

Recent Results from PandaX experiment

Qihong Wang

Fudan University

On behalf of the PandaX Collaboration

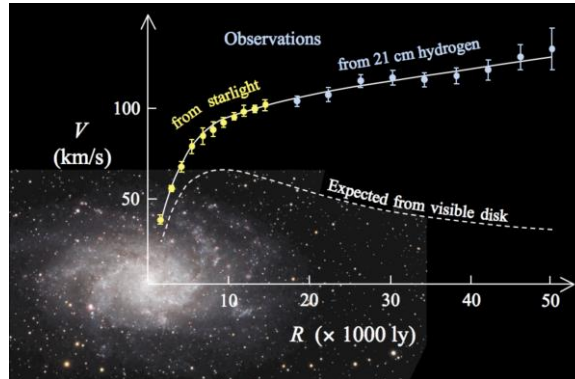


PANDA X
PARTICLE AND ASTROPHYSICAL XENON TPC





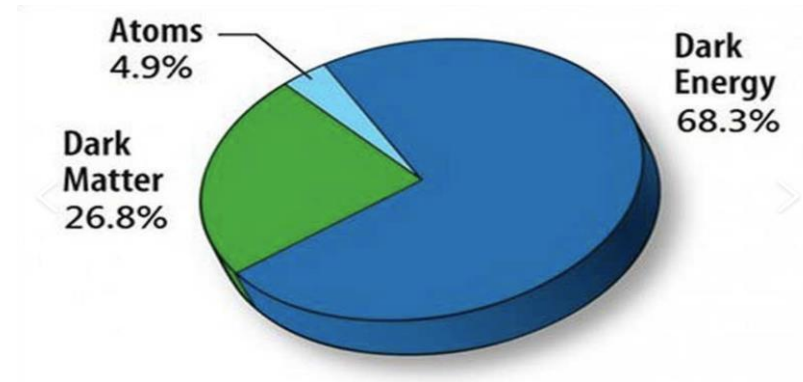
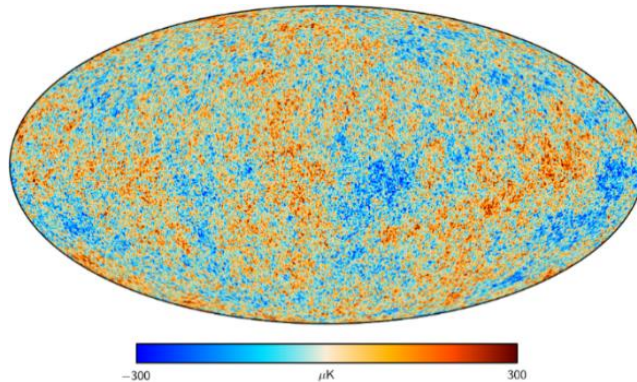
Galactic rotation curve



Bullet Cluster

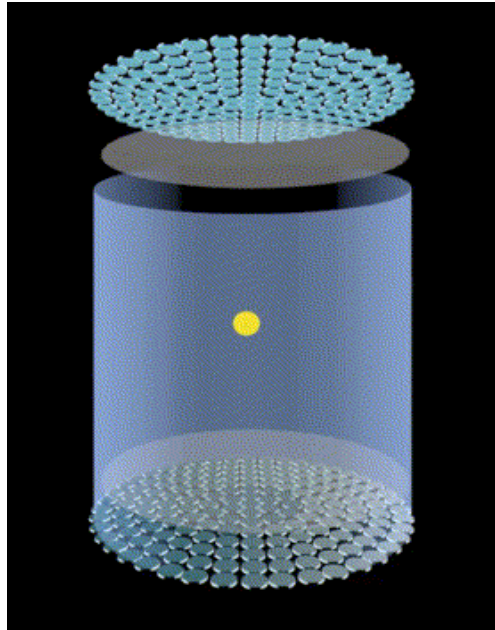
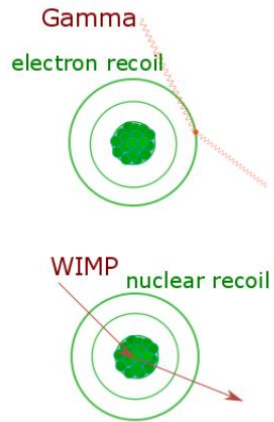


Cosmic Microwave Background

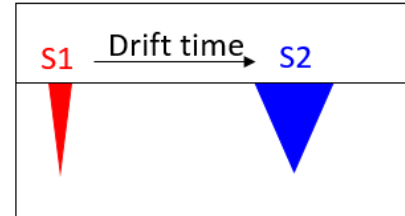


- Gravitational evidences suggest dark matter is the dominant form of matter in Universe!
- The nature of dark matter is still a mystery: WIMPs?

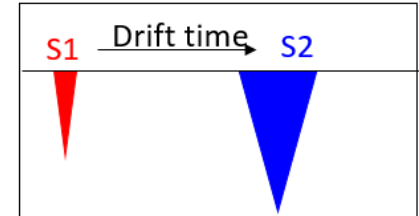
Dual phase xenon TPC



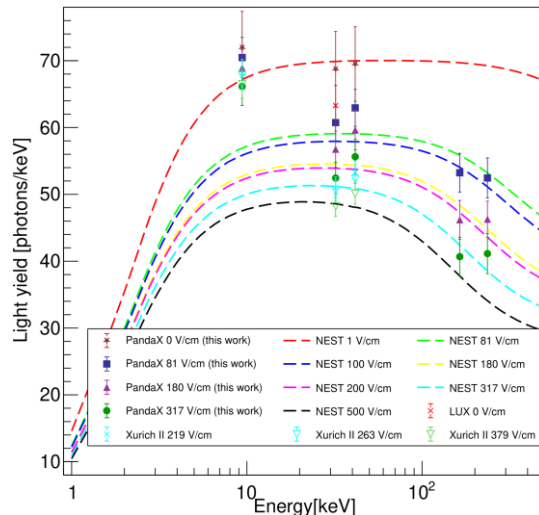
Dark matter: nuclear recoil (NR)



γ background: electron recoil (ER)



$$(S2/S1)_{NR} \ll (S2/S1)_{ER}$$



JINST 17 P01008 (2022)

Detector capability:

- Large monolithic target
- Good ER/NR rejection
- 3D reconstruction and fiducialization
- High light and charge yields

Particle and Astrophysical Xenon Experiments

Collaboration formed



2009.3

PandaX-I started

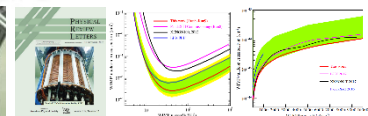


2014.3

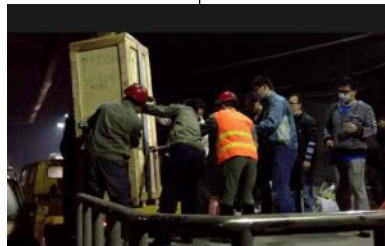
PandaX-II, 580 kg operation



2016.7
-2019.7



2019.8-



PandaX-I apparatus moved to Jinping

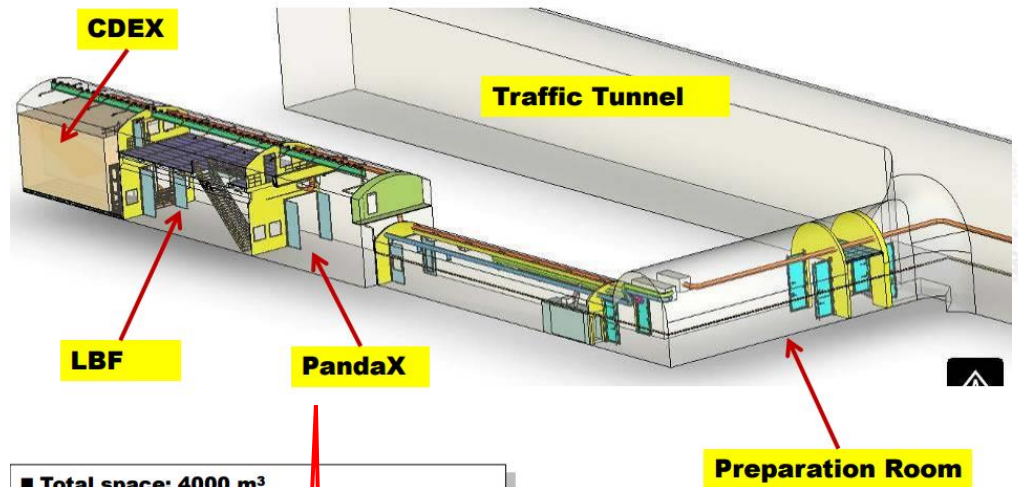
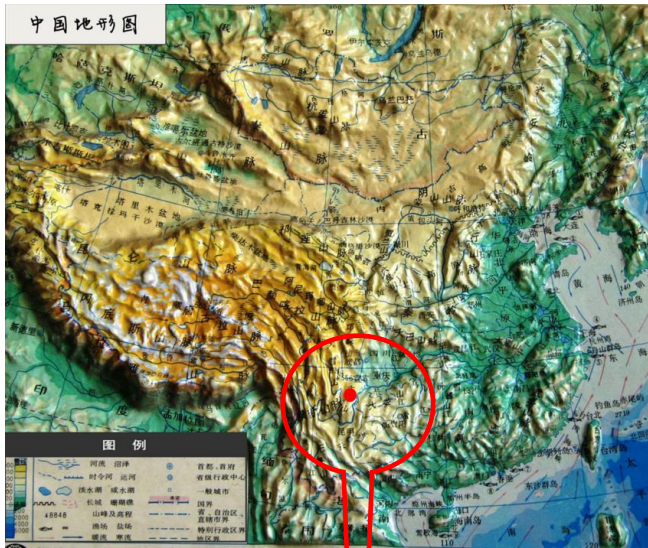


