

Priscilla (Prisca) Cushman  
Professor of Physics, University of Minnesota

Graduate student with the hyperon group at FNAL

Postdoc with Rockefeller on CERN UA6

Assistant Prof. at Yale: BNL muon g-2 experiment

SSC Fellowship and R&D on hybrid APDs for SDC

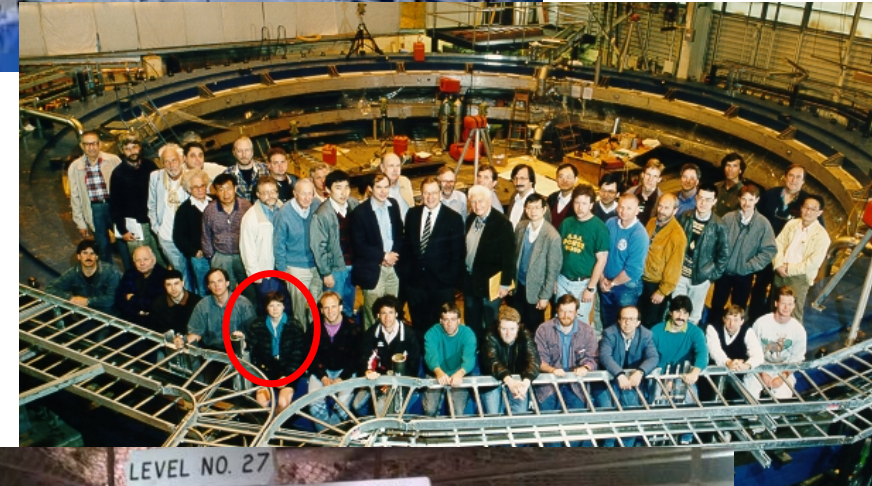
Recruited to University of Minnesota to start an SSC group

SSC cancelled 1 month after I moved

Began work on CMS (HCAL L3 manager)



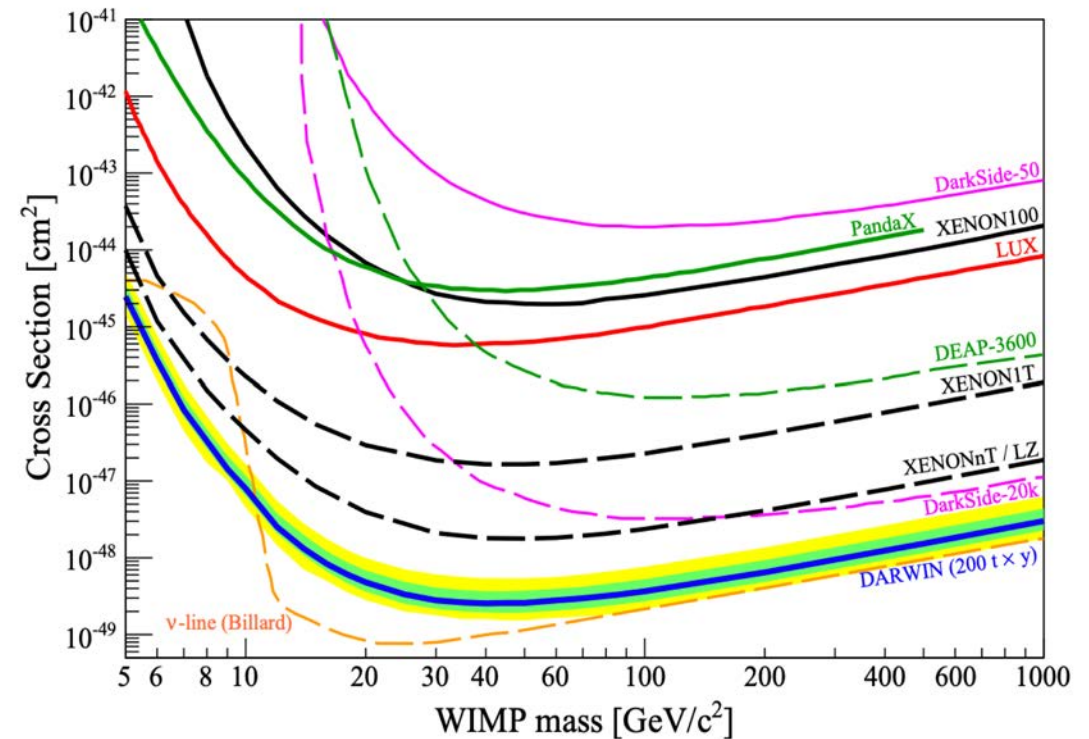
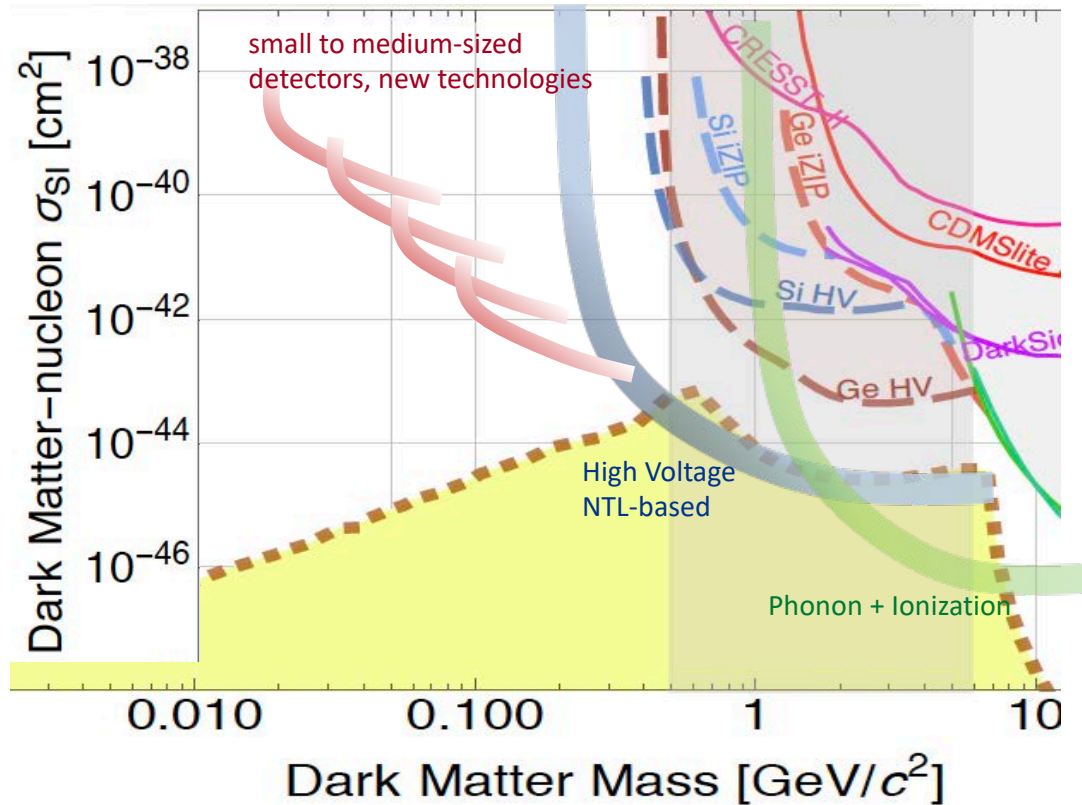
Joined CDMS and helped install in the Soudan Mine  
Spokesperson for SuperCDMS SNOLAB



