](https://arxiv.org/abs/2106.16243)

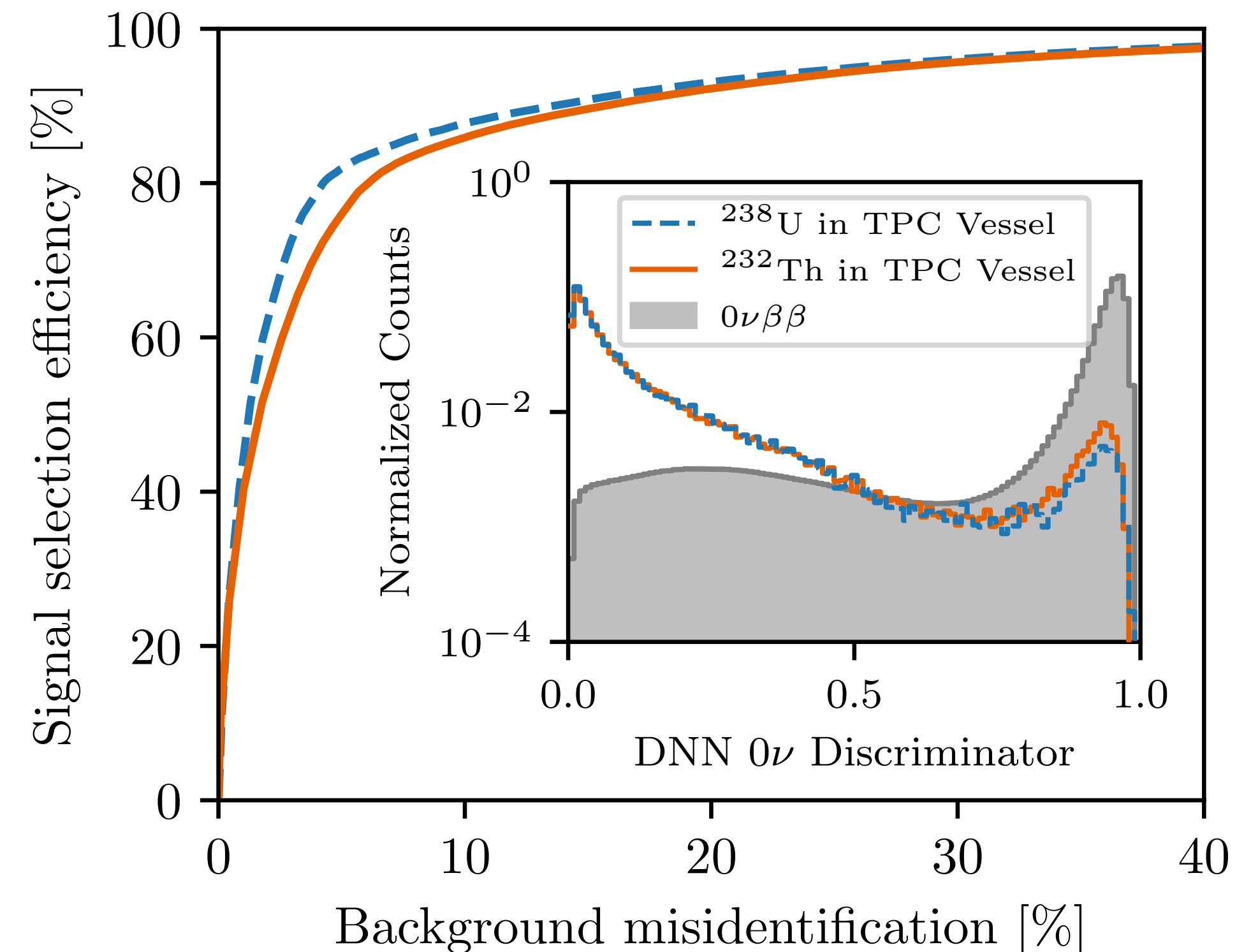
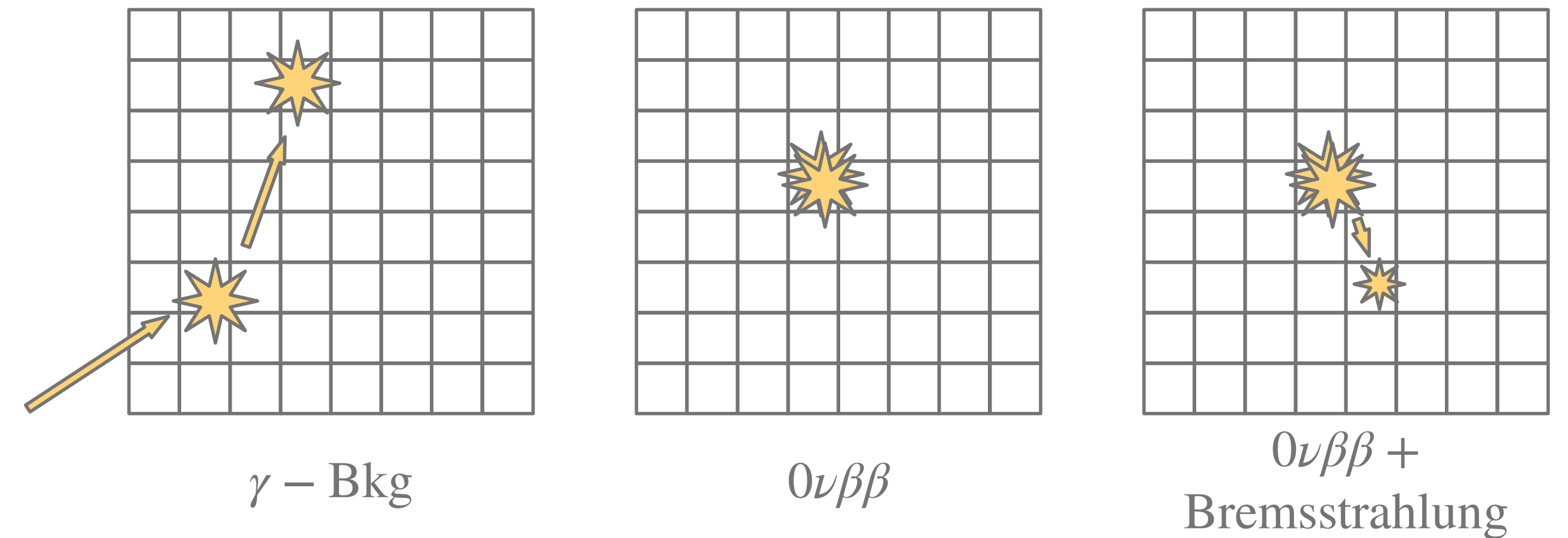