PandaX-I, 120 kg operation

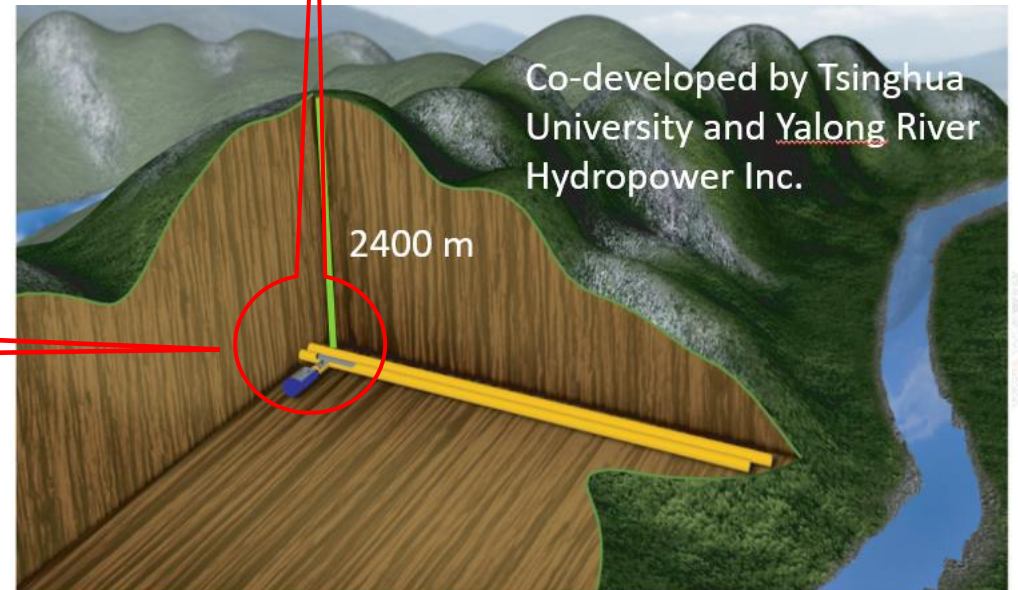


PandaX-4T, 3.7 t moved to CJPL-II

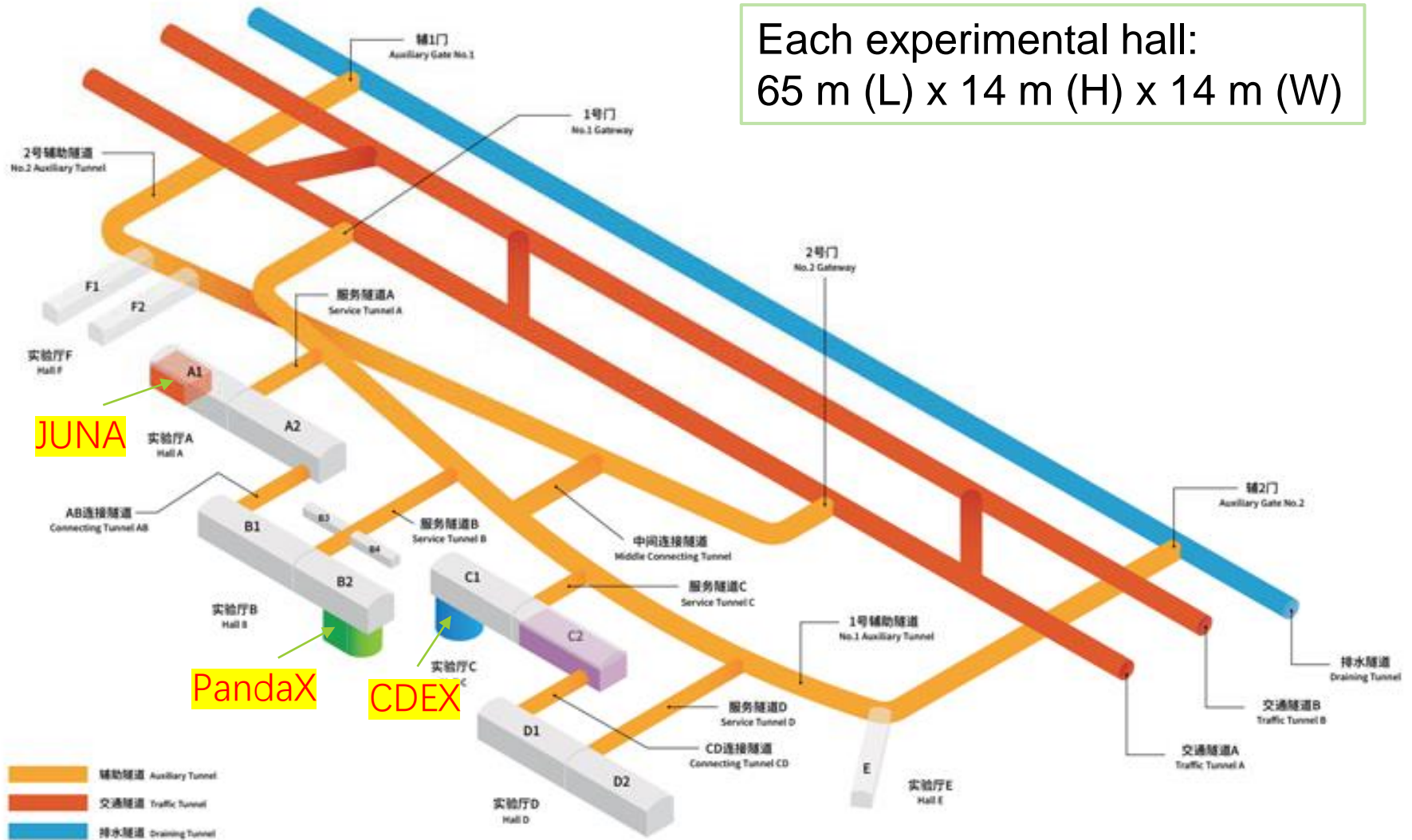
China Jinping underground Laboratory

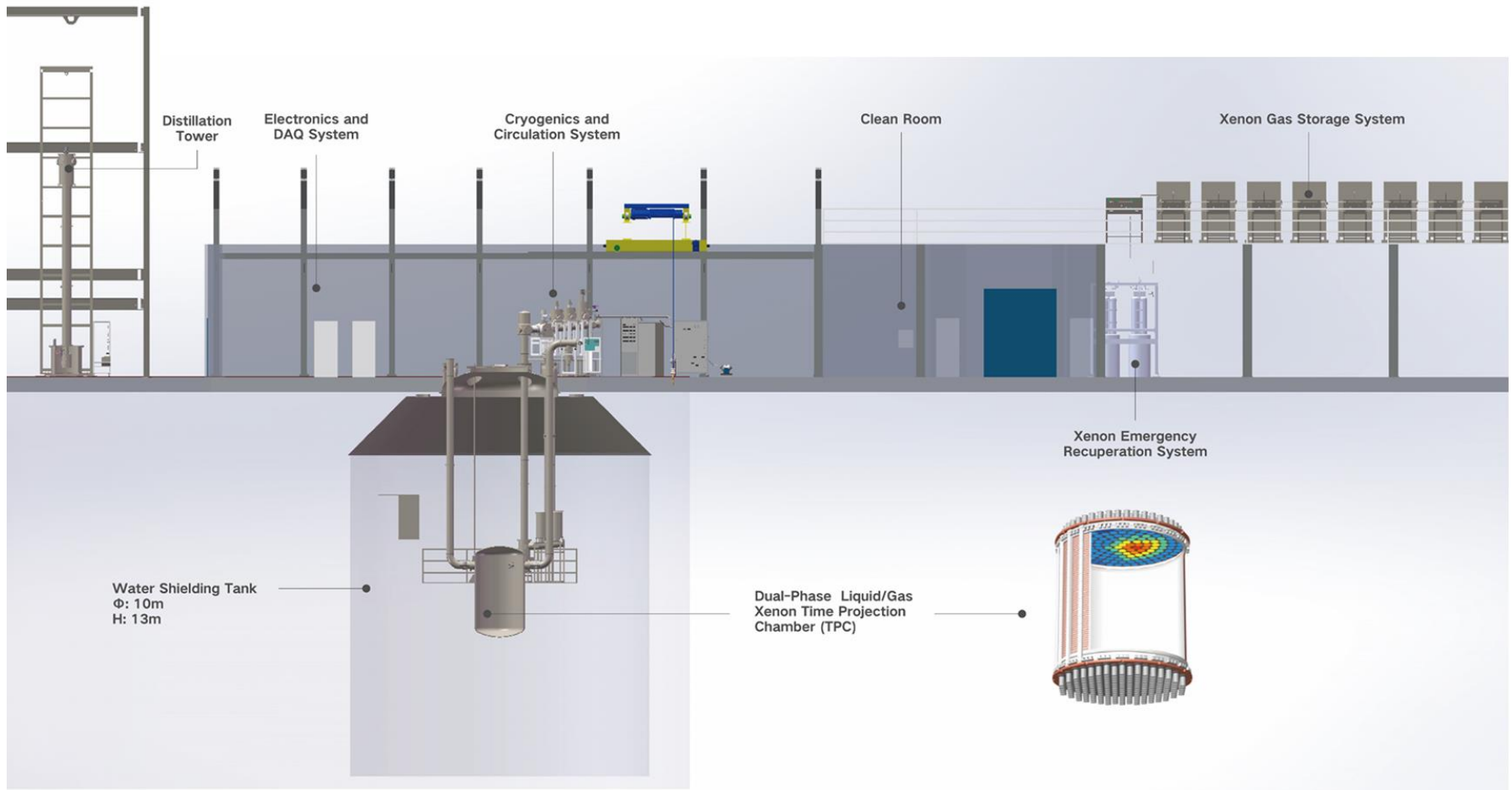


- Total space: 4000 m³
- Main Lab Space: 6.5(W) x 6.5(H) x 42(L)



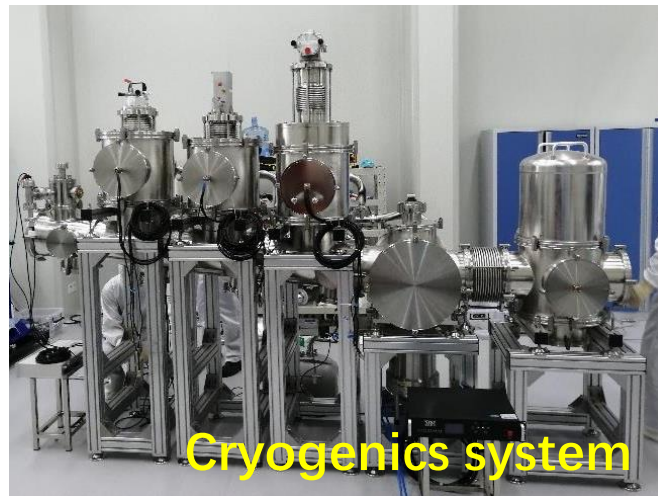
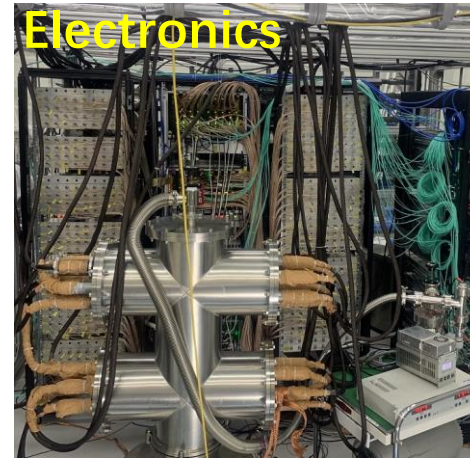
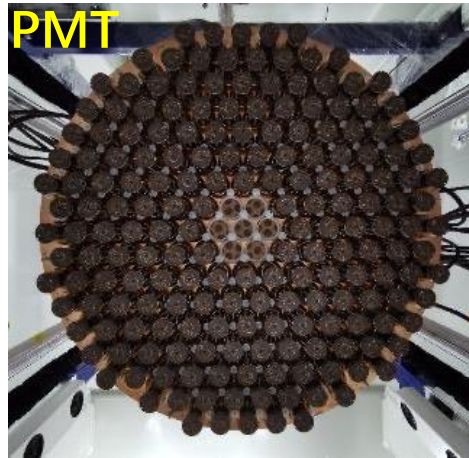
Each experimental hall:
65 m (L) x 14 m (H) x 14 m (W)





- ❑ Ultrapure water shield: 13 m (H) x 10 m (D) ~ 900 m³
- ❑ TPC: 1.2 m (H) x 1.2 m (D)
- ❑ 3-in PMTs: 169 top/199 bottom