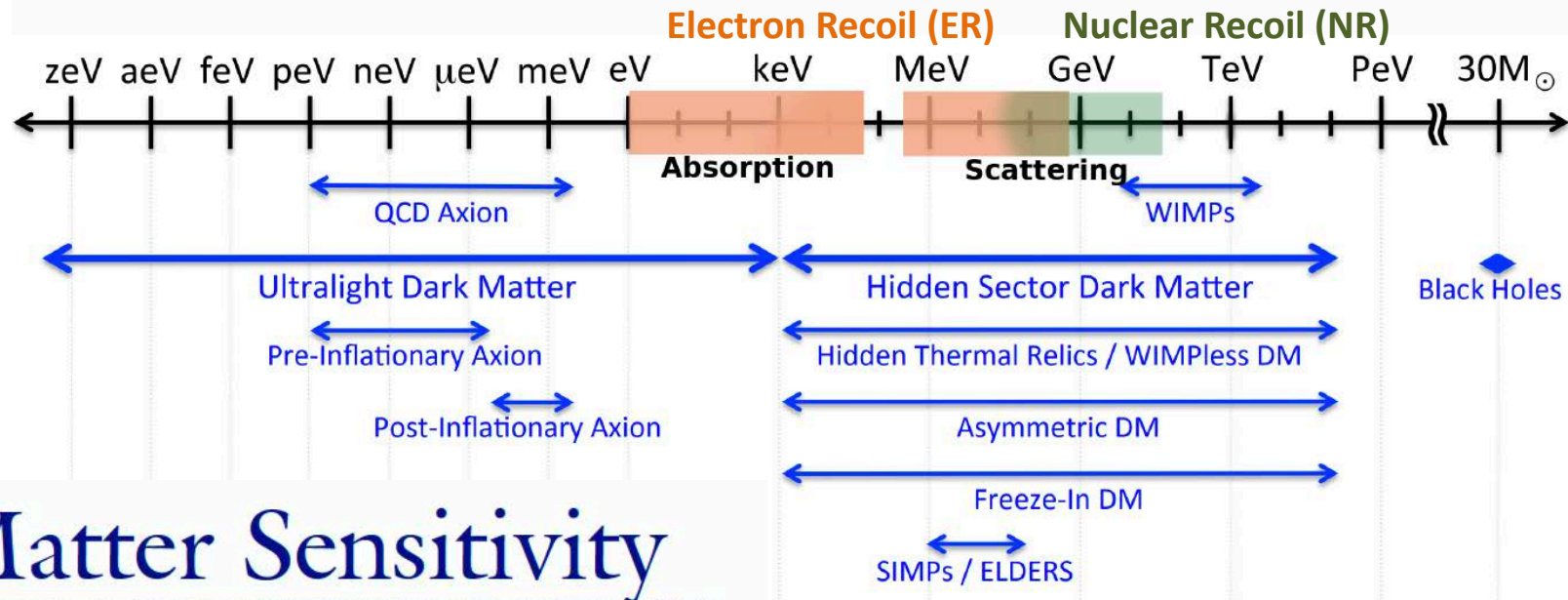
# Direct Detection of nuclear-recoiling Dark Matter currently covers two different mass regions with different technologies