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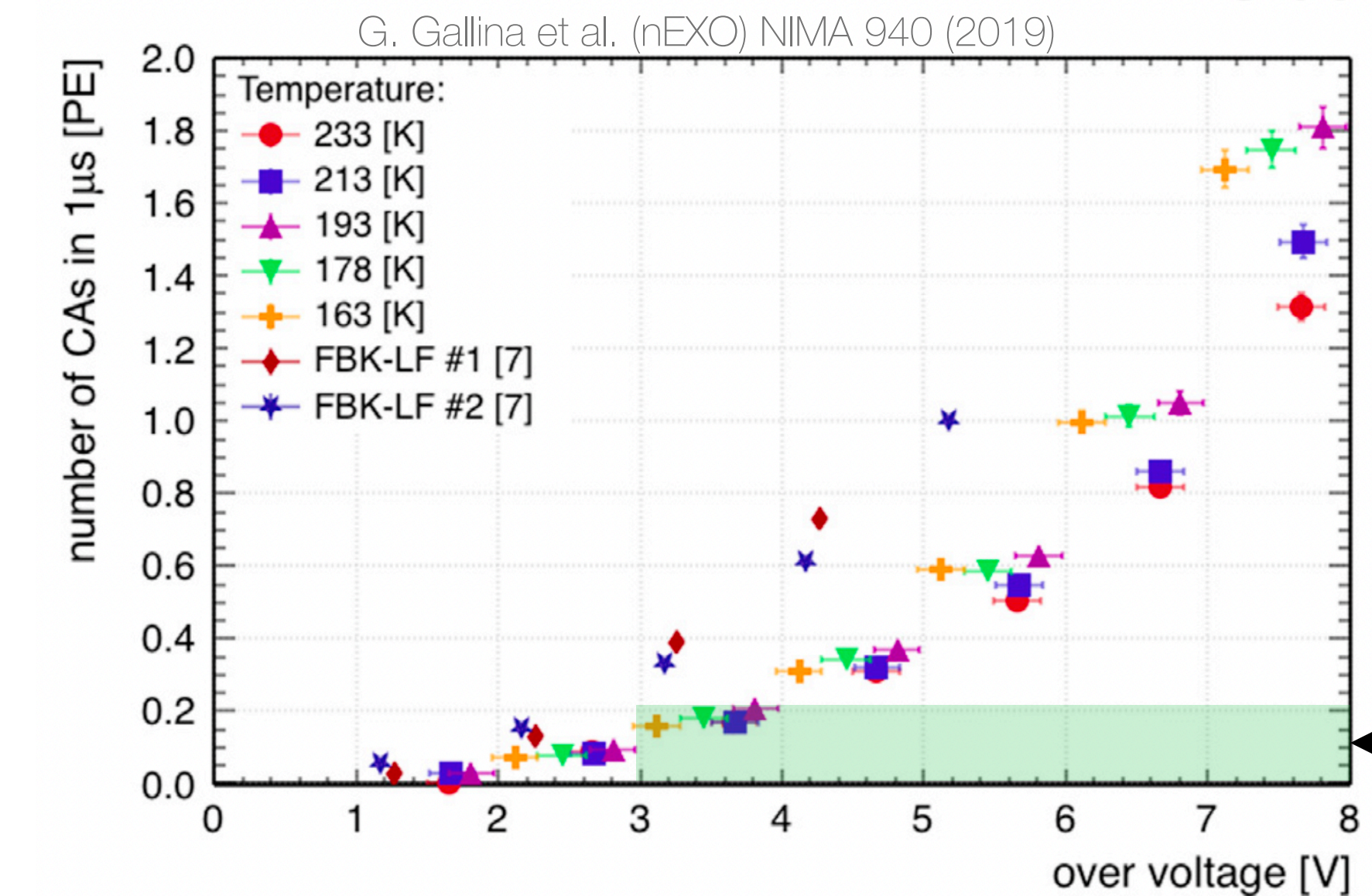