PandaX-4T Subsystems

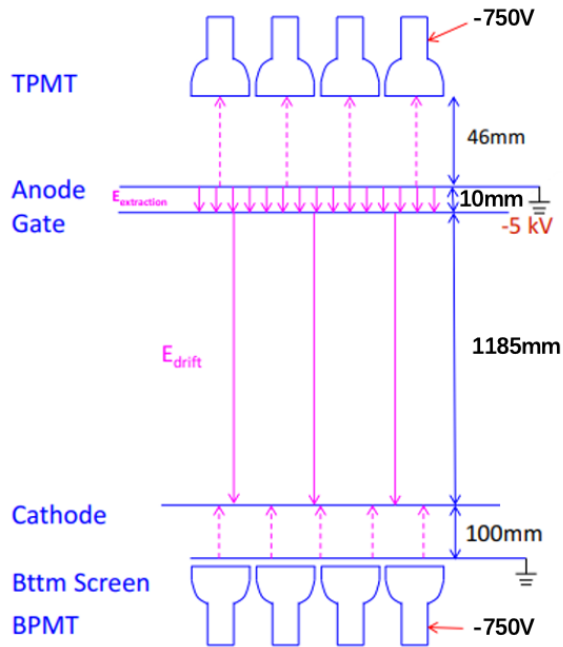


Ultrapure water filling

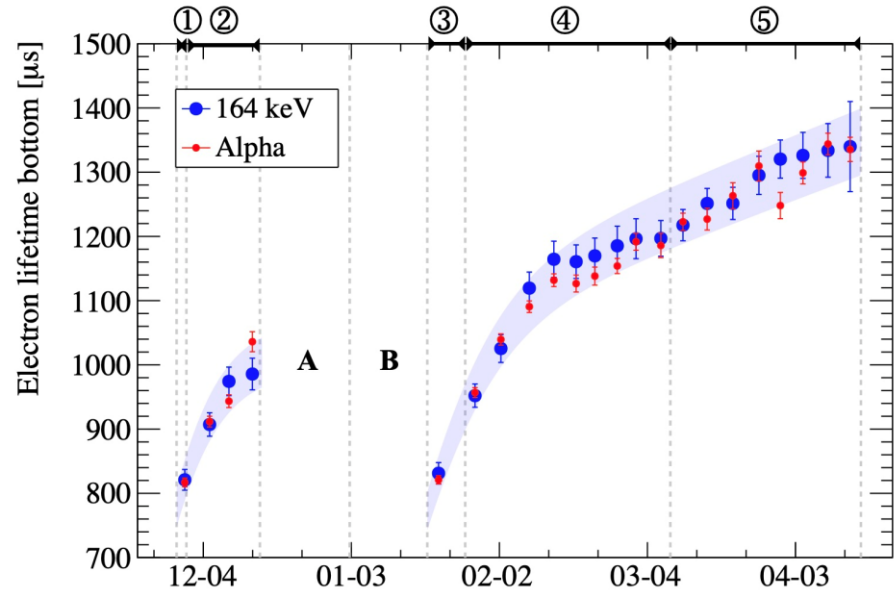
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TPC conditions & data taking history



①-⑤: Commissioning data taking subset

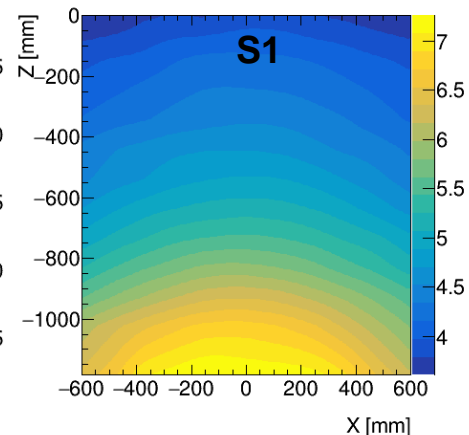
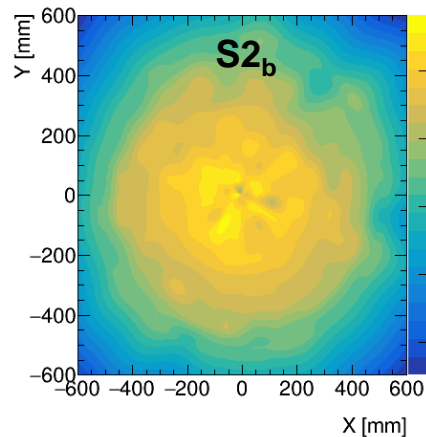
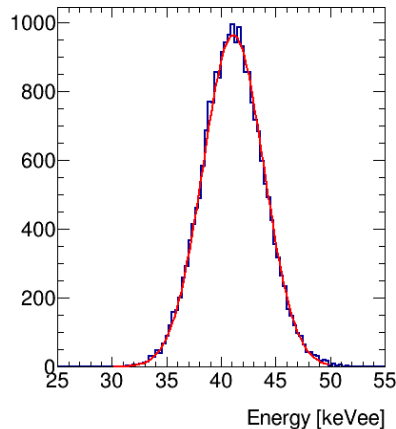


A: HV training
B: Circulation pump replacement

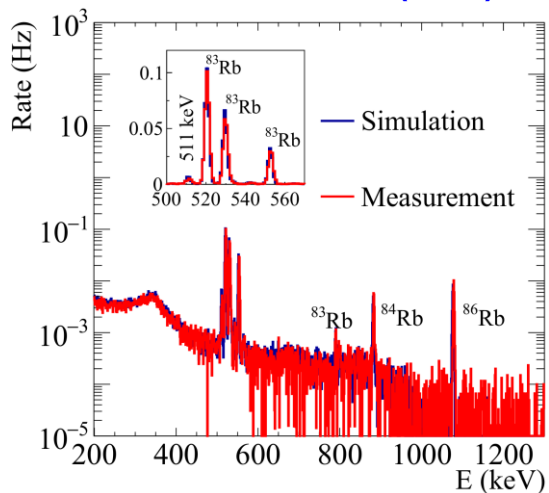
Set	1	2	3	4	5
Duration (day)	1.95	13.25	5.53	35.58	36.51
$\langle \tau_e \rangle$ (μs)	800.4	939.2	833.6	1121.5	1288.2
dt_{max} (μs)	800	810	817	841	841
V_{cathode} (-kV)	20	18.6	18	16	16
V_{gate} (-kV)	4.9	4.9	5	5	5
PDE (%)	9.0 \pm 0.2		9.0 \pm 0.2		
EEE (%)	90.2 \pm 5.4		92.6 \pm 5.4		
SEG _b (PE/e)	3.8 \pm 0.1		4.6 \pm 0.1		

- Different HV to avoid excessive discharges
- Electron lifetime: *in situ* S2 vertical uniformity calibration
- Ref: the maximum drift time \sim 840 μs
- Stable data running period: 95.0 calendar days (86 days after selection)

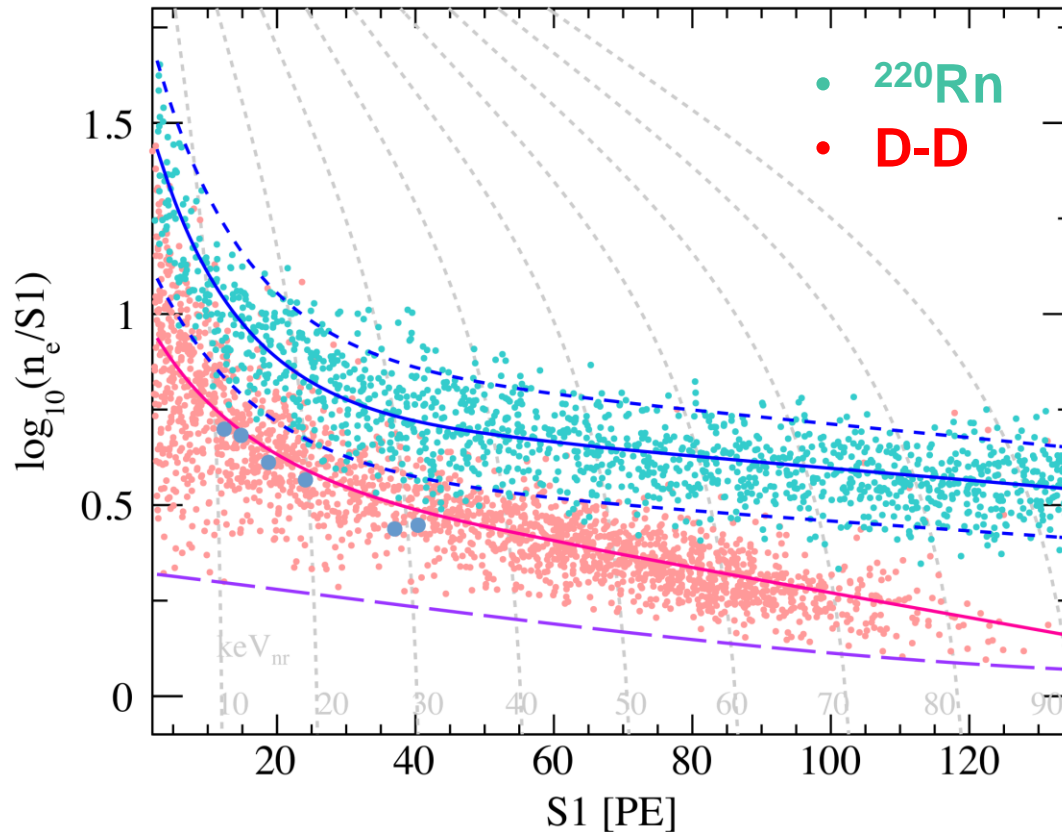
Gaseous source injection panel



PRC 105, 014604 (2022)



- ❑ Internal calibration sources: ^{83m}Kr
- ❑ ^{83m}Kr production with 3.4/20 MeV proton beams
($p + \text{natKr} \rightarrow ^{83}\text{Rb} \rightarrow ^{83m}\text{Kr}$)
- ❑ Energy resolution @ 41.5 keV: 6.8%
- ❑ S2_b RMS in FV: 15%
- ❑ S1 RMS in FV: 19%



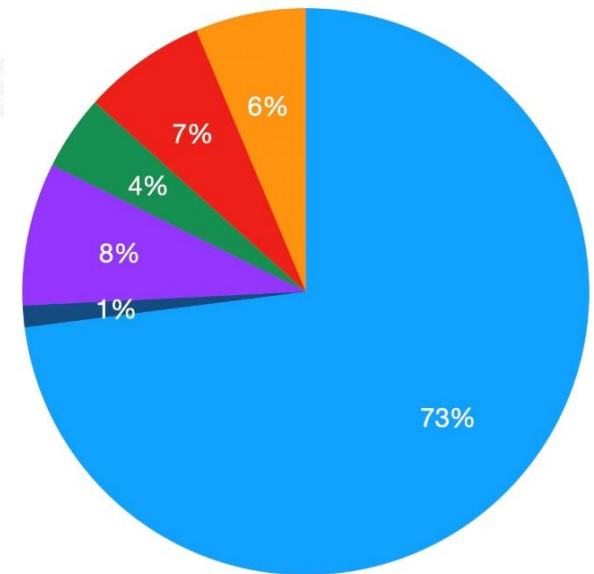
- ❑ Nuclear Recoil (NR) calibration with D-D & Am-Be neutrons
- ❑ Electron recoil (ER) calibration with ²²⁰Rn
- ❑ ER leak ratio (below NR median) = $6/1393 = 0.43\% \pm 0.18\%$