Solid state detectors are struggling with exposure \$\$, but continue to push lower in mass with smaller detectors and new low-threshold devices

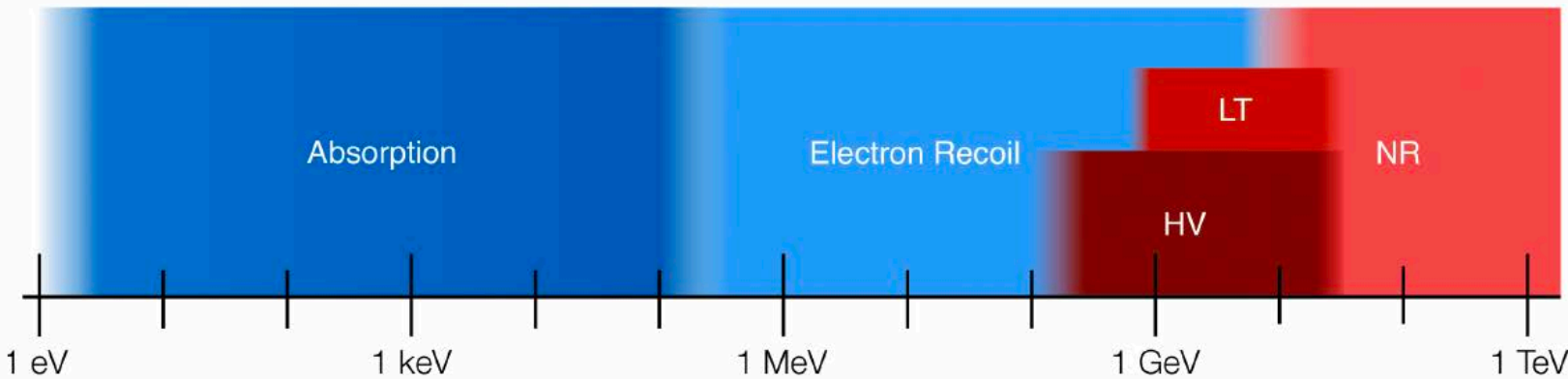
Nobel Liquid TPCs cover the WIMP landscape  
They are becoming neutrino detectors.



Cosmic Visions Workshop  
 Landscape reminded us that it  
 isn't all about nuclear recoils

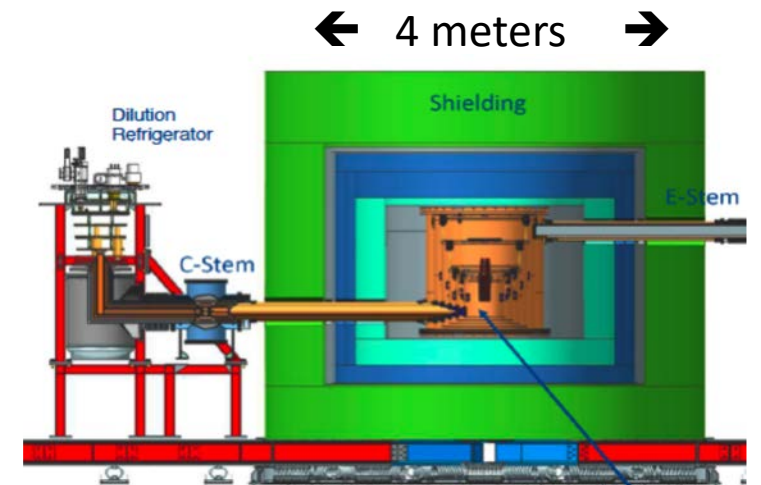


# SuperCDMS Dark Matter Sensitivity



**Mid-size facility needed in this range**  
 smaller than the noble liquid TPC facilities  
 e.g. SuperCDMS will operate 24 detectors in  
 mixed modality. Room for ~200 in the shield

<b>Traditional NR:</b>	iZIP, Background free	> 5 GeV
<b>Low Threshold NR:</b>	iZIP, limited discrimination	> 1 GeV
<b>HV Mode:</b>	HV, no discrimination	~0.3 - 10 GeV
<b>Electron Recoil:</b>	HV, no discrimination	~0.5 MeV - 10 GeV
<b>Absorption (Dark Photons, ALPs)</b>	HV, no discrimination	~1 eV - 500 keV (peak search)



## Other cryogenic experiments also targeting this region

New SuperCDMS R&D (HVeV and CPD)

EDELWEISS (bolometers, HV)