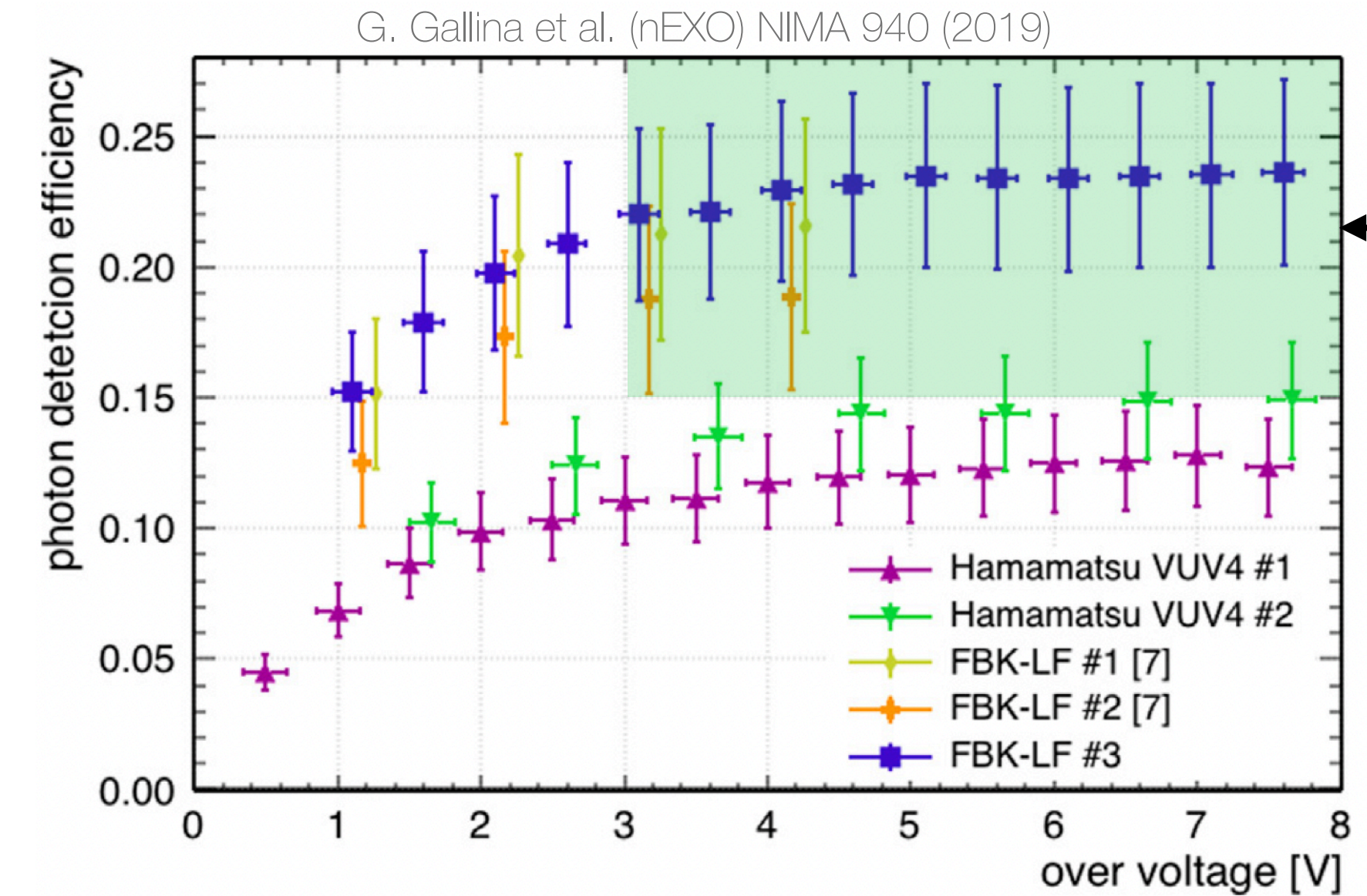
# DNN-based Signal and Background Discrimination