Background composition

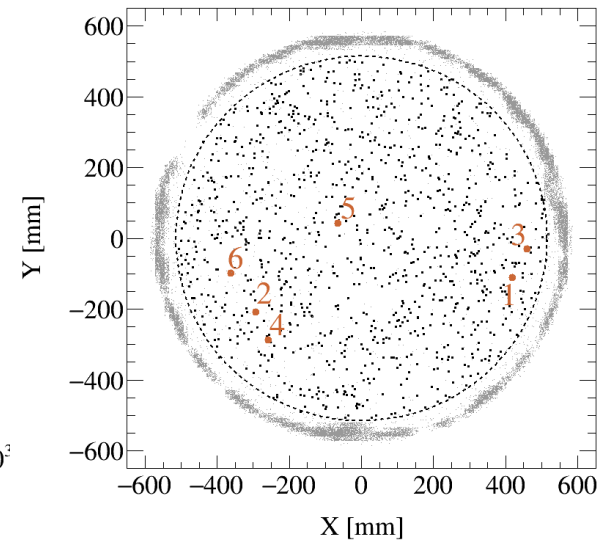
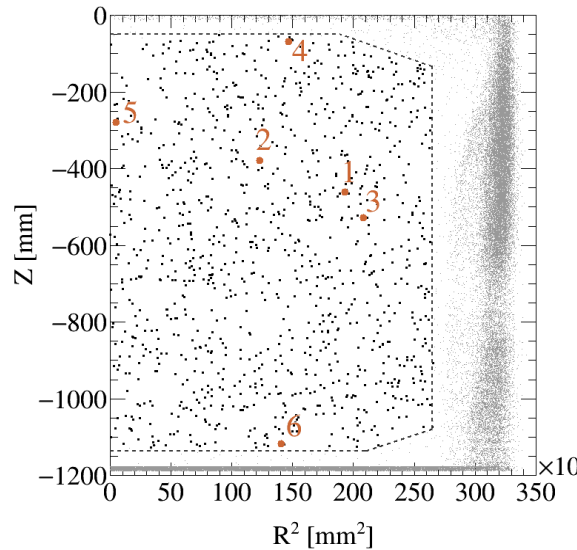
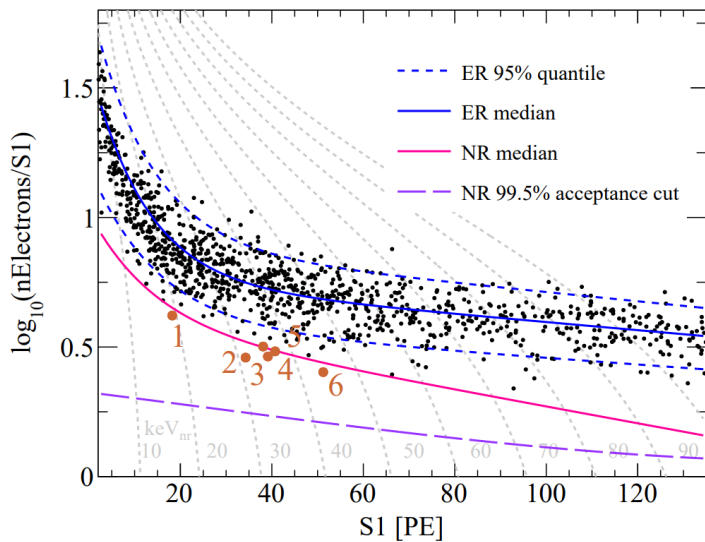
Component	Nominal (evts)
^3T (from fit to data)	532 (32)
Flat ER* (18-30keV side band)	492 (31)
Rn	347 (190)
Kr	53 (34)
Material	40 (5)
Xe127	8 (1)
Neutron	0.9 (0.5)
Neutron-X	0.2 (0.1)
Surface	0.5 (0.1)
Accidental	2.4 (0.5)
B8	0.6 (0.3)
Sum	1037 (45)

- Flat ER (Rn+Kr+Material) is determined from side band in DM data
- Background per unit target is improved from PandaX-II by 4 times (<10 keV)

Expected below-NR-median events: 9.8 (0.6) evts

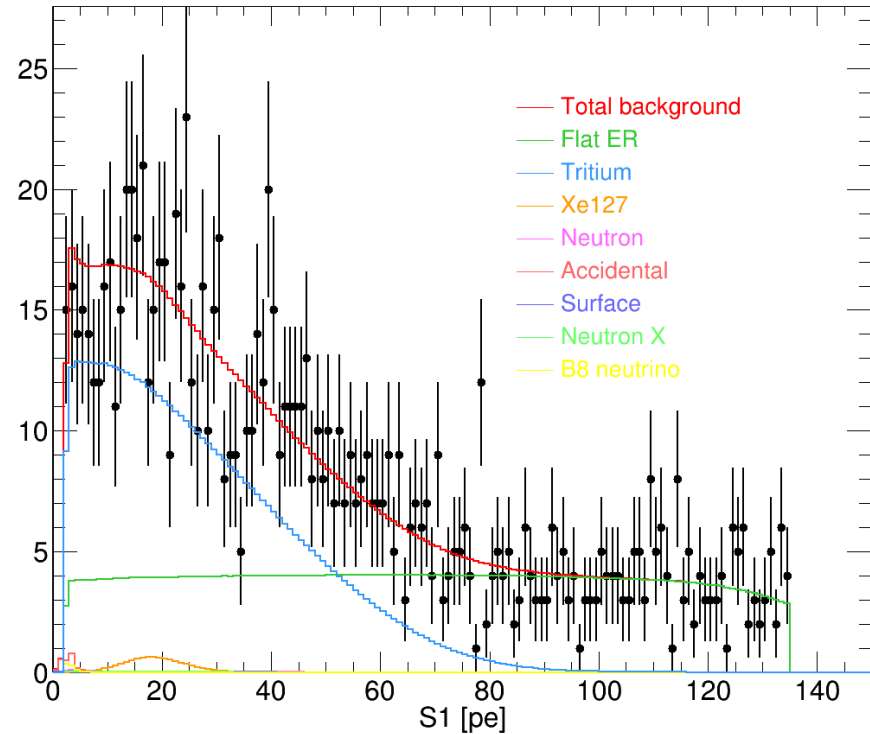


- ER (flat ER + tritium)
- Xe127
- Accidental
- Surface
- Neutron
- B8

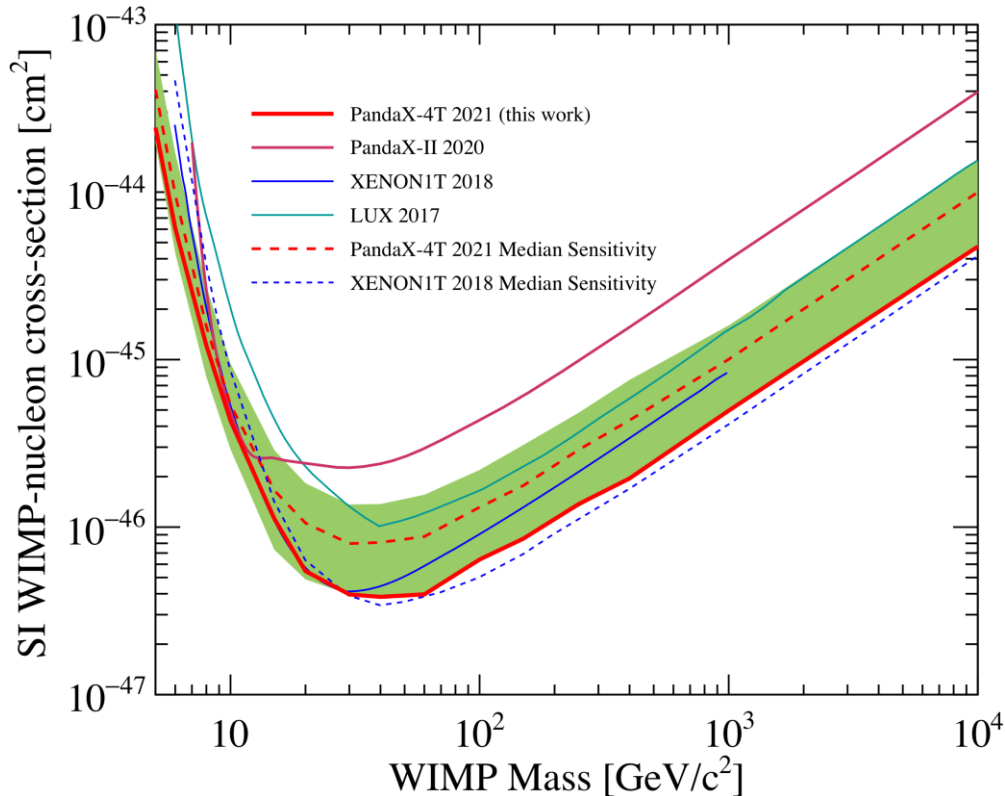


- ❑ $S1 = (2, 135)$ PE, $S2_{\text{raw}} > 80$ PE, $S2 < 20000$
- ❑ FV = 2.67 tonne
- ❑ 1058 candidates (expected 1054 ± 39), 6 below NR median curve (expected 9.8 ± 0.6)
- ❑ Events uniformly distributed in the FV, as expected if dominated by tritium and radon

$$\mathcal{L}_{\text{pandax}} = \left\{ \prod_{n=1}^{\text{nset}} \left[\text{Pois}(\mathcal{N}_{\text{obs}}^n | \mathcal{N}_{\text{fit}}^n) \times \prod_{i=1}^{\mathcal{N}_{\text{obs}}^n} (l_s^{n,i} + \sum_b l_b^{n,i}) \right] \right\} \times \left[G(\delta_s, \sigma_s) \prod_b G(\delta_b, \sigma_b) \right],$$