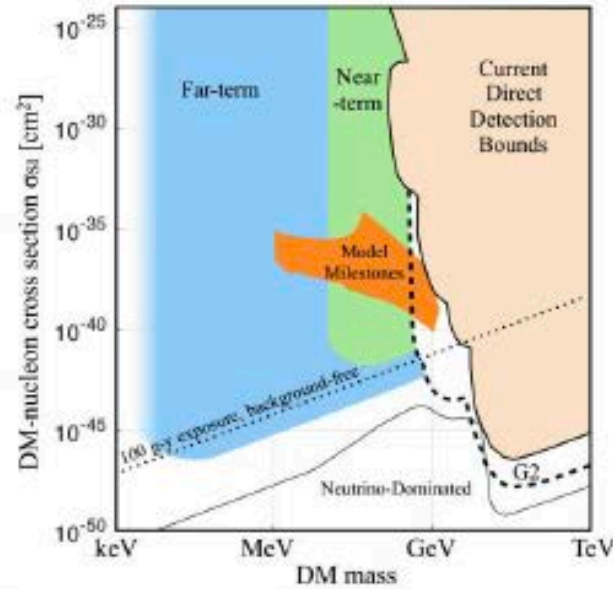
CRESST (CaWO<sub>4</sub> - also LiAlO<sub>2</sub>, Al<sub>2</sub>O<sub>3</sub>)

Sensei, DAMIC-M, Oscura (silicon CCDs)

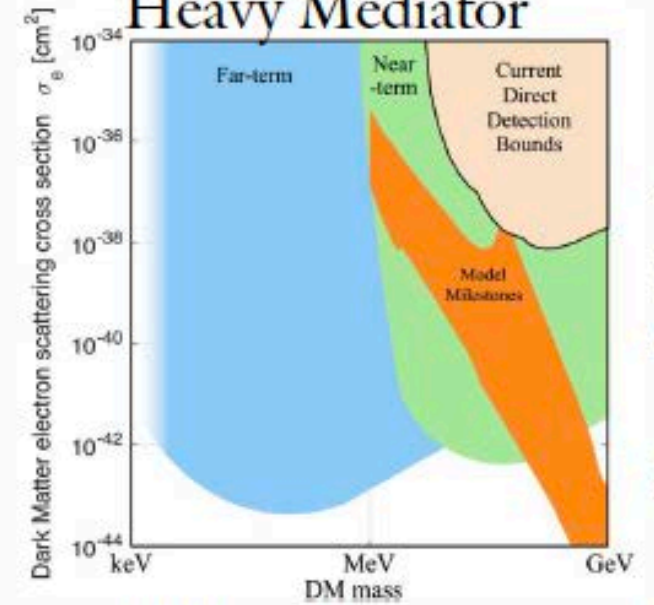
HeRALD (Superfluid He)

SPICE (polar crystals)

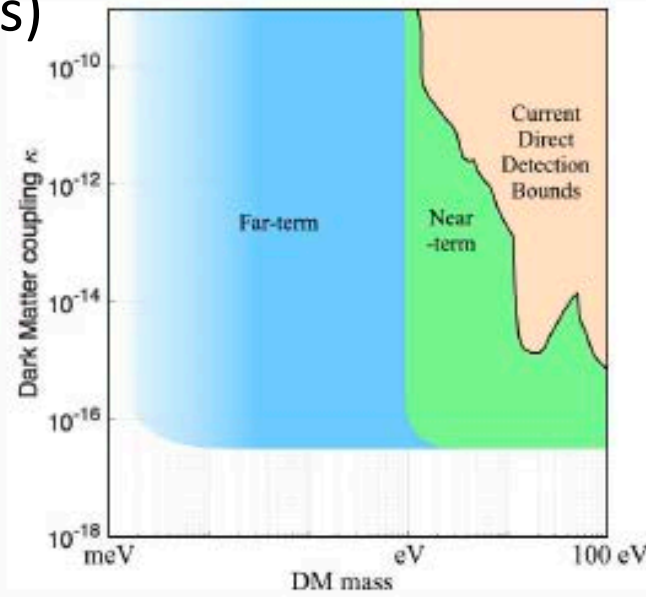
## NR Scattering



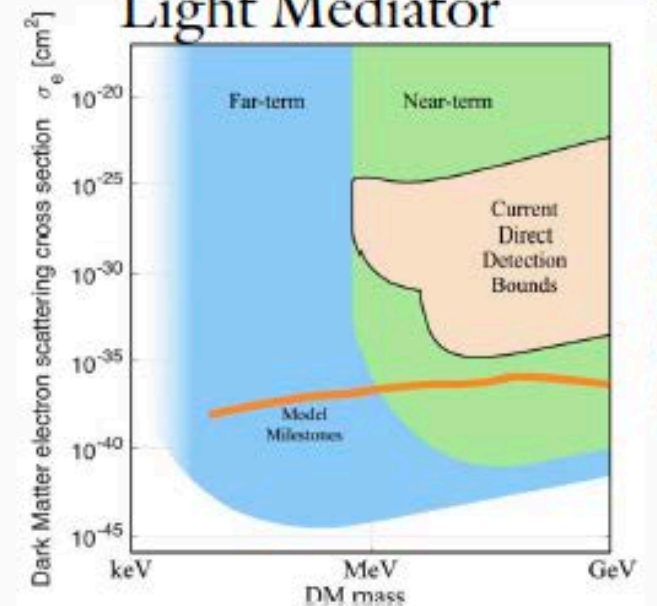
## ER Scattering Heavy Mediator



## Dark Photon



## ER Scattering Light Mediator



## The problem is predicting the next best Dark Matter detector to put in SURF

The problem with dark matter searches is our ignorance! We don't know its mass or how it interacts  
Underground needs differ by DM technology and candidate.

