Kilotonne scale Xe detectors for Ovbb and new physics searches

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Background in underground physics:

- Dark matter (CDMS II, 2006-2012) [Soudan]
- Neutrinoless double beta decay:
 - EXO-200, 2012-present [WIPP]
 - nEXO, planned in coming years [Conceptual design @ SNOLAB]
 - Future ktonne Xe detector?



Science questions

- Big questions:
 - What is the origin of the small (but non-zero) neutrino masses?
 - Why is there matter in the Universe (rather than antimatter)?
 - Is lepton number conserved?
 - → Searches for neutrinoless double beta decay, possibly eventually reaching the minimal masses possible in normal hierarchy
 - Other motivations for very large Xe TPCs (with ~keV threshold):
 - Dark matter (mostly beyond WIMPs), CEvNS from solar, atmospheric, or supernovae neutrinos, tagging of CC interactions of solar v (e.g. CNO v or ⁷Be lineshape)

https://journals.aps.org/prd/abstract/10.1103/PhysRevD.102.072009



Existing/proposed ¹³⁶Xe 0vbb detectors



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Ovbb at 10³⁰ yr

 If ton-scale detectors do not discover 0vbb, eventual sensitivities of 10³⁰ yr may be required to explore majority of parameter space in normal hierarchy

100% $|\mathcal{M}^{0\nu}| = 2.7$ (Median published NME) 10³¹ 10 half-life [yr] 10₃₀ (*m_{ββ}*) [meV] 10% $m_{136} \sim 300 \text{ t}$ | *θ θ* 10²⁹ 1% $m_{136} \sim 5 t$ 10²⁸ 0.1% 0.1 0.1 0.1 0.01 10 0.01 10 100 100 *m*₁ [meV] *m*₁ [meV]

Allowed parameter space (normal hierarchy only, Nu-Fit 5.0):

Xe acquisition

- Primary challenge to extending Xe TPCs (gas or liquid) to ktonne scale is likely the acquisition of the Xe itself
 - E.g., LArTPCs at this scale exist, and substantially larger experiments are underway (see talks yesterday)
- For Xe, the enabling technology is isotope procurement
 - Current production (reliant on steel industry) is 50-100 tonnes per year
 - Alternative production path likely required to realize detectors beyond 10s of tons scale
 - Note that the amount of air required is << that needed to capture 1 Gton CO₂, although engineering challenges can be different due to lower concentration (89 ppb for Xe)
- Enrichment is not necessarily required, but if desirable enrichment by centrifuge with dedicated plants may be possible (but expensive)
- Backgrounds at this scale must also be demonstrated for any technology (see following slides)

Backgrounds

- If Xe could be procured in ktonne scale quantities, several advantages to incorporating it directly into a TPC:
 - Self-shielding: external U/Th backgrounds (as well as Rn induced backgrounds) substantially mitigated by scaling to large size
 - Energy resolution: sufficient to remove otherwise irreducible background from 2vbb decay
 - Solar v backgrounds: ^{nat}Xe (with ~10% isotope fraction) or ^{enr}Xe minimize inactive material in the detector volume, and single versus double beta discrimination possible in GXe
 - Unknown backgrounds: homogeneous detector gives multiple handles for tagging possible backgrounds, e.g. from possible rare cosmogenic activation products
- Sensitivity estimates for gas/liquid concepts appear to indicate sensitivities as long as 10³⁰ yr are achievable (300 t of ¹³⁶Xe)

Shielding of external backgrounds in LXe TPCs:



Detector concepts

- Ultimate 0vbb detector might be ktonne scale, although intermediate detectors also possible, e.g.:
 - Sensitivity $\sim 10^{29}$ yr with 50 t of ^{enr}Xe
 - Possible complementarity with WIMP searches at the neutrino floor
 - Other possibilities might be 300 t ^{nat}Xe detector (~10²⁹ yr), which can later be filled with ^{enr}Xe (~10³⁰ yr) if desired
 - Optimal choice between gas and liquid depends on backgrounds and engineering constraints
 - LXe detectors more compact for same mass
 - GXe detectors typically operate between 15-50 bar (6x to 30x lower density)



LXe detector size vs mass:

Summary

- Ktonne scale Xe detectors might be possible with new methods of Xe acquisition
 - R&D underway to explore various options (led by M. Heffner, LLNL)
 - Publication describing these ideas is in preparation
- If Xe could be procured in these quantities, extremely sensitive detectors for rare events may be
 possible, including 0vbb at ~10³⁰ yr half-lives
- Large underground spaces would be required by such detectors, but similar e.g. to scale of currently planned LAr TPCs