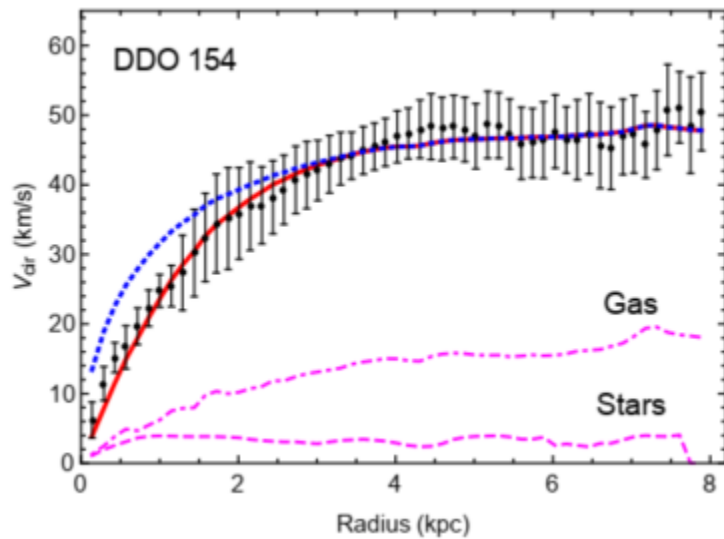
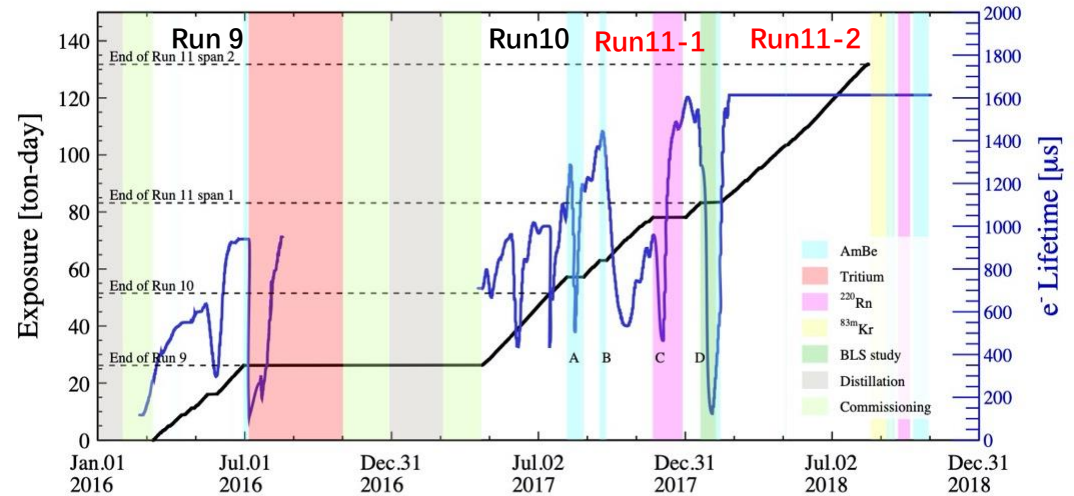
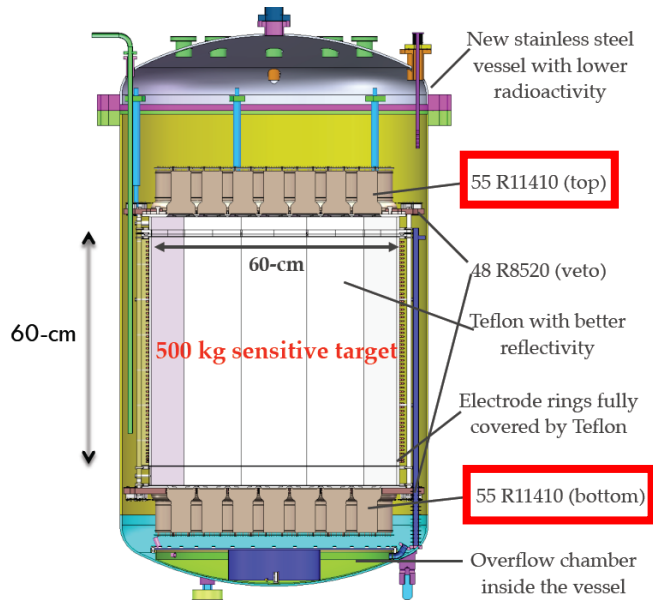
- ❑ Fit data with unbinned likelihood with all signal/background PDFs in $(S1, S2_b)$
- ❑ No excess found, background-only p-value **0.58**
- ❑ Projected S1 spectrum agrees with expected background



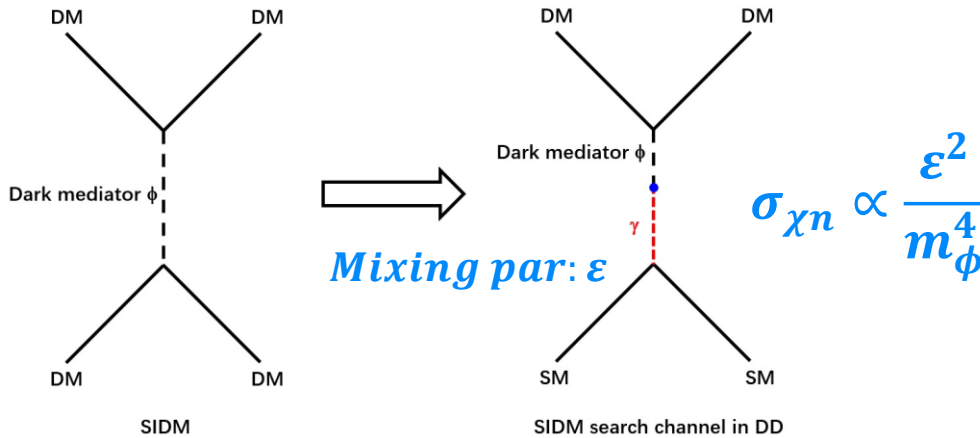
Phys. Rev. Lett. 127, 261802 (2021)

- ❑ Exposure: 0.63 tonne·year
- ❑ Sensitivity improved from PandaX-II final analysis by 2.9 times ($30 \text{ GeV}/c^2$)
- ❑ Our limit is ~ 1.24 times stronger than XENON1T around $30 \text{ GeV}/c^2$
- ❑ Dived into previously unexplored territory!
- ❑ Approaching the “low E” neutrino floor

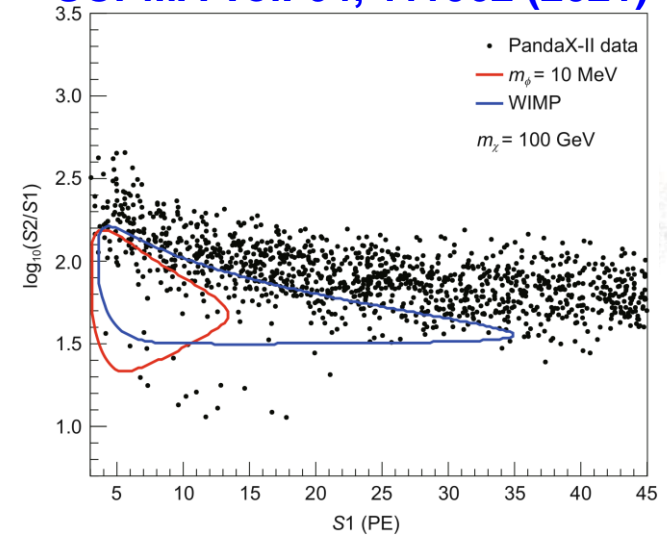
PandaX-II & Self-interacting DM



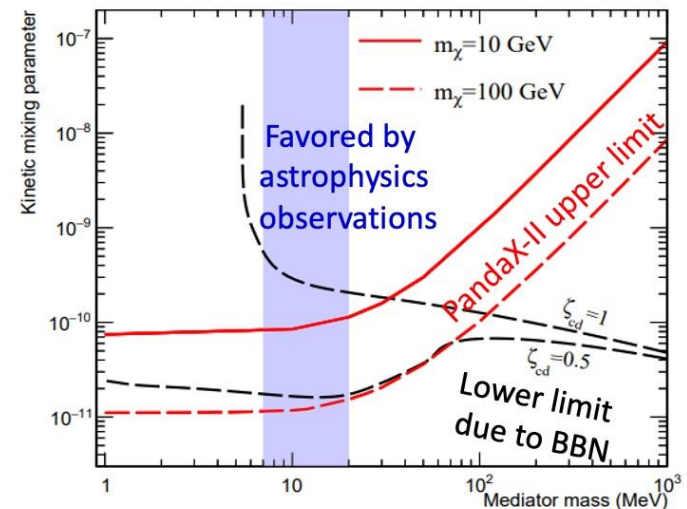
- PandaX-II: 132 tonne-day exposure
- Re-analysis on other DM models
- To solve discrepancies between CDM predictions and galactic observations
 - Self-interacting dark matter (SIDM)
 - Light mediator: $m_\phi \sim 10$ MeV



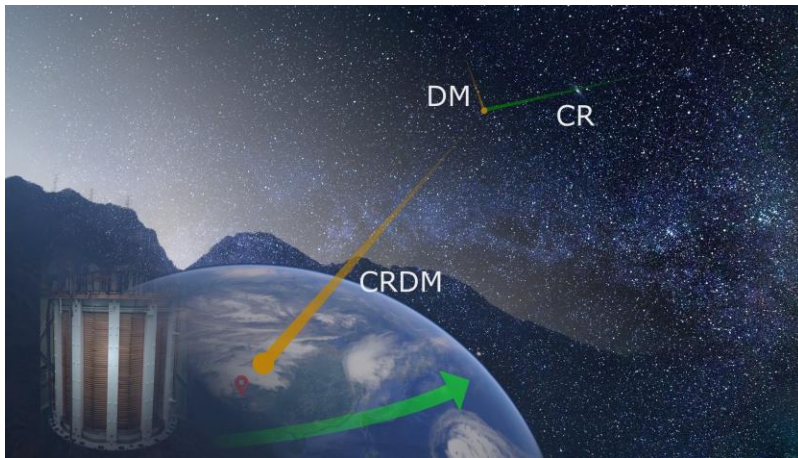
SCPMA Vol. 64, 111062 (2021)



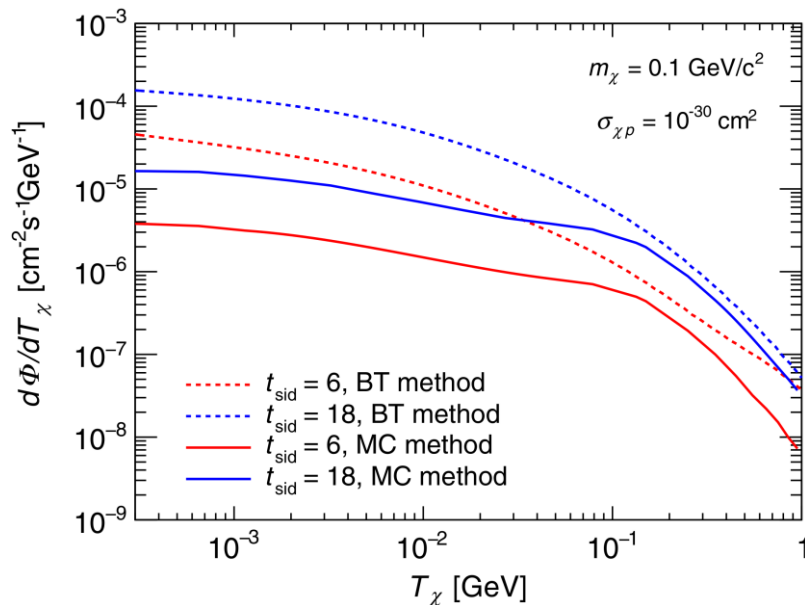
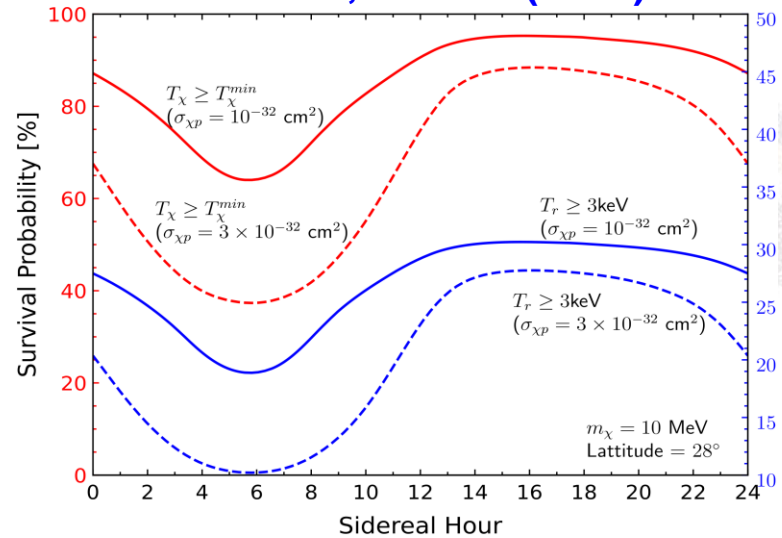
- Light mediator: $m_\phi \lesssim q$
 - Signal spectrum peaked more towards low energy
- PandaX-II results
 - No significant excess is found above background
 - Upper limits on $\sigma_{\chi n} \propto \epsilon^2 / m_\phi^4$
 - Combination with astrophysics and cosmology
 - Under SIDM, for 10-200 GeV DM mass, dark sector is colder than visible sector in early universe