It is already true that new technologies are no longer about exposure, but about new parameter space  
Many can be done above ground.

→ Accelerator and beam dump searches

→ Axions and wave-like DM

→ Even Low mass DM with solid state and/or superfluid He: R&D starts at the surface.

Modest space requirements underground

Future Noble Liquid TPCs need BIG cavern installations, many are already spoken for. After LZ, what?

Paradigm shift coming up. If no WIMP-like DM found down to the neutrino floor, attention will shift to  
Annual modulation (crystals) - Multiple northern and southern locations

Moderate Depth, Excellent purification techniques

Directional detection (wire chambers/TPCs etc.) large installations still needed for full exposure

AND different candidates → less need for large underground space

# SURF should brand itself as the premiere neutrino laboratory

Build on your strengths to create an “ecosystem” of international underground labs with complementary expertise

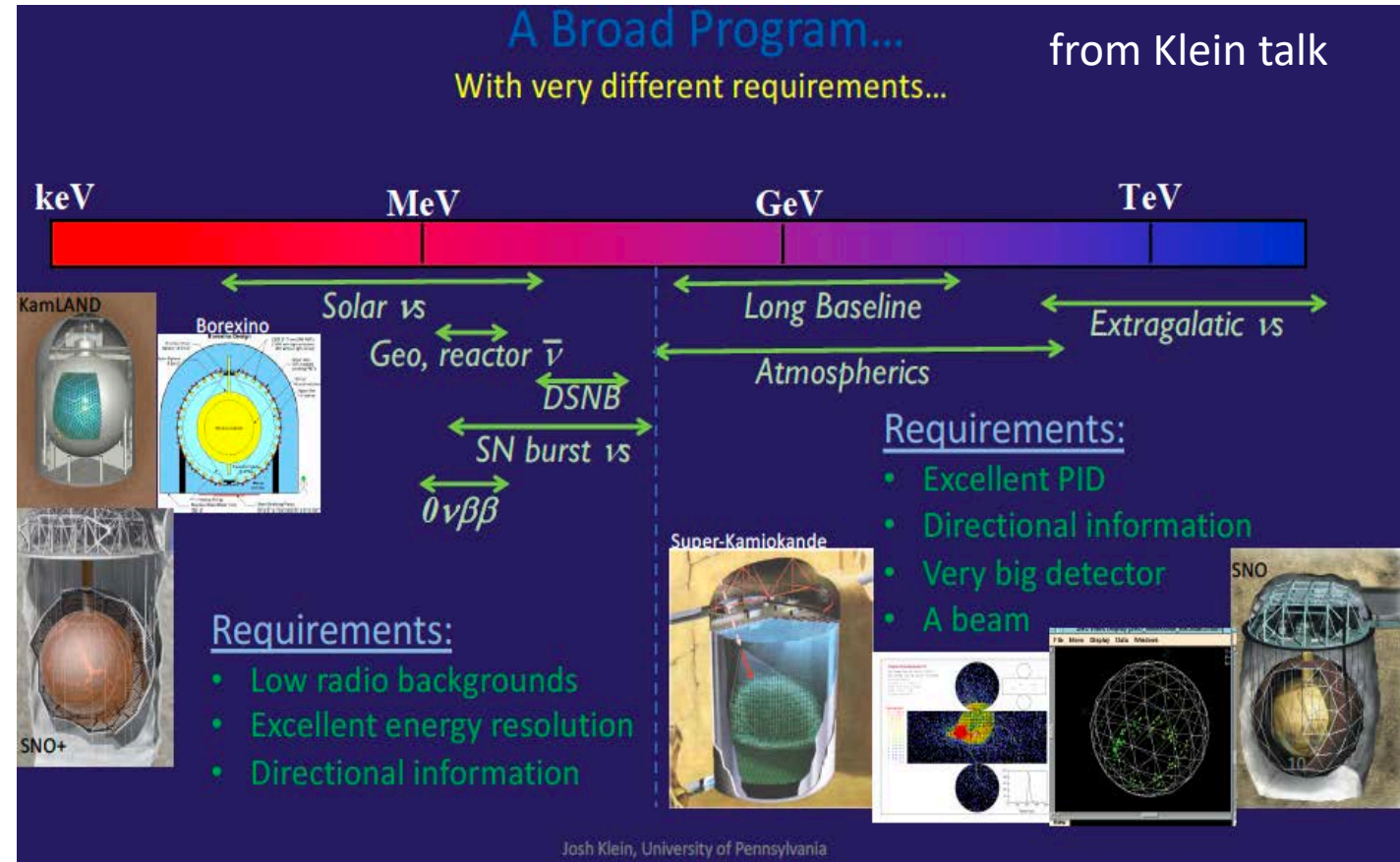
## Strengths

Large caverns (engineering expertise)  
Multiple campuses and room for expansion  
DUSEL legacy and Low bkgd counting infrastructure  
Surface sites at local institutions

A neutrino beam!!  
A flagship neutrino experiment (DUNE)  
Space for more exciting BIG neutrino installations  
Neutrinoless double beta decay (Majorana)  
CASPAR - upgrades? new location?