- Motivated by success in EXO-200
  - Allows equivalent discrimination with improved signal efficiency due to recovering of signal events accompanied by Bremsstrahlung
- Trained on waveform-level simulations, as we would with real data
- Training dataset:  $0\nu\beta\beta$  and  $\gamma$ 's with
  - Uniform energy between **900 keV – 3600 keV**
  - Uniform and random distribution in the detector
- Disentangled DNN variable from the other two fit dimensions
- Expect  $\sim 80\%$  signal efficiency at  $\sim 5\%$  background misidentification



# Characterization of SiPM Performance

- nEXO is running an extensive characterization campaign with several setups measuring
  - Absolute PDE in vacuum
    - Ostrovskiy et al. (nEXO) IEEE TNS 62 (2015)
    - A. Jamil et al. (nEXO) IEEE TNS 65 (2018)
    - G. Gallina et al. (nEXO) NIMA 940 (2019)
- Have identified devices that meet our requirement
- Working together with vendors to increase operational range



Within nEXO requirements

# Characterization of SiPM Performance

- nEXO is running an extensive characterization campaign with several setups measuring
- Reflectivity in vacuum and LXe
  - P. Nakarmi et al. (nEXO) JINST 15 (2020)
  - P. Lv et al. (nEXO) IEEE TNS 99 (2020)
  - M. Wagenpfeil et al. (nEXO) In prep. (2021)
- Photons reflected from SiPM surface can be detected by other SiPMs
- Reflectivity of passive TPC components crucial for good light collection efficiency