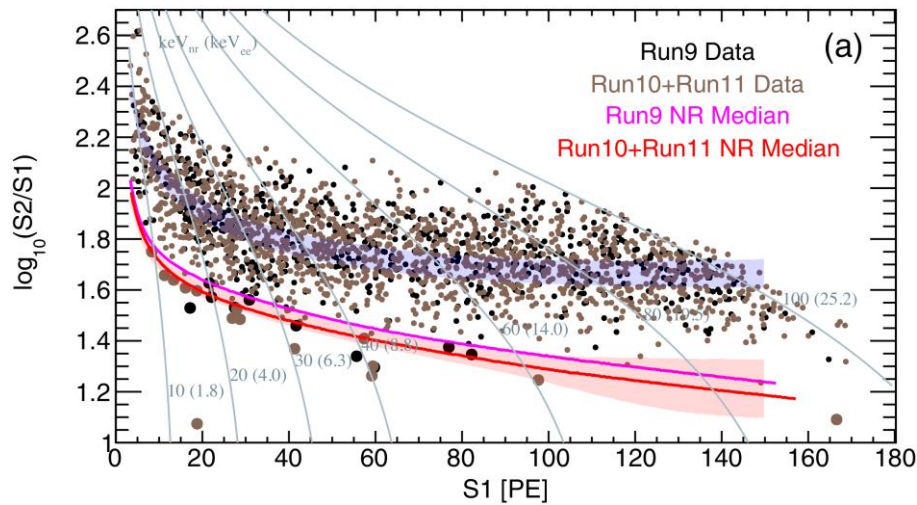
Cosmic-ray boosted dark matter (CRDM)



PRL 126, 091804 (2021)

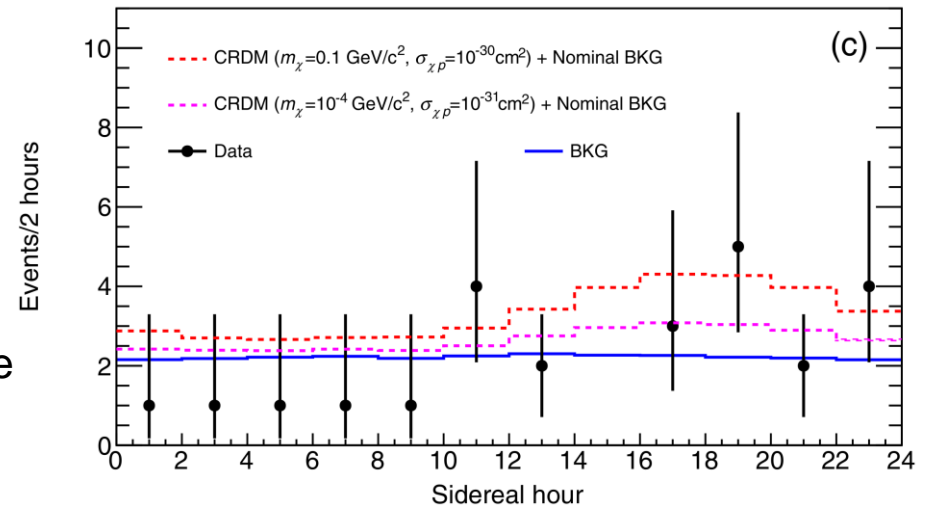
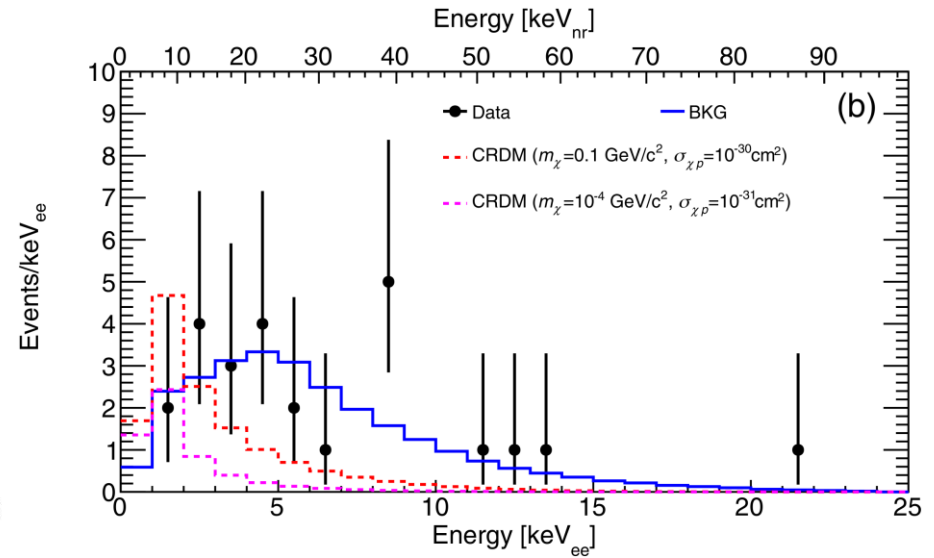


- ❑ Cosmic-ray boosts the kinetic energy of Galactic DM with sub-GeV mass
- ❑ Sidereal diurnal modulation in CRDM rate and recoil energy spectra
- ❑ Calculate attenuated CRDM flux at CJPL using full MC simulation

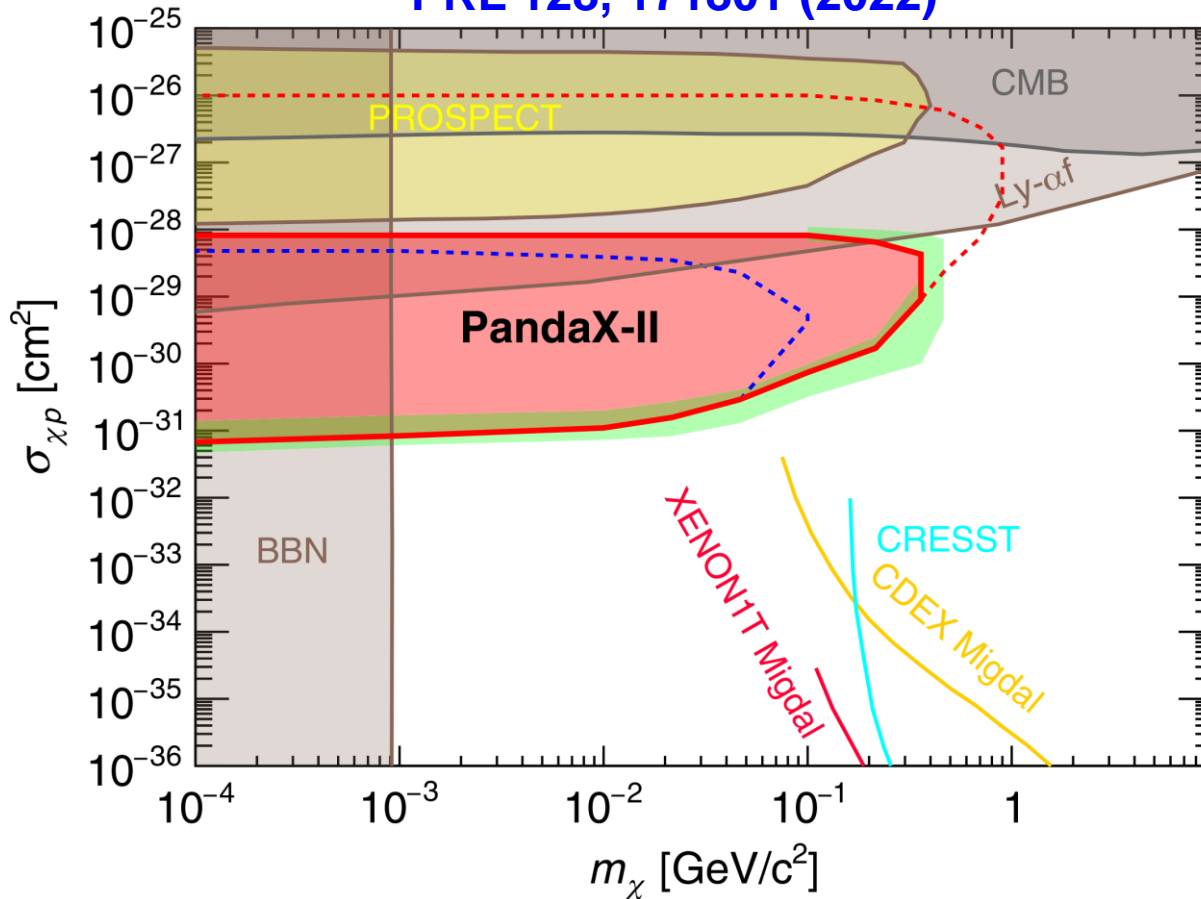


PRL 128, 171801 (2022)

- ❑ 100 tonne·day dataset of PandaX-II
- ❑ Using events below NR median: 25 events (expected 26.6 background)
- ❑ No significant CRDM signal is found above background.



PRL 128, 171801 (2022)



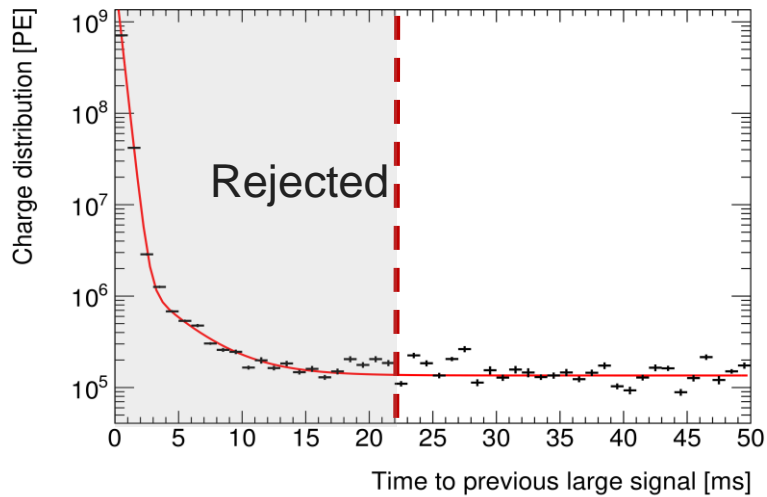
- ❑ Exclude $\sigma_{\chi p}$ between 10^{-31} and 10^{-28} cm² for DM mass [0.1 MeV, 0.1 GeV]
- ❑ Expand to the region beyond the astrophysical and cosmological probes

- ❑ PandaX-4T has completed its commissioning run
- ❑ With a 0.63 tonne·year exposure, PandaX-4T produced the strongest WIMP-nucleon interaction constraint
- ❑ An offline tritium removal campaign has been performed, new physics run is on going
- ❑ PandaX-II analysis on other physics topics: SIDM, CRDM
- ❑ In parallel, the collaboration is developing the plan for the next generation experiment at CJPL, we welcome collaborators!