Sell new experiments on the neutrino physics and related infrastructure.

Maybe the next noble liquid TPC is a neutrino experiment and dark matter is a background (*just kidding....?*)



# Another strength to capitalize on is Low Background Infrastructure and assay techniques

Sheer amount of screening needed → MORE dedicated space for ultra-sensitive screening

Make connections to other screening facilities and unify the process

This does not easily happen from the top down – Progress tends to be driven by the experiments


e.g. LZ and the UK collaborators → Boulby

But that leaves out smaller players and non-physics applications needed to build up a multi-disciplinary bio/geo presence.

Large space needed for

- Crystal growth and fabrication
- Copper Electroforming
- Stockpile of cosmogenically-sensitive materials
- Water shielded ultra-sensitive screening space

The ultimate low background facility was explored in the context of DUSEL and all the engineering still exists.



Boulby Underground Laboratory

Scovell's talk

## Challenges for the future

- Next generation particle physics will put huge pressure on facilities
  - Potentially 10s of 1000s of measurements needed
- At LRT2019, I calculated at least 77 germanium crystals dedicated to material assay of varying sensitivity
  - Challenge here becomes one of worldwide coordination, cross-calibration & QA/QC
- For **radon emanation**, mass-spec, NAA and **surface alpha** there may still be too few facilities to meet future demand
  - Some labs starting to ramp up but there is certainly room for more!
  - Most relatively low cost but high impact
- And don't forget cleanliness!

*<math>\mu\text{Bq/kg}</math> required for next-gen!\**      *Numbers growing!*

# A DUSEL vision of a world-class Low Background Facility

Entire facility is class 10,000 clean room,  $< 20 \text{ Bq/m}^3$

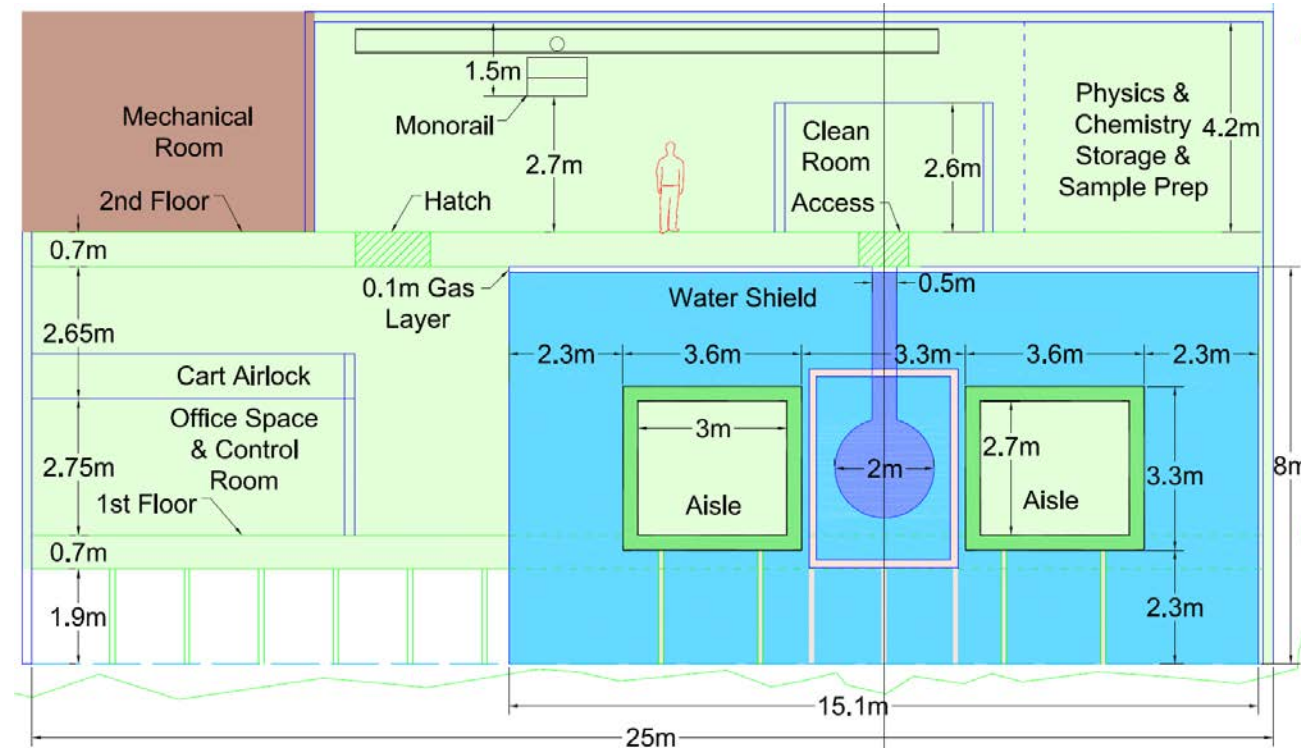
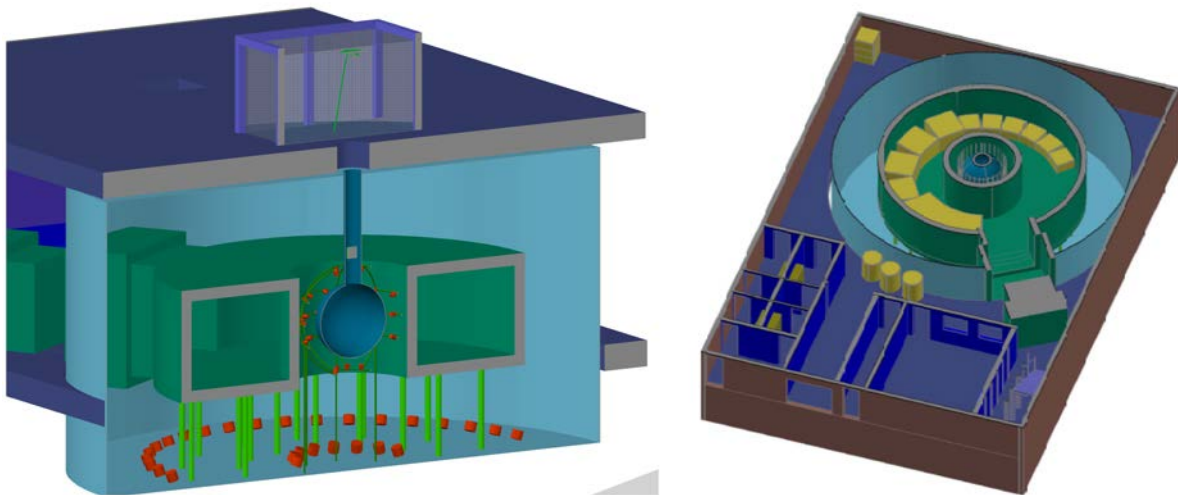
- Radon-mitigated zones ( $< 1 \text{ Bq/m}^3$ ) and assembly areas ( $< 0.1 \text{ Bq/m}^3$ )
- Radon-free storage and unified LN system
- Wet benches, clean machining, hoods, Several class 1000 clean rooms

Instrumented Water Shield with toroidal interior acrylic room

- Much cheaper than individual lead castles. Instrumented water shield provides veto capacity, bkgd monitoring.
- Houses ultra-sensitive screeners, staging space for experiments & R&D Prototypes
- Serves as the outer shield of the Immersion Tank

Central Top-loading Immersion Tank

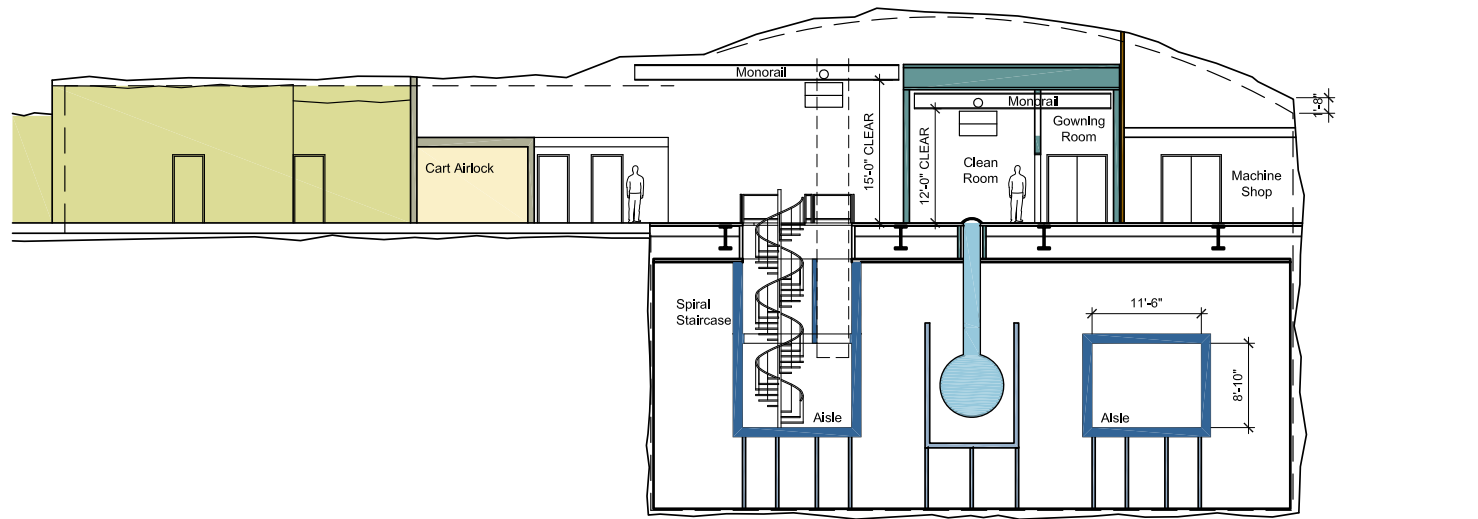
- modeled on the Borexino CTF – whole body counting