Backups

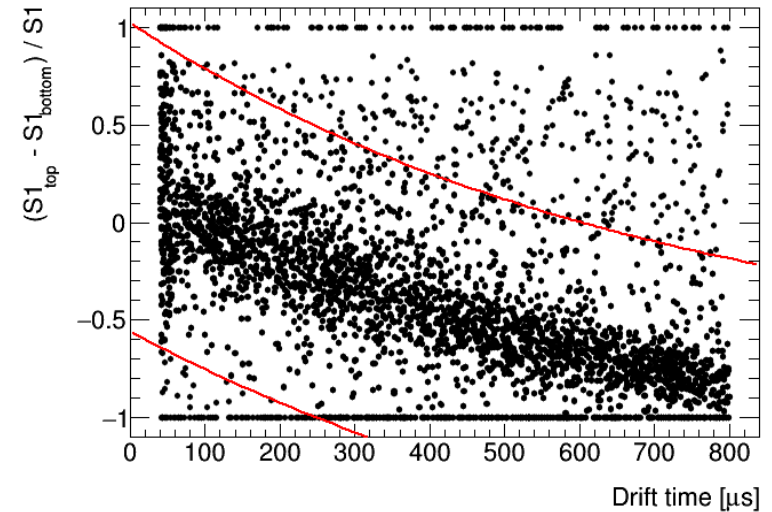


Time to previous large signals

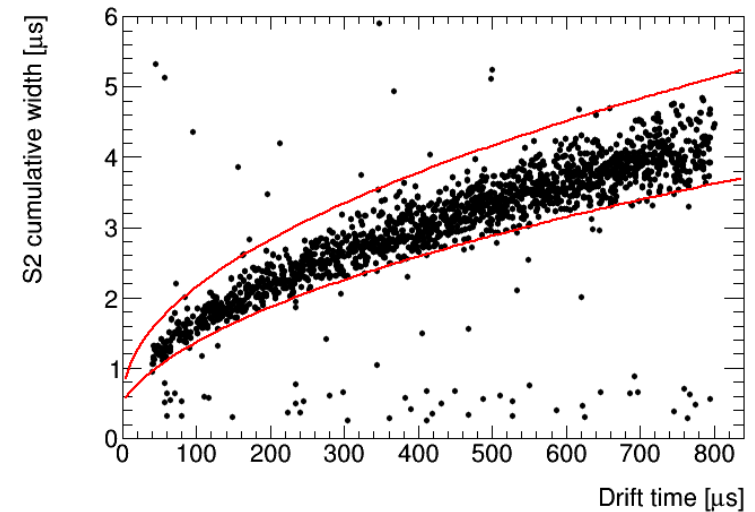


- ❑ Livetime fraction: 92.7%
- ❑ Rate excursion cuts: 97.7%

T-B asymmetry cut for S1

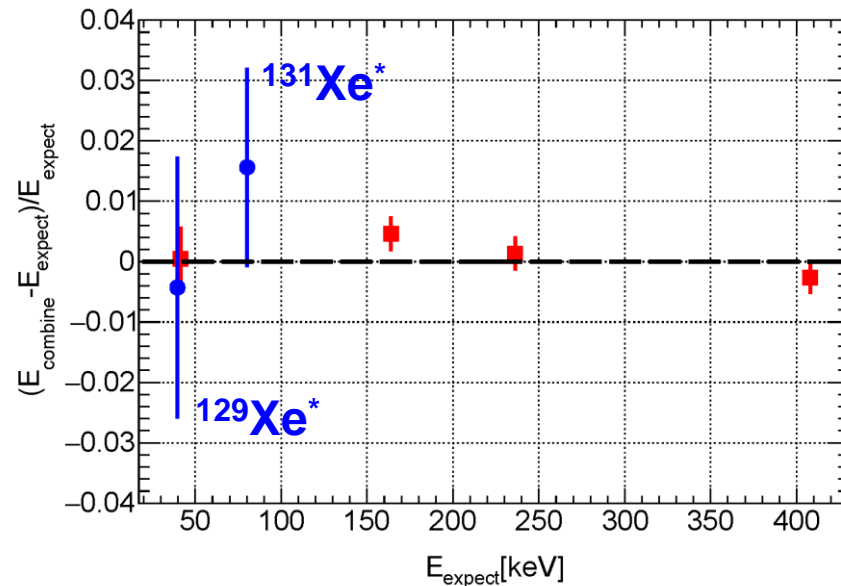
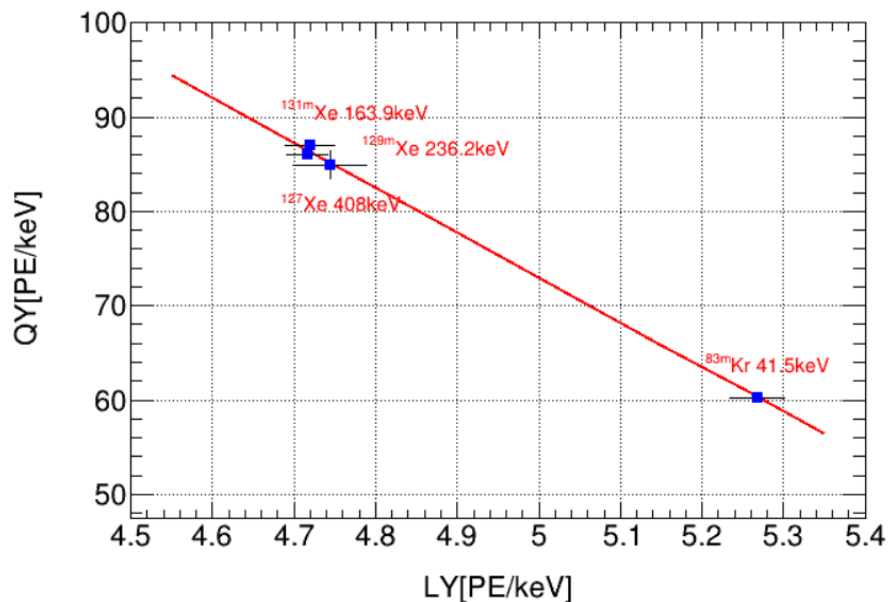


Diffusion cuts



Energy reconstruction

$$E = 13.7\text{eV} \times \left(\frac{S1}{\text{PDE}} + \frac{S2_b}{\text{EEE} \times \text{SEG}} \right)$$

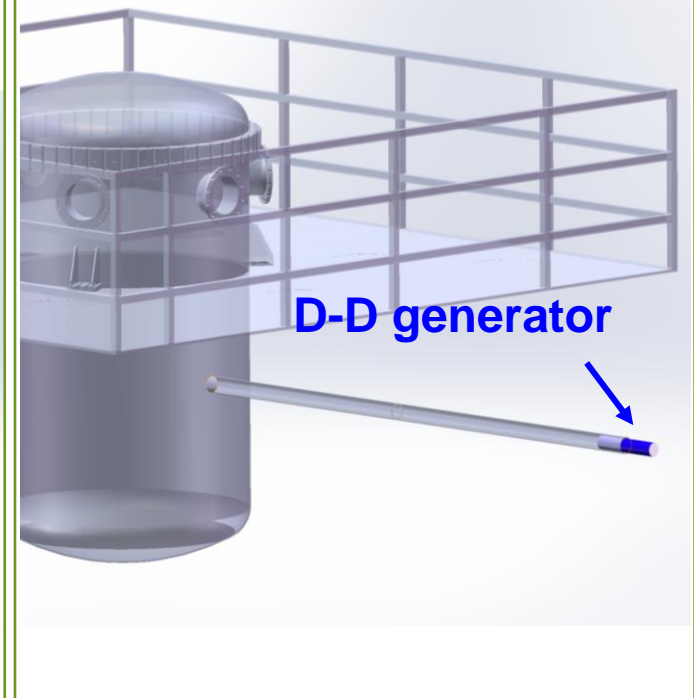
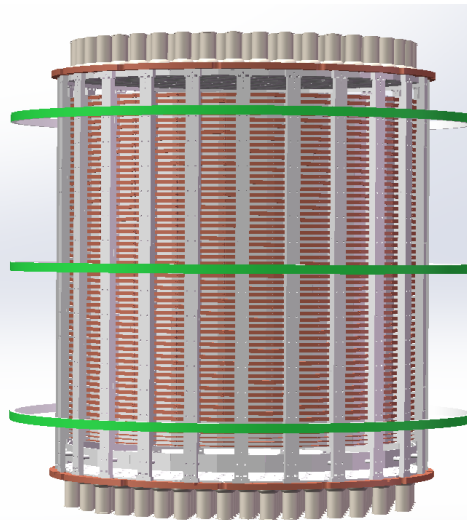


#Set	PDE [%]	EEE [%]	SEG [PE/e]
1-2	9.0 ± 0.2	90.2 ± 5.4	3.8 ± 0.1
3-5	9.0 ± 0.2	92.6 ± 5.4	4.6 ± 0.1

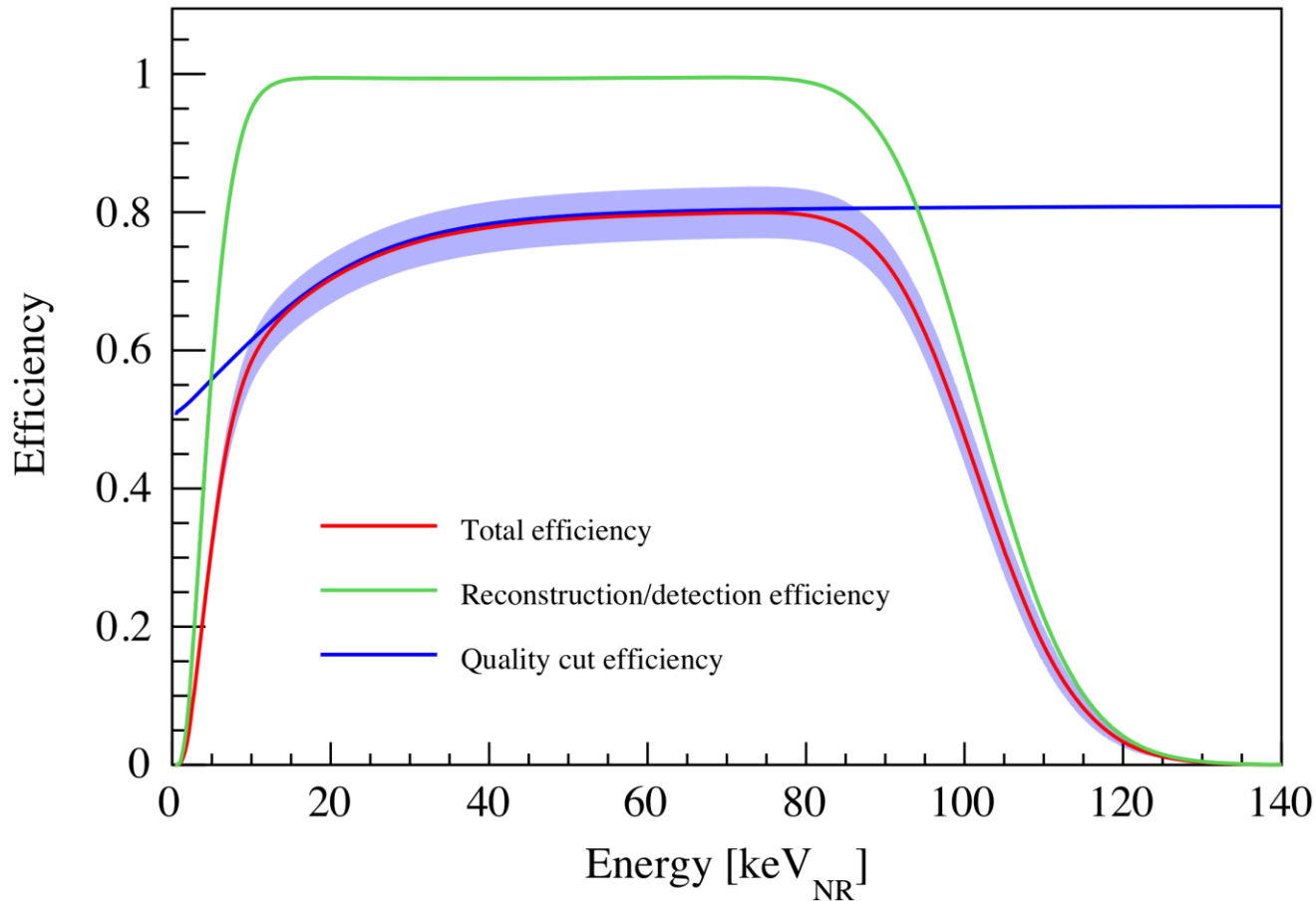
Gaseous source injection panel



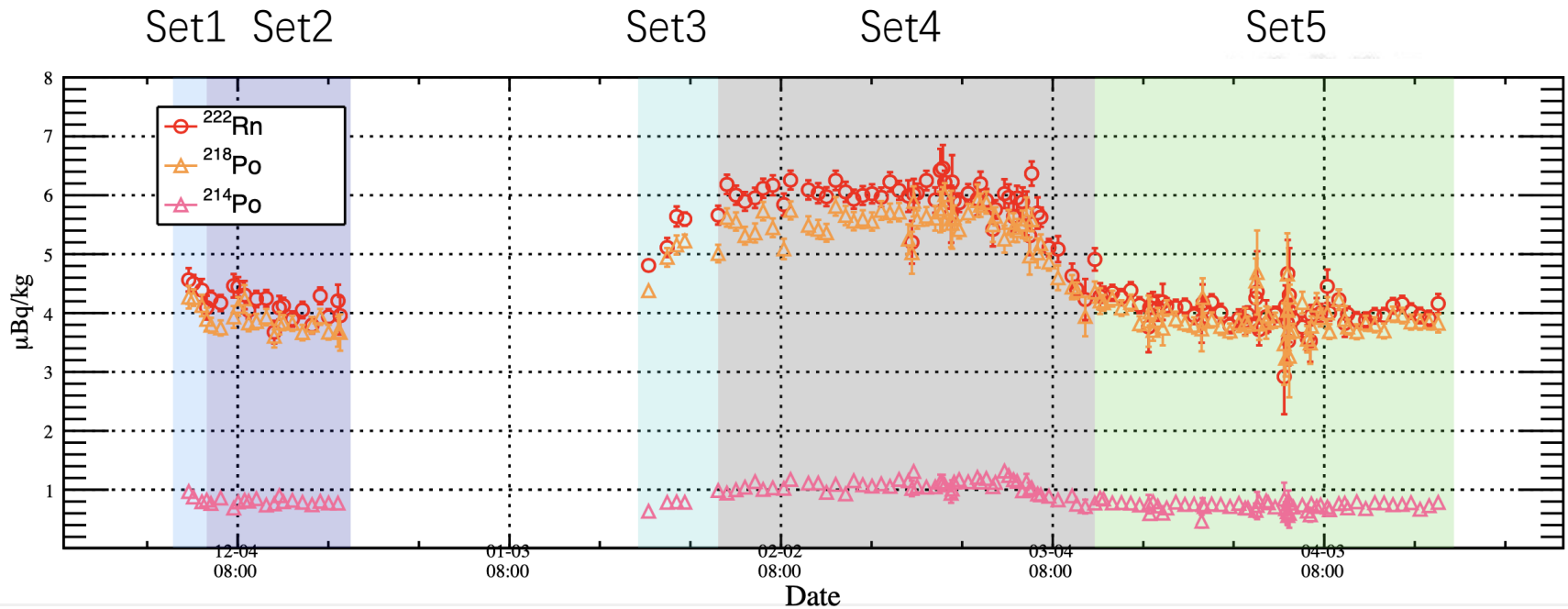
Calibration tubes



Calibration source	Position
$^{83m}\text{Kr}/^{220}\text{Rn}$	Injected from gas panel
$^{241}\text{Am-Be}$	Calibration tubes
D-D neutron	Beam pipe

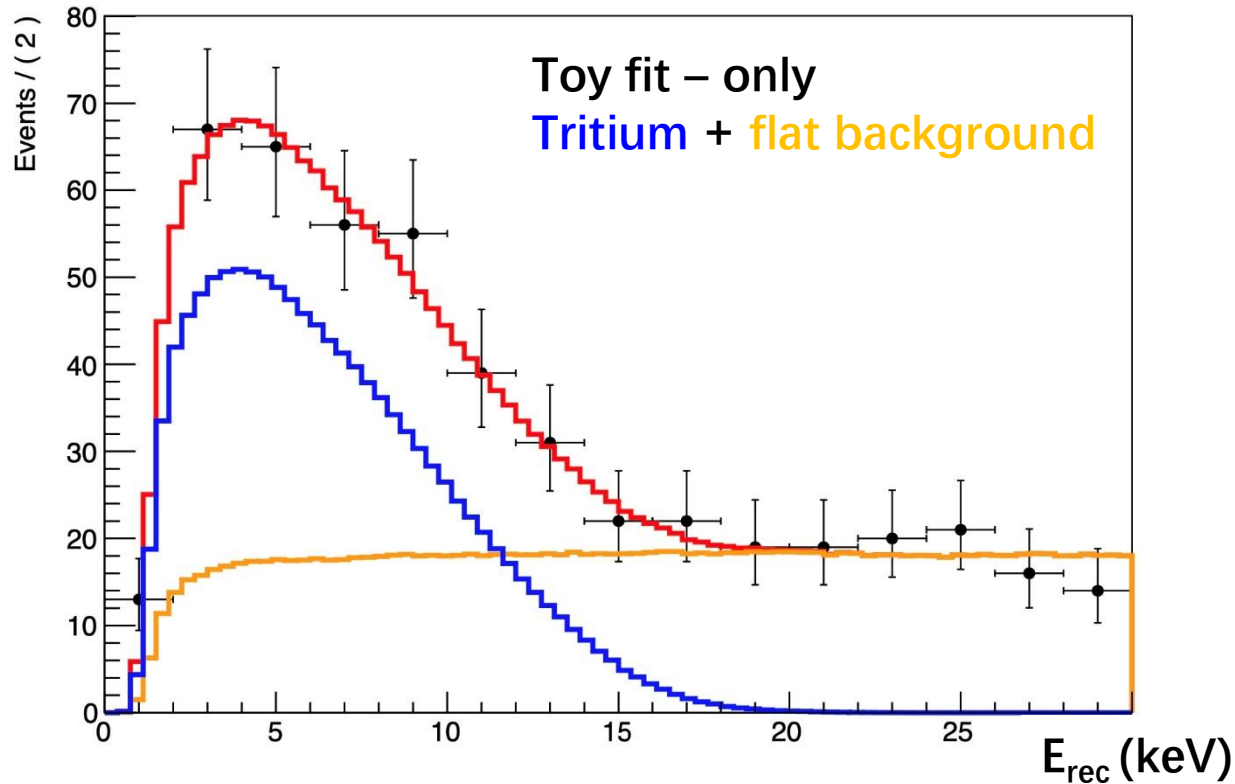


- ❑ Same S1 and S2 efficiency obtained from the ER and NR data
- ❑ Plateaued efficiency at 40 keV_{nr} ~78%.

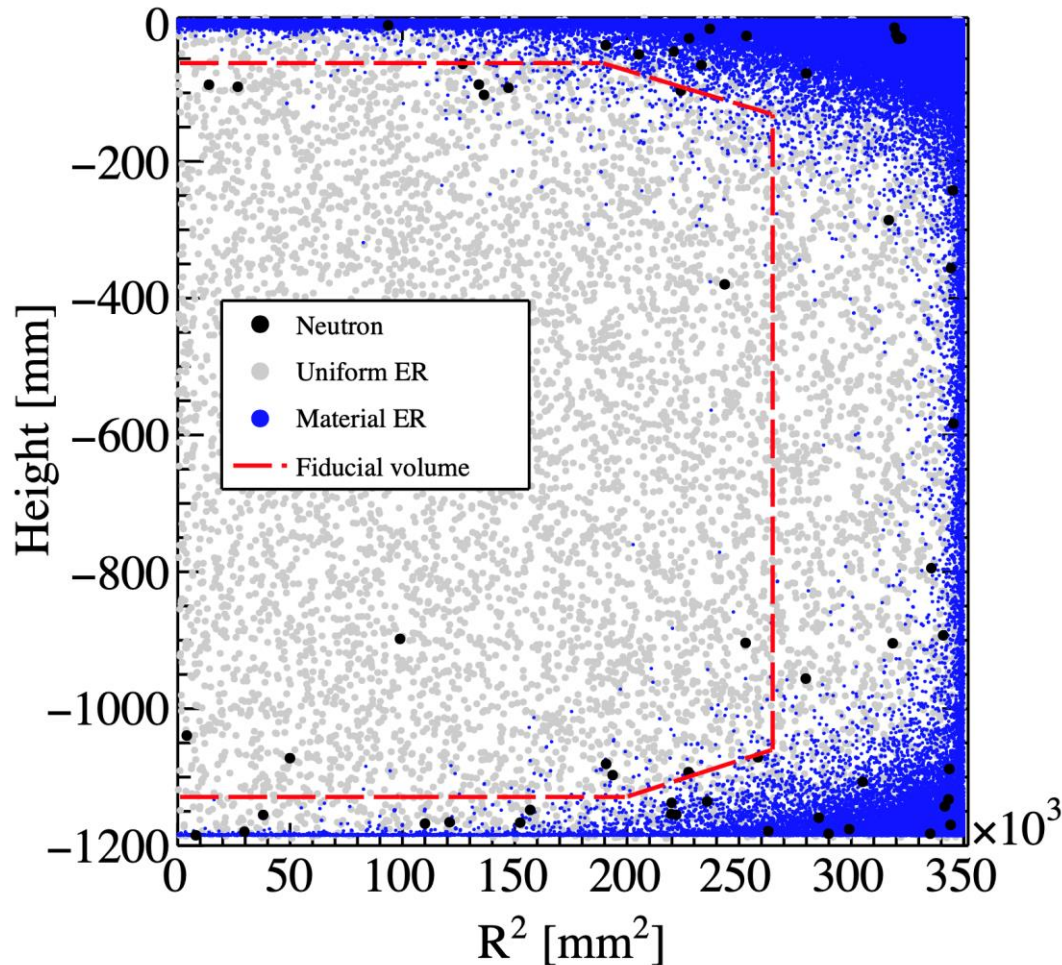


- ❑ Set 3→4: online Kr distillation (10 SLPM)
- ❑ Set 4→5: distillation off to reduce Rn emanation from the tower
- ❑ Low background directly extracted: 347 ± 190 events

Subset 4



- ❑ Tritium spectrum identified in the data
- ❑ Likely originated from a tritium calibration at the end of PandaX-II
- ❑ Level floating in the final dark matter fit: $\sim 5(0.3) \times 10^{-24}$ (mol/mol)



- ❑ Based on background simulation (10 t-year)
- ❑ Uniform ER (including tritium) normalization come from data
- ❑ Define FoM = \sqrt{B}/M
- ❑ **Optimized FV = 2.67 tonne**
- ❑ FV cuts in the data maintaining the same FV (correcting for reconstruction bias)

Likelihoods of the 6 below-NR events