# A DUSEL vision of a world-class Low Background Facility

Original Design fit into “Lab Module 2” in DUSEL plans at Homestake.  
A fully engineered, smaller, non-site specific design also exists



1 FAARM - LONGITUDINAL SECTION  
A3 3/32" = 1'-0"

- Unclassified Space
- Transition Space
- CMU Walls
- Clean Room
- Modular Walls
- Torus Walls

## Inner Tunnel Lab

$\gamma$ -flux  $7.974 \times 10^{-5} \text{ cm}^{-2} \text{ s}^{-1}$   
n-flux  $4.817 \times 10^{-10} \text{ cm}^{-2} \text{ s}^{-1}$

4 < ppt (GeMPI, arrays)  
6 < ppb (well, clover, coax)  
2 Beta Cages  
Prototyping Space  
(DM or  $0\nu\beta\beta$  or novel assay)

Radon Mitigation  
Common cryogen plumbing and  
LN boil-off for screeners

## Central Pool

0.1 counts/day,  $E > 250 \text{ keV}$   
sensitivity of  $10^{-14} \text{ g/g U/Th}$   $10^{-12} \text{ g/g K}$   
modeled on Borexino CTF  
2m diam nylon vessel filled with LS  
Observed by low rad QUPIDs  
Top-loading from dedicated Clean Room

# Reconsider many DUSEL goals in the next decade.

For example, no one has successfully created a true multi-disciplinary underground lab. Smaller projects are not self-supporting; they need to be subsidized until they mature. The DUNE project can provide this.

They also need a unifying theme and an Institute: **Climate Change is this century's challenge**

This theme already shares many of the established geo/bio/engineering science goals.

An Institute at SURF would also promote work at other sites, unified by the theme and led by geo/bio/engineering

## Examples

- Energy production
  - geothermal energy
  - understanding fracking
- Changes to the water cycle
- Climate and the critical zone
- CO<sub>2</sub> sequestration
- Vulcanism, subduction, CO<sub>2</sub> cycles
- Biogeochemical cycles,
- New microbes for digestion/waste

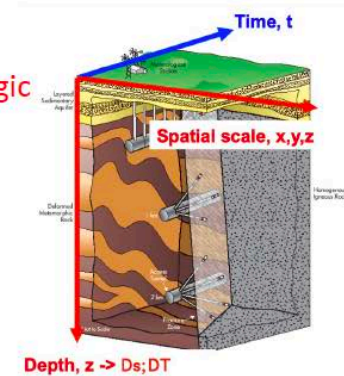
## Attributes of SURF for BG&E Grand Challenges

### Pro

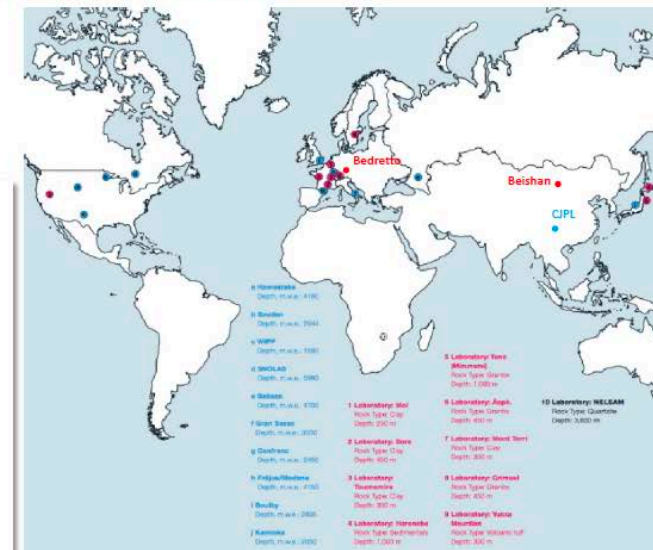
- Large Spatial Scale – Access to heterogeneous/opaque block at km-scale
- Large Depth – Elevated stress and temperature
- Long Term Occupancy – Continuity
- Low Background Noise – Seismically
- Proximal Access – To processes/expts. At depth
- Active Experimentation – Ameliorates constraints of the (very specific) geologic environment

### Con

- Merely One Environment – Rock type/non-sedimentary
- Many Competing Locations – Some with more-favorable/specific attributes



Elsworth Slide





DPF Community Planning Exercise

## UNDERGROUND FACILITIES

### Frontier Conveners

Name	Institution	email
Laura Baudis	University of Zurich	<a href="mailto:laura.baudis@physik.uzh.ch">laura.baudis[at]physik.uzh.ch</a>
Jeter Hall	SNOLAB	<a href="mailto:Jeter.Hall@snolab.ca">Jeter.Hall[at]snolab.ca</a>
Kevin Lesko	Lawrence Berkeley National Laboratory	<a href="mailto:ktlesko@lbl.gov">ktlesko[at]lbl.gov</a>
John Orrell	Pacific Northwest National Laboratory	<a href="mailto:john.orrell@pnnl.gov">john.orrell[at]pnnl.gov</a>

**Advertisement:** A robust Snowmass White Paper on a unified plan for complementary strengths of Underground Labs worldwide would help SURF.

Explore the ecosystem! Define it